

# **Broadband imaging of anisotropy and composition in the Earth's mantle**

Ludwig Auer

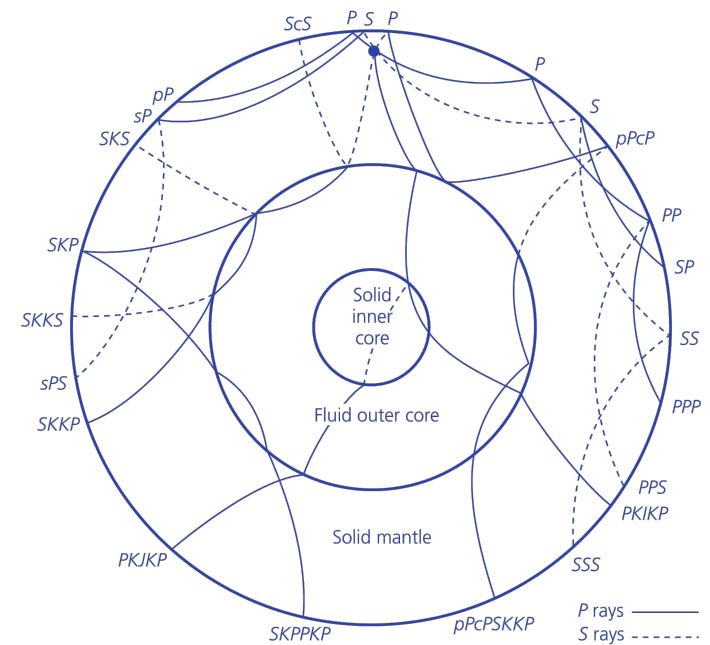
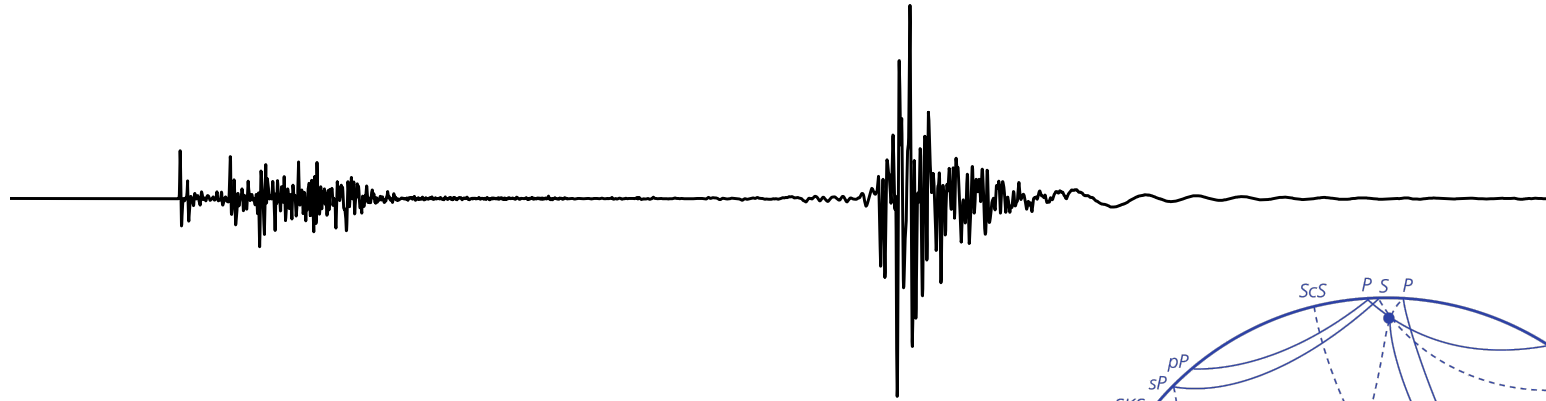
Institute of Geophysics – ETH Zürich

Committee: Domenico Giardini, Philippe Lognonné ,  
Lapo Boschi, Tarje Nissen-Meyer

March 14<sup>th</sup> 2016

# (1) Introduction: Tomography approach

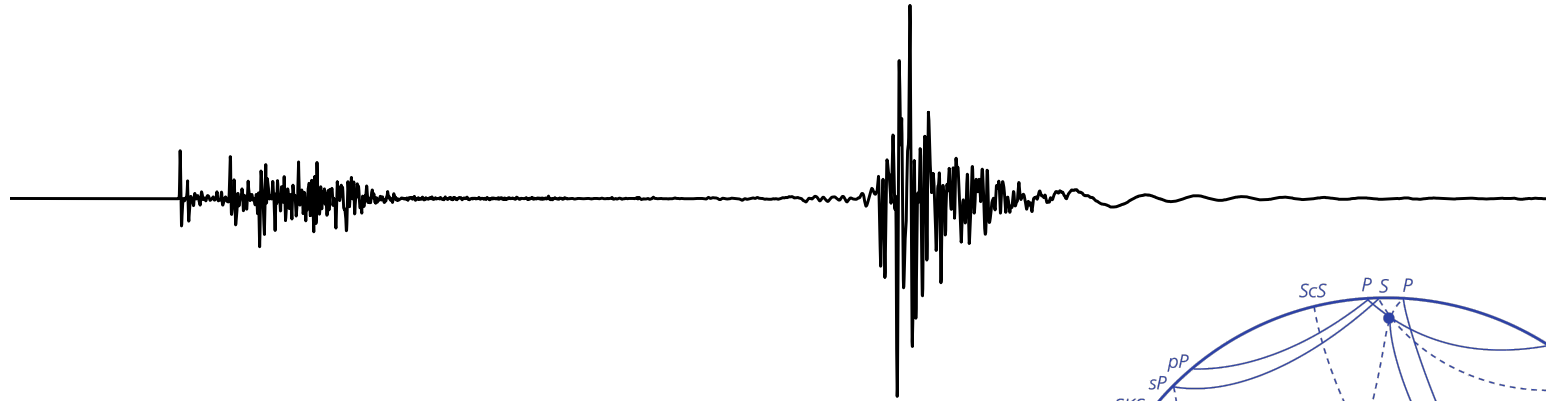
Why broadband seismic information?



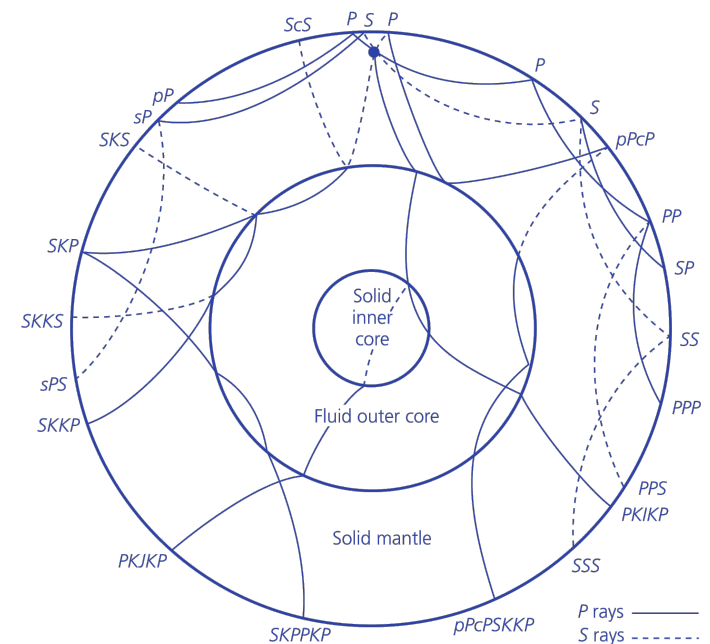


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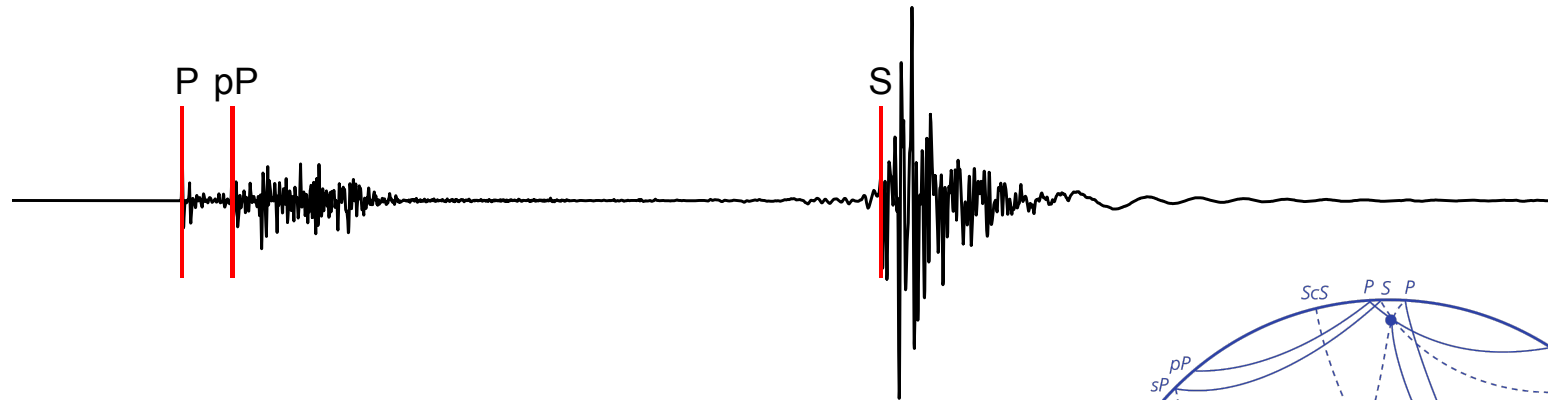


- Different time-frequency portions of the seismic record sense different regions

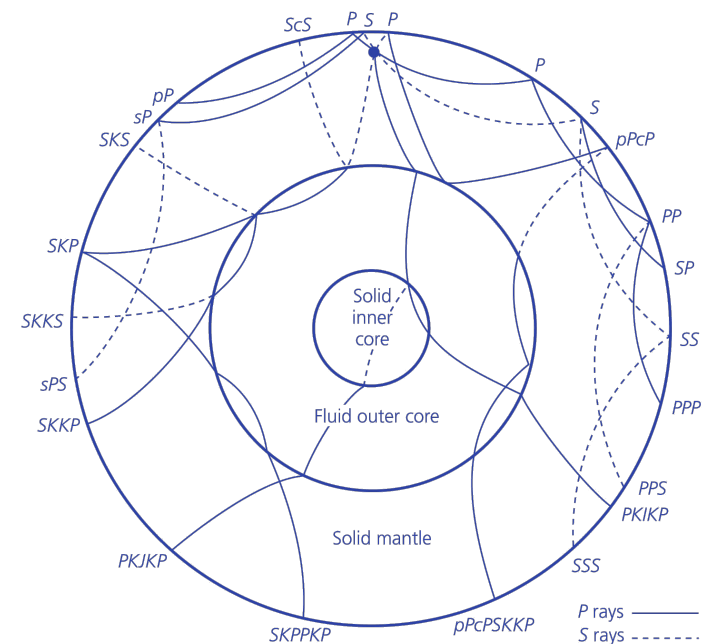


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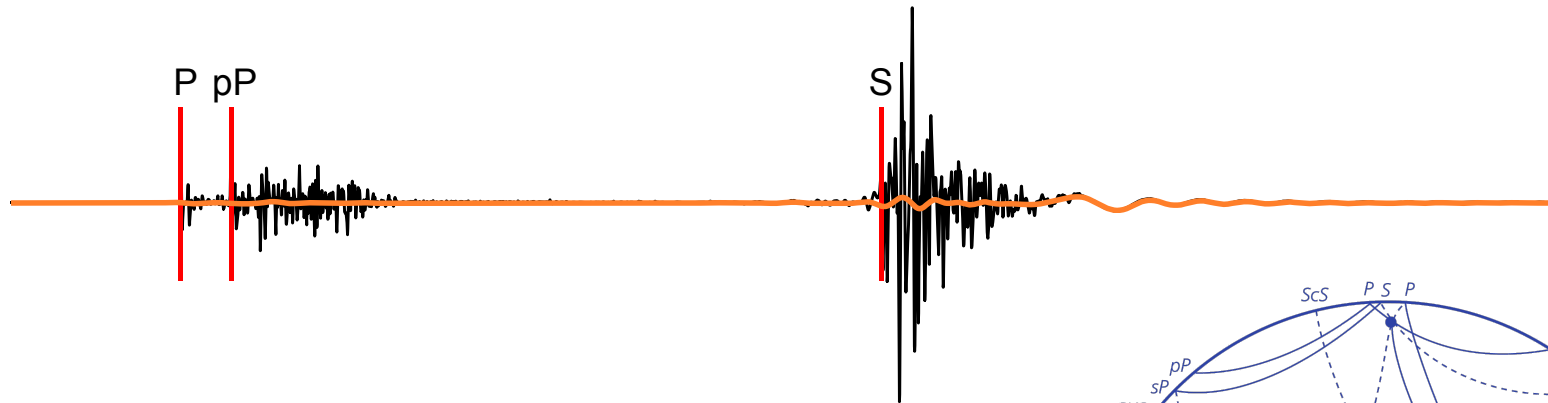


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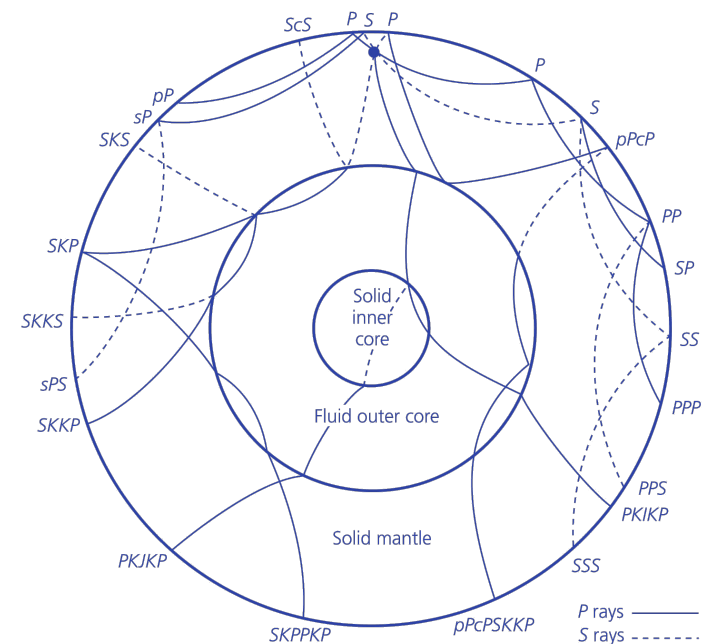


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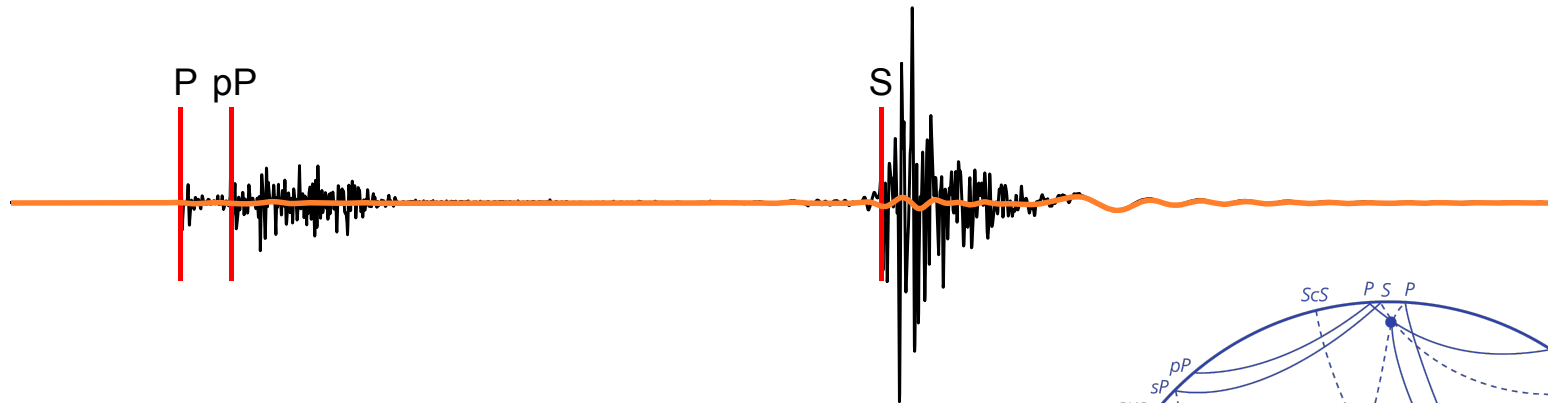


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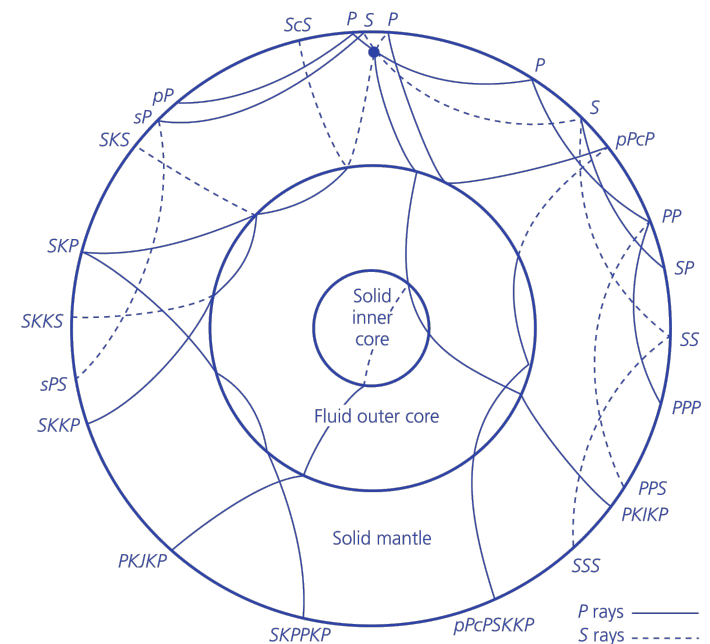


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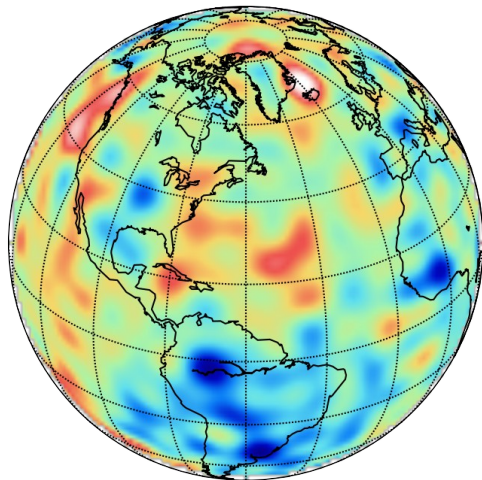
- Different time-frequency portions of the seismic record sense different regions
- Developing meaningful whole-mantle demands using **high-frequency** and **lower-frequency** information together



# (1) Introduction: Tomography approach

Imaging multi-scale structure and dynamics

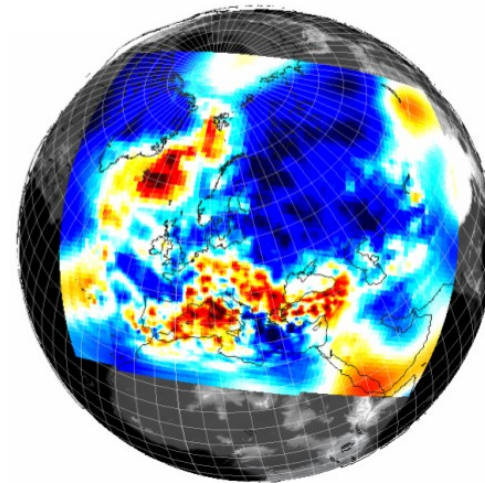
Global scale



Approximate theory and  
1D starting models

Ray theoretical global model  
S40RTS, (Ritsema 2011)

Regional scale



Fully numerical and 3D  
background models

Adjoint tomography of Europe and  
the Middle East (Fichtner 2012)

# (1) Introduction: Tomography approach

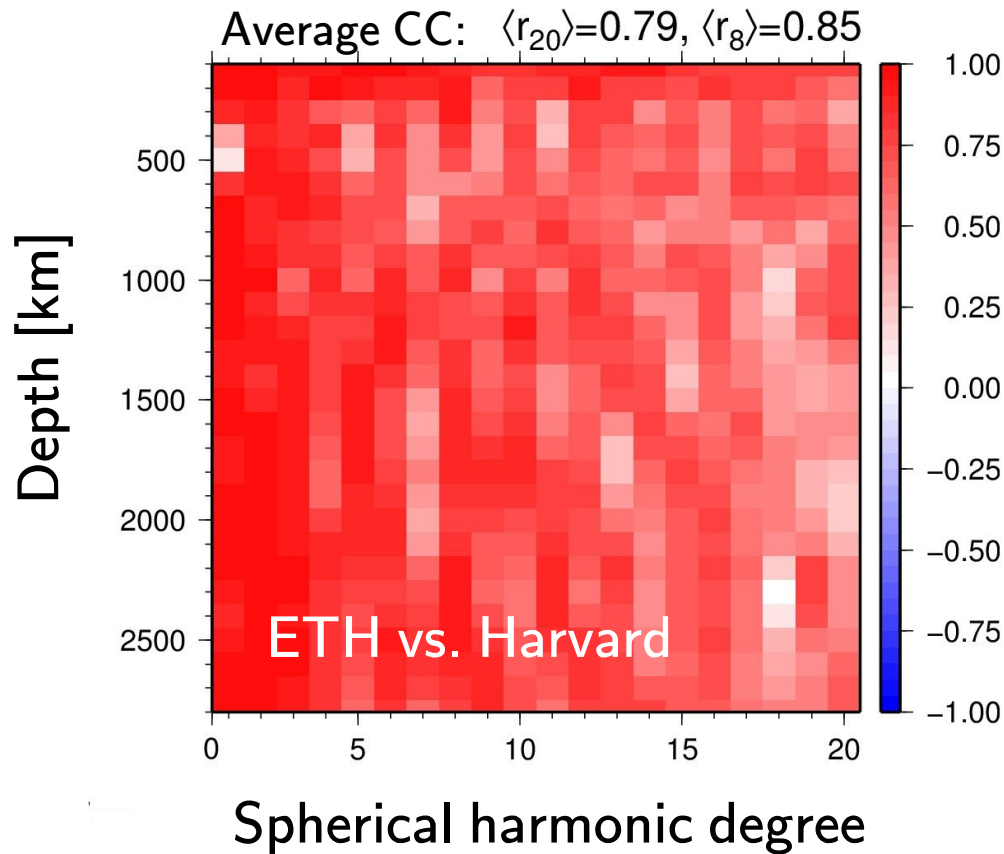
Imaging multi-scale structure and dynamics

Compromise “full”-waveform tomography and ray theory, jointly model global and regional structure?

Multi resolution models?

# (1) Introduction: Earth's 3D structure

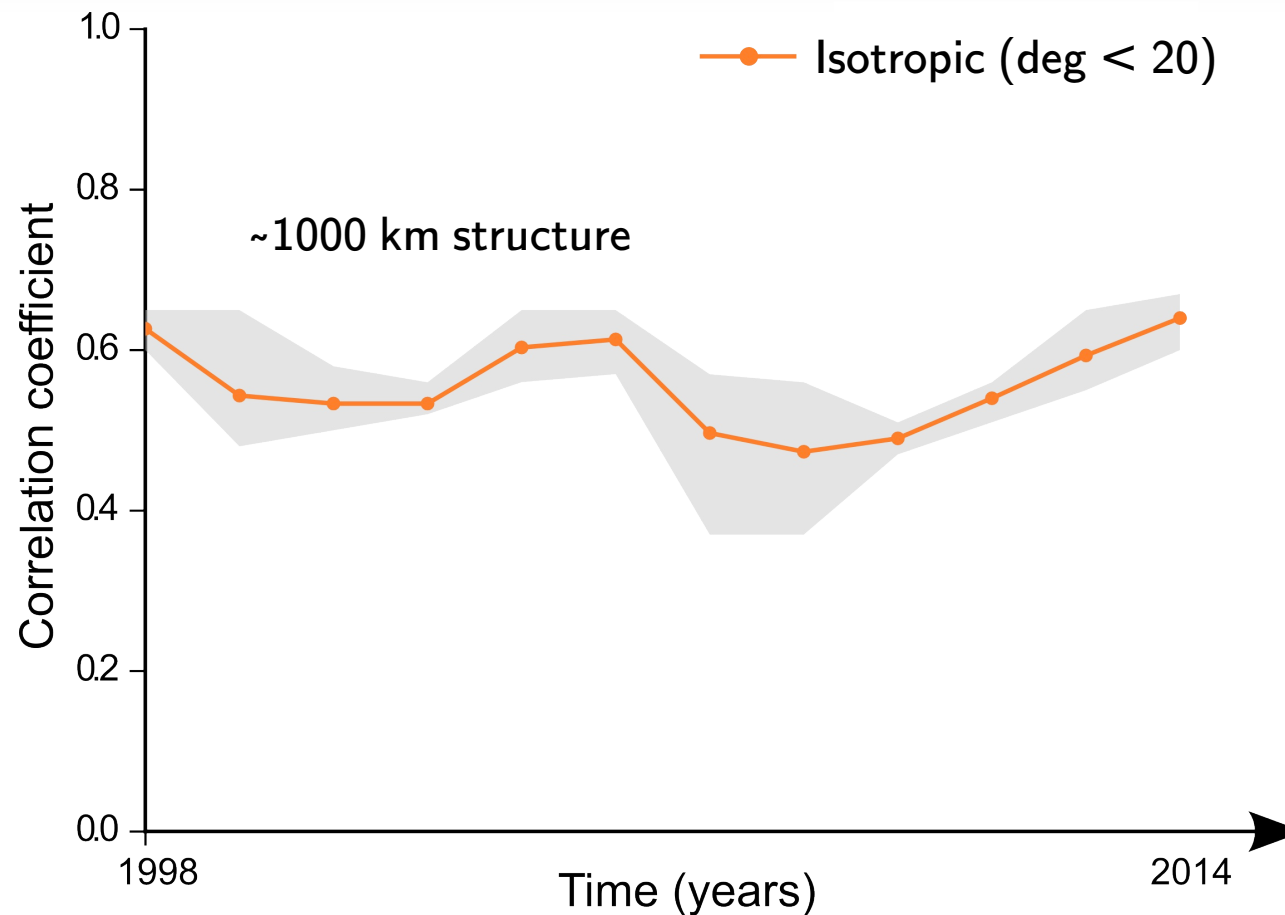
## SH model comparison (isotropic $v_s$ )



- Isotropic tomographic models correlate well for the SH harmonic degrees
- Model consistency validates modeling algorithms

# (1) Introduction: Earth's 3D structure

## Model congruency over the years

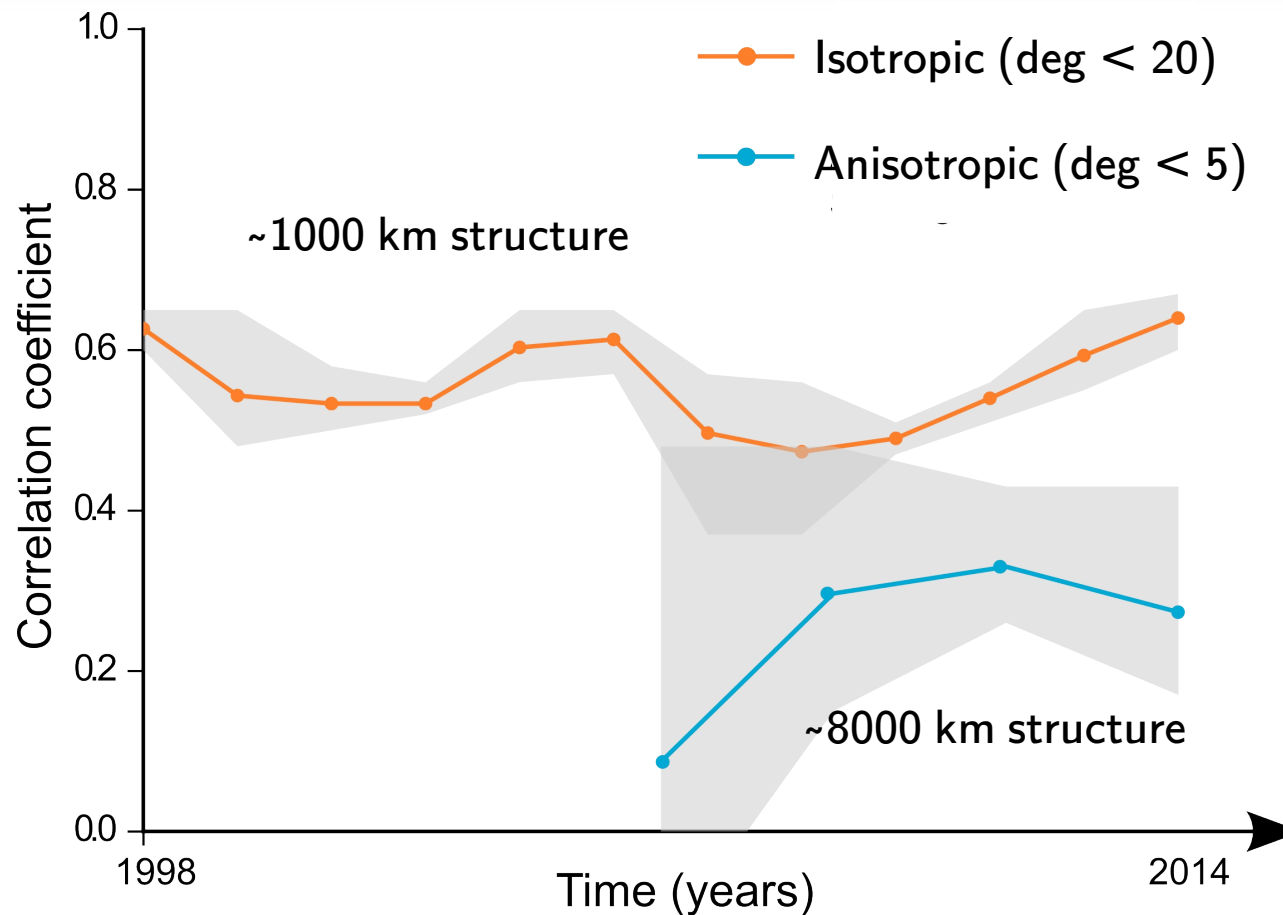


- Models still lack consistency at short spatial wavelength
- For Geodynamically relevant parameterizations situation is even worse
- E.g. anisotropy ...



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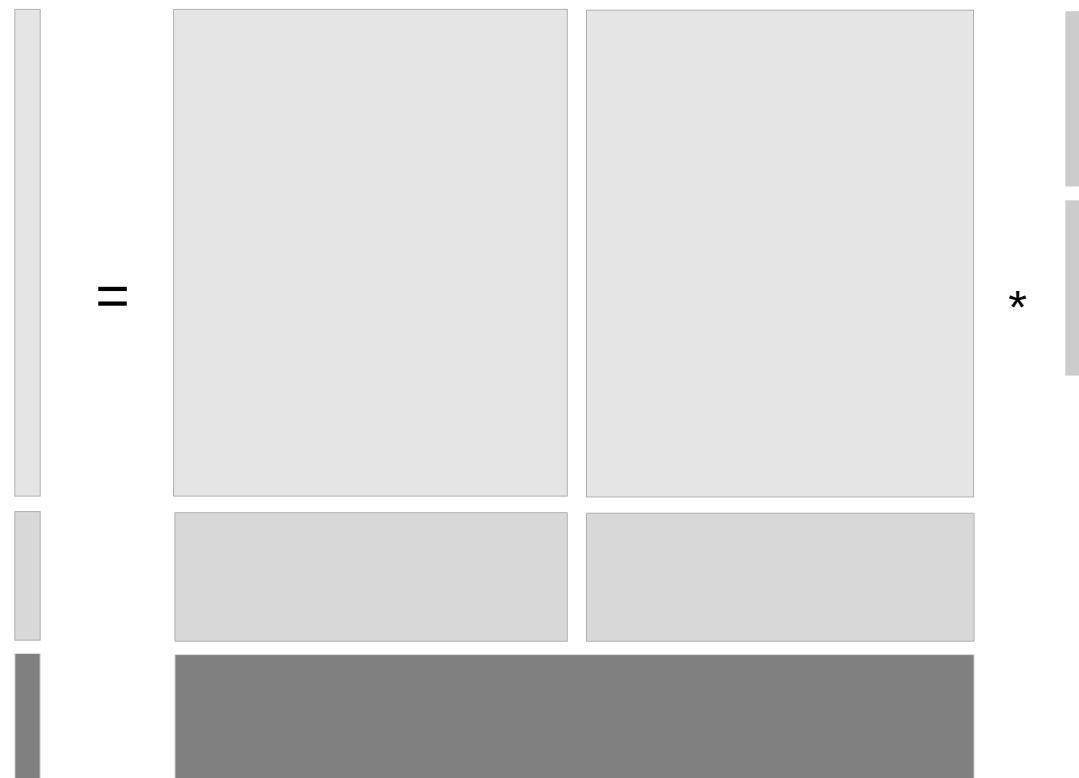
# (1) Introduction: Thesis outline

Chap.

2. Whole-mantle anisotropy from surface and body waves at adaptive resolution (*Auer et al. 2014*)
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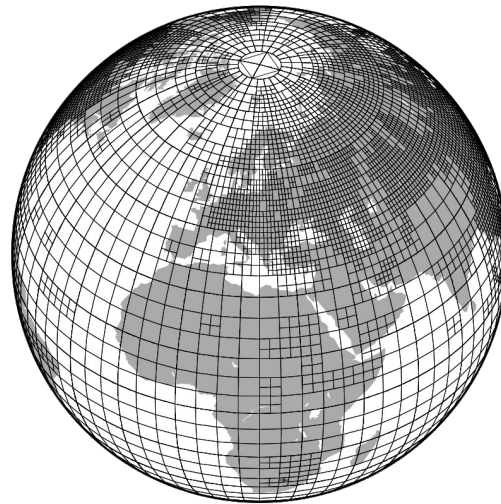
## (2) Whole-mantle anisotropy

Summary of linear system  $d=Gm$



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Summary of linear system  $d=Gm$

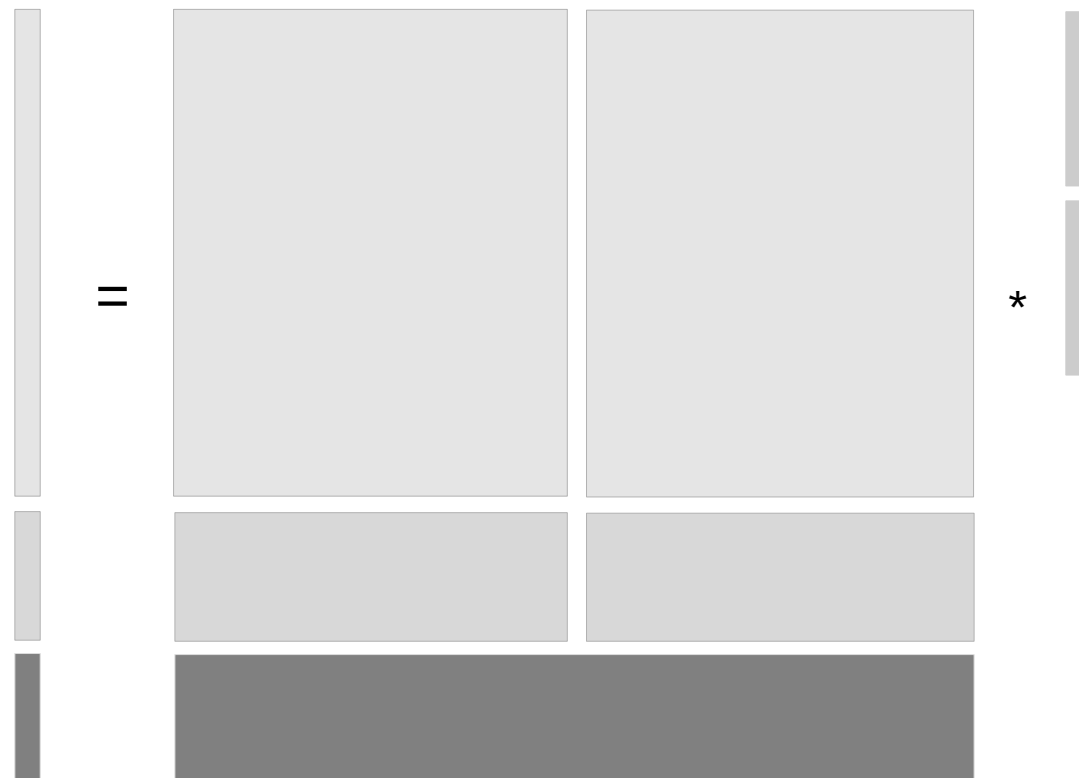


Voxel discretization

- Multi-scale
- Orthogonal
- Simple

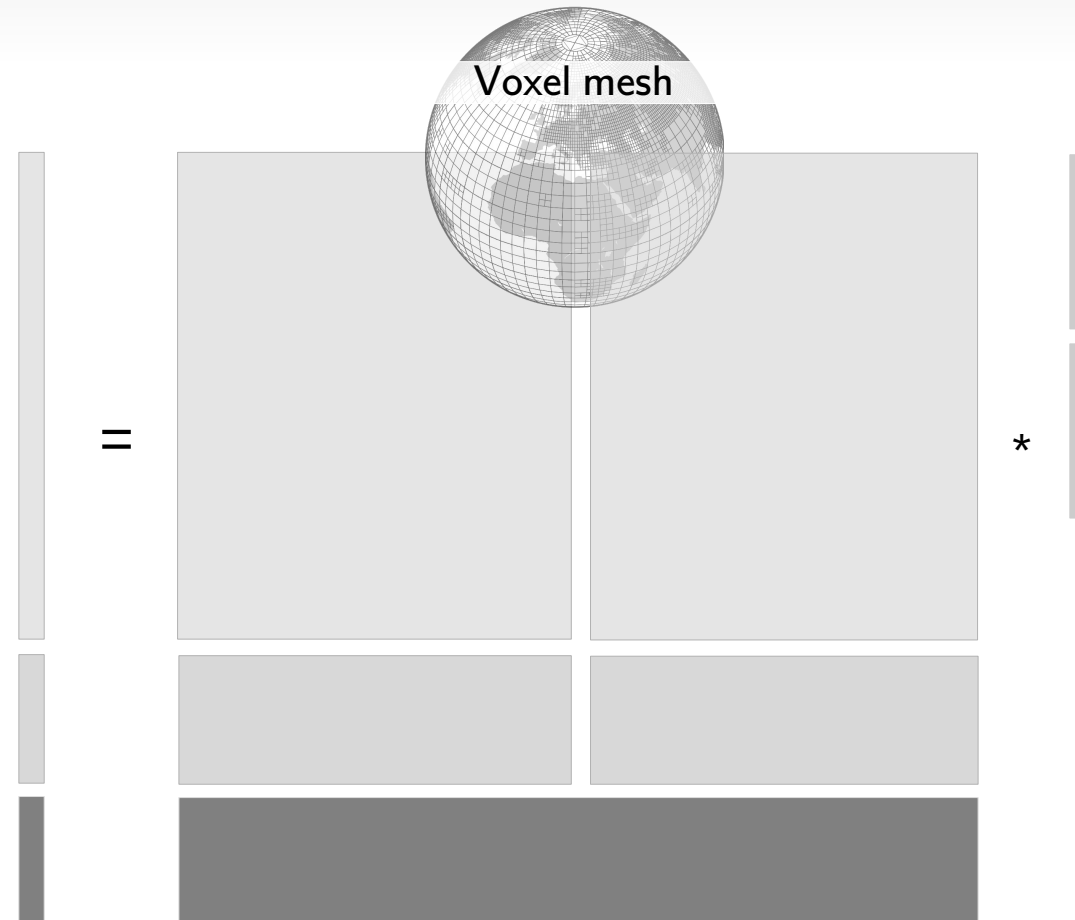
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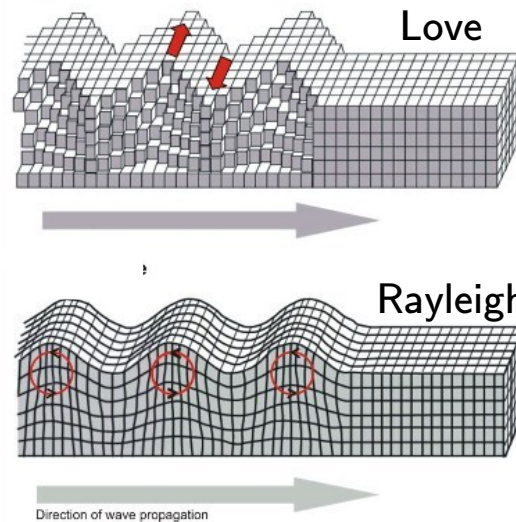


# (2) Whole-mantle anisotropy

## Summary of linear system $d=Gm$



### Surface wave dispersion



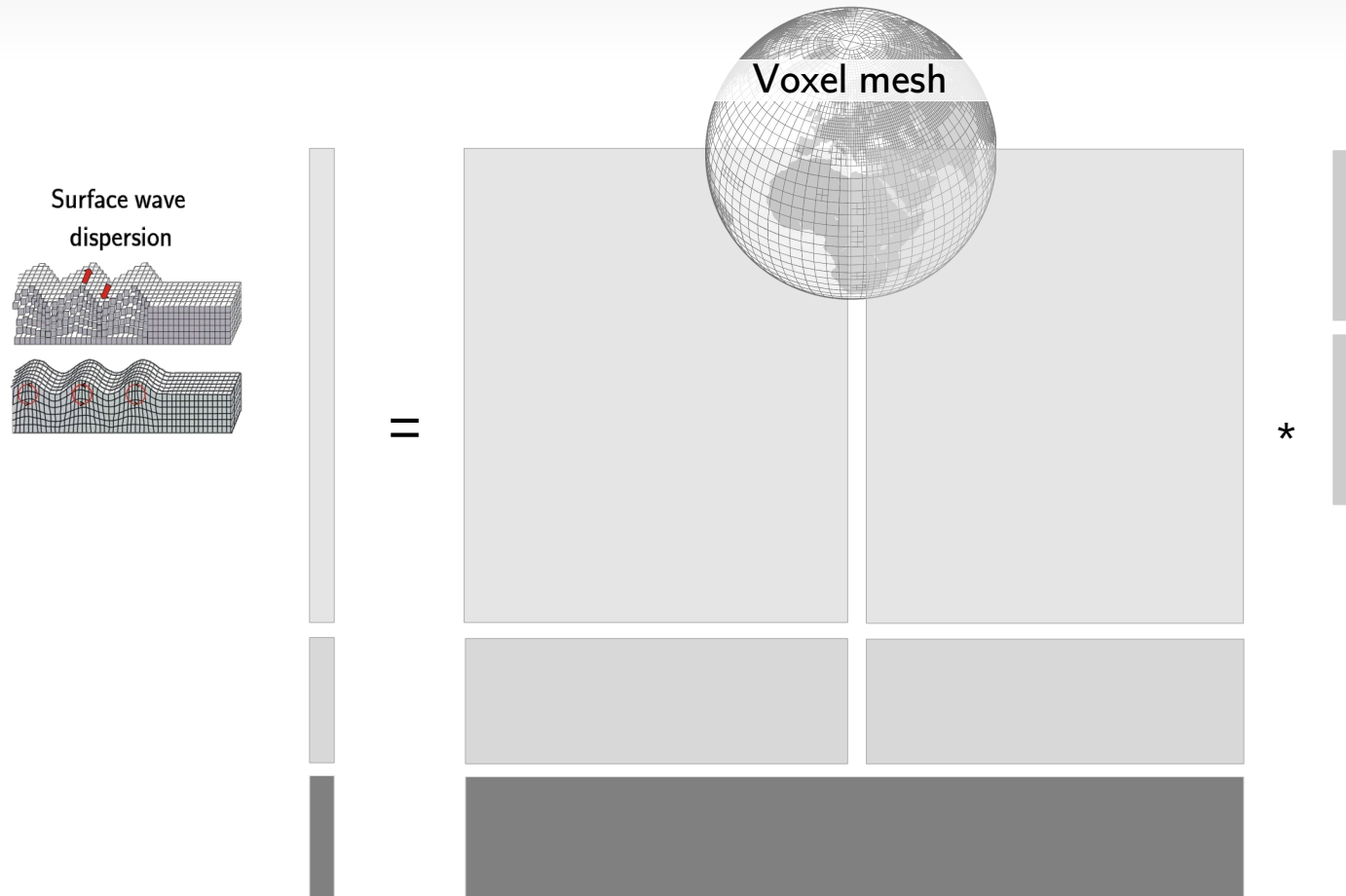
### Published fundamental modes and overtones

Overview of surface and body wave datasets				
Author	Spheroidal	s. rays	Toroidal	s. rays
Ritsema et al. (2004)	FM 37s-374s	2977971	FM 37s-375s	1605517
	1st OT 37s-274s	391495	1st OT 37s-269s	666057
	2nd OT 37s-365s	343398	2nd OT 37s-151s	242996
	3rd OT 37s-273s	316161	3rd OT 37s-100s	118018
	4th OT 37s-149s	300515	4th OT 37s-69s	54279
	5th OT 37s-114s	240256	5th OT 37s-56s	31106
	6th OT 37s-78s	162507		
Visser et al. (2008)	FM 35s-175s	477776	FM 35s-173s	399792
	1st OT 35s-172s	396431	1st OT 35s-176s	331168
	2nd OT 35s-149s	364140	2nd OT 35s-115s	250315
	3rd OT 35s-87s	253143	3rd OT 35s-78s	154160
	4th OT 35s-61s	159448	4th OT 35s-62s	81592
	5th OT 35s-56s	114037	5th OT 35s-56s	42756
	6th OT 37s-50s	71652		
Ekström (2011)	FM 25s-250s	1022706	FM 25-250s	342261

Table 1: Overview of surface wave datasets

## (2) Whole-mantle anisotropy

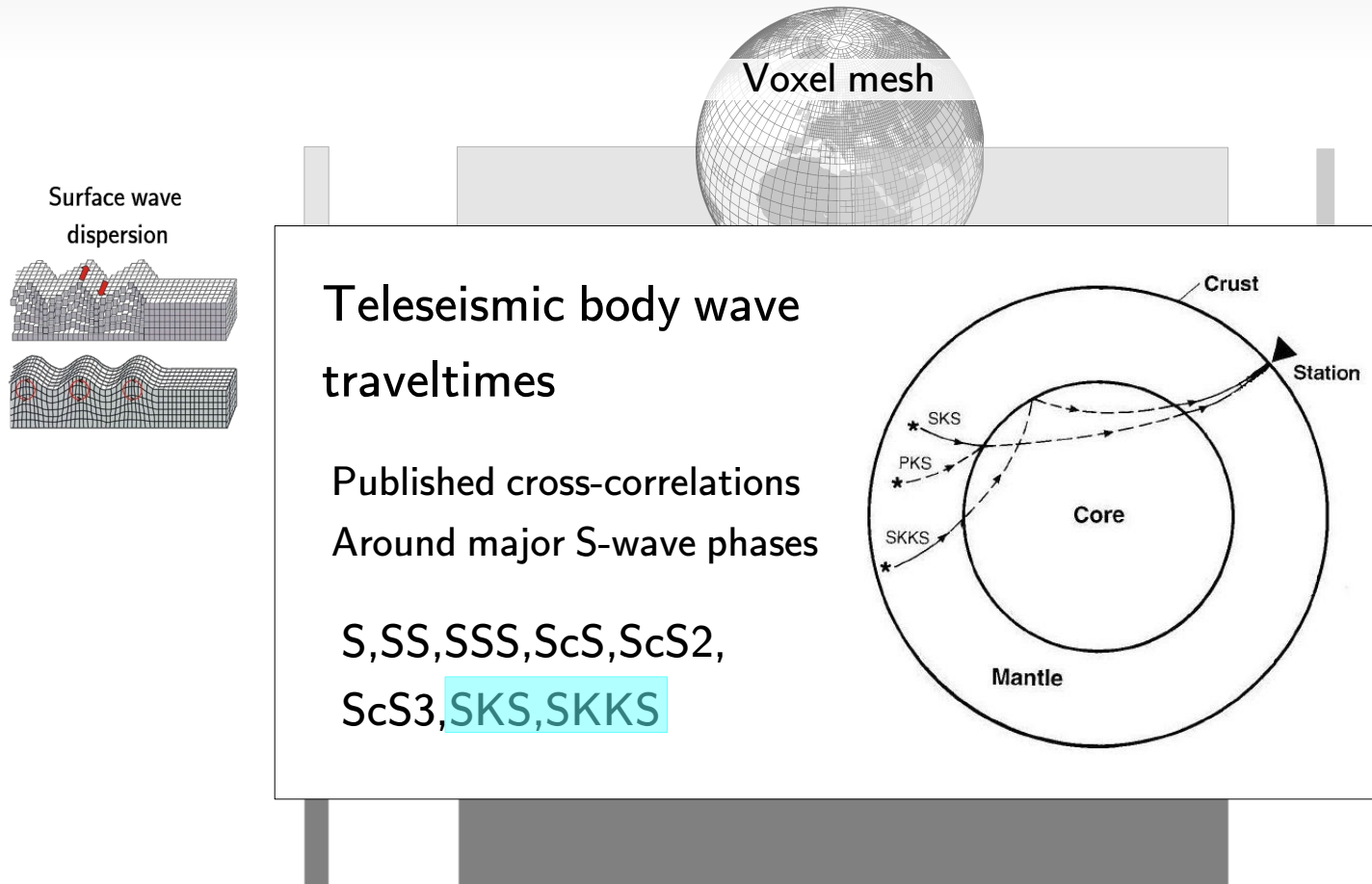
Summary of linear system  $d=Gm$





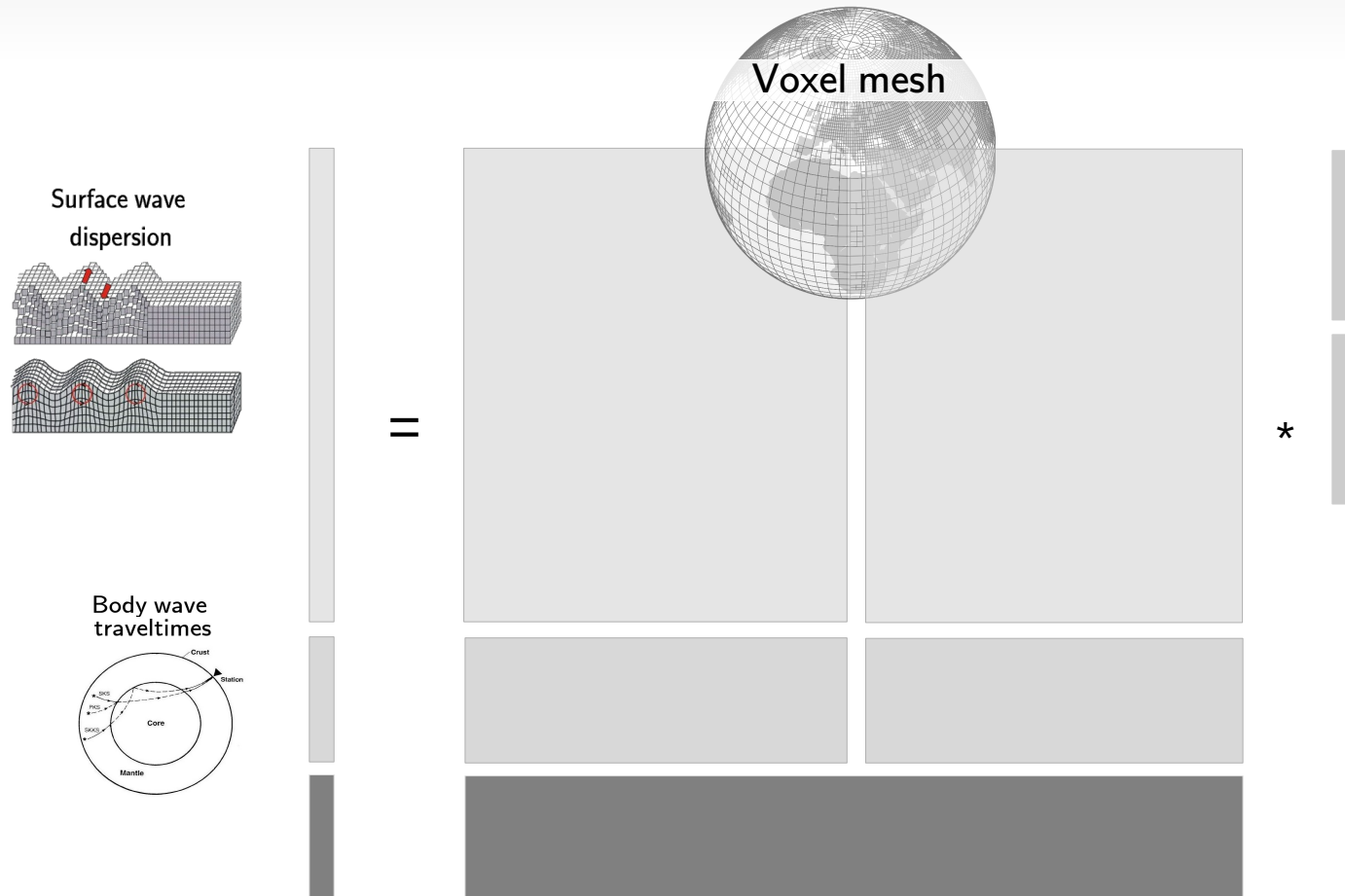
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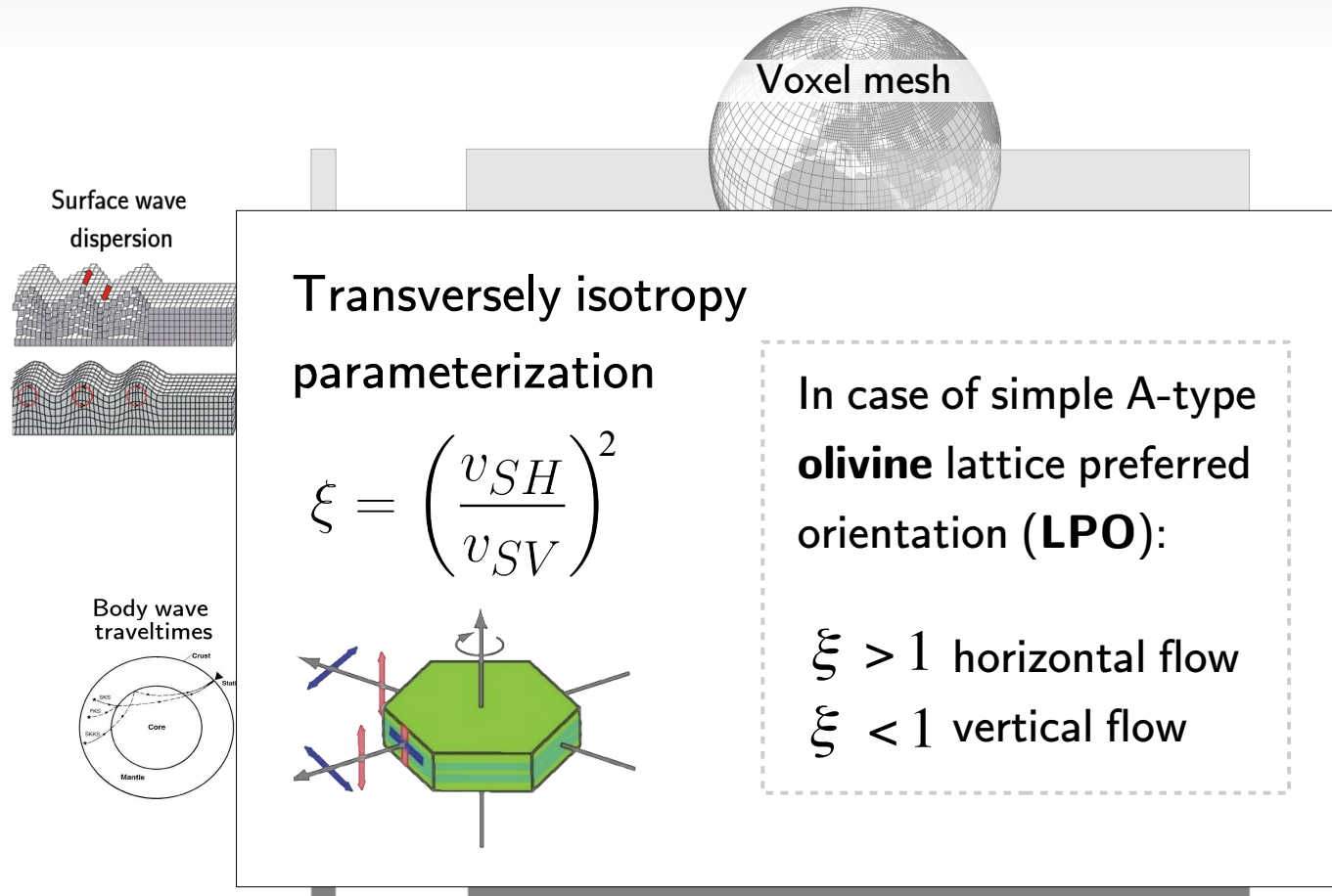
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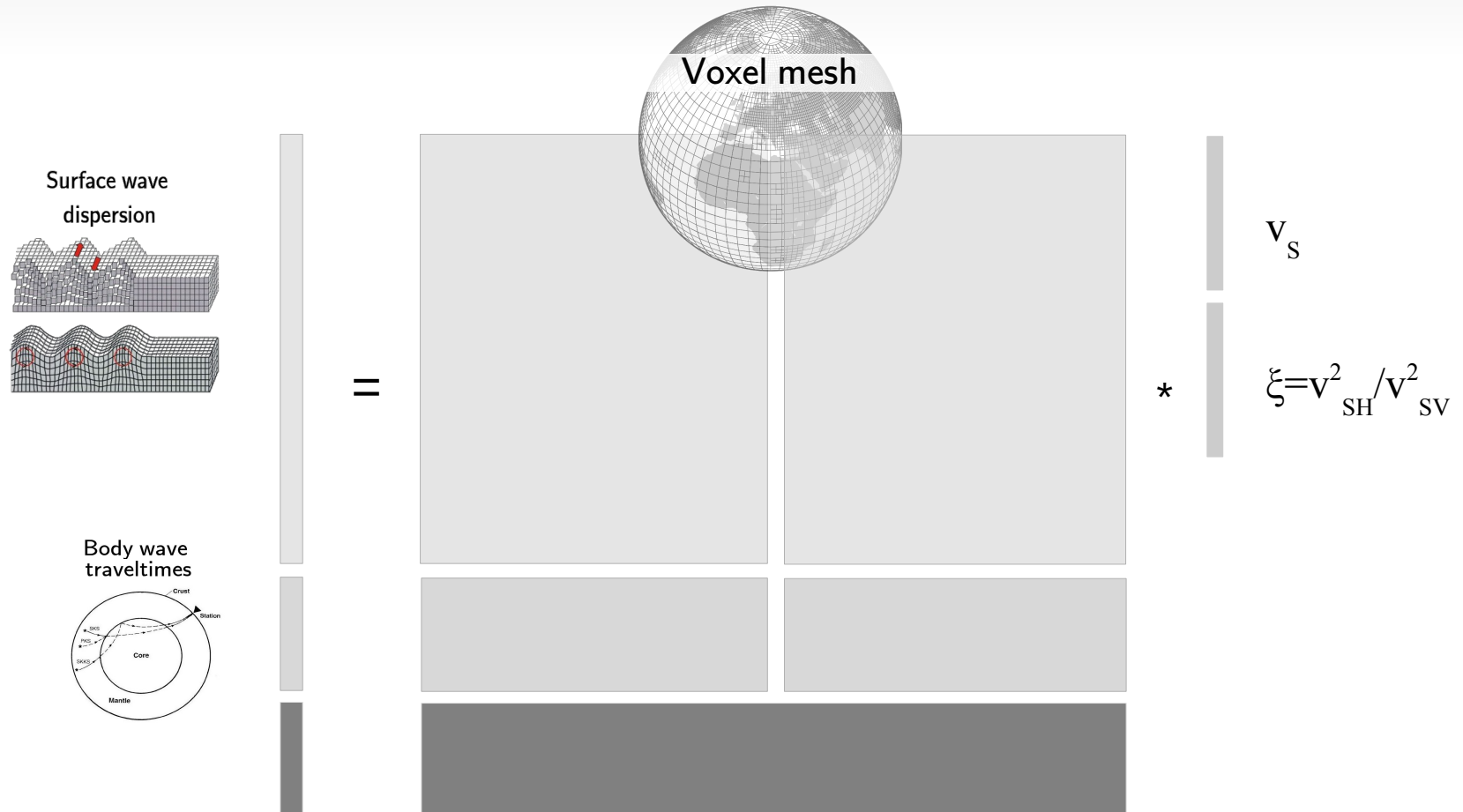
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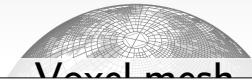
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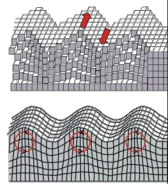


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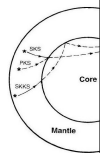
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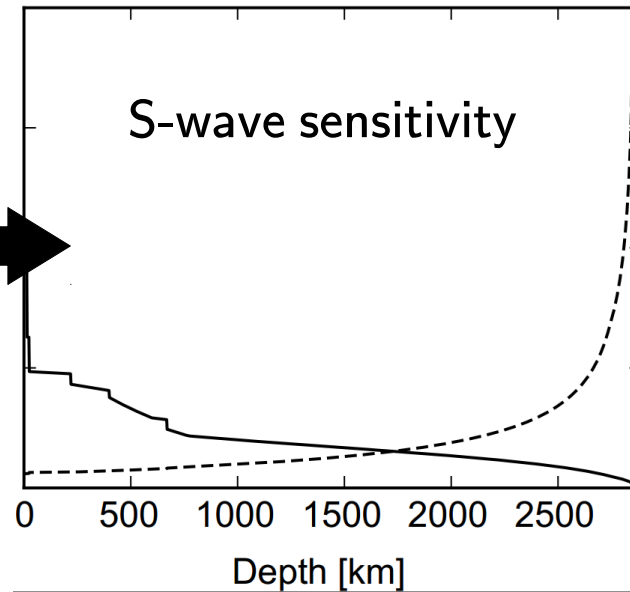
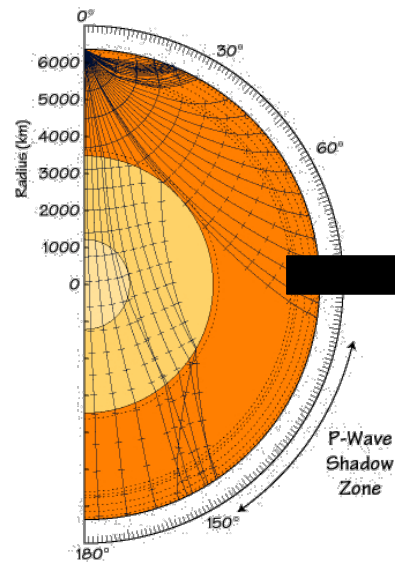
Surface wave dispersion



Body wave travel



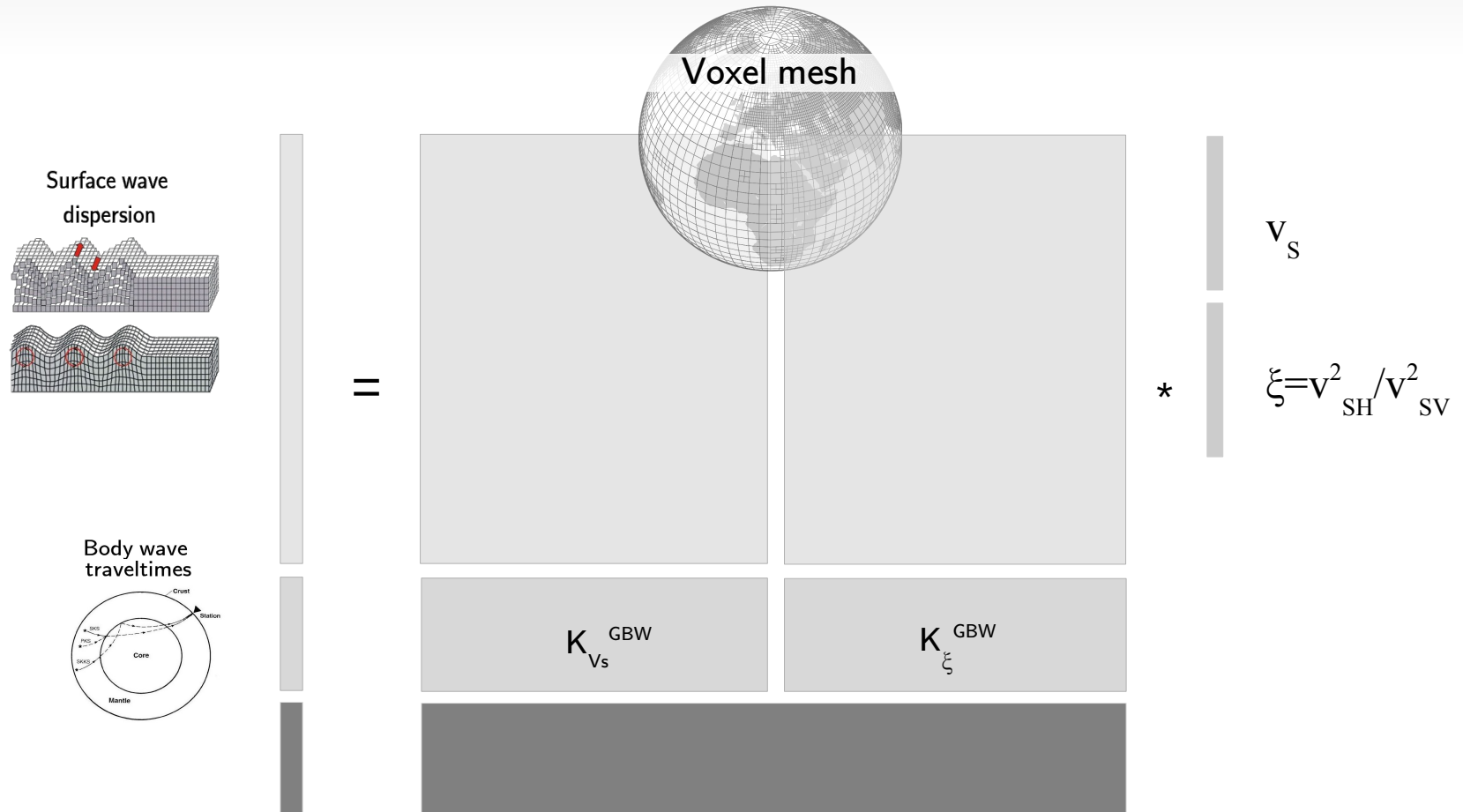
## Body wave ray theory



$$S = \frac{v_{SH}^2}{v_{SV}^2}$$

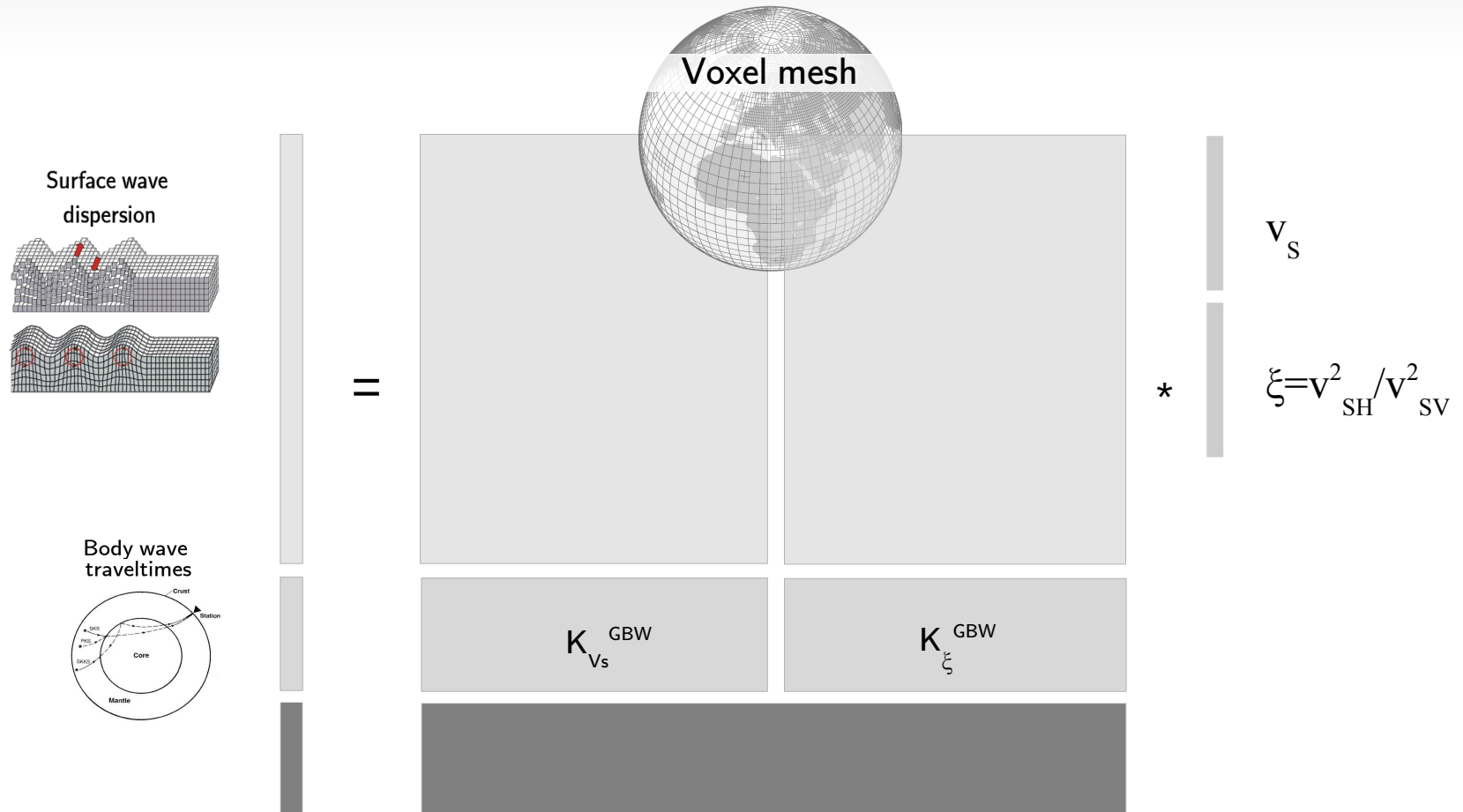
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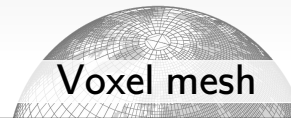
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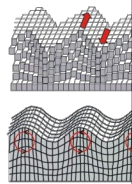


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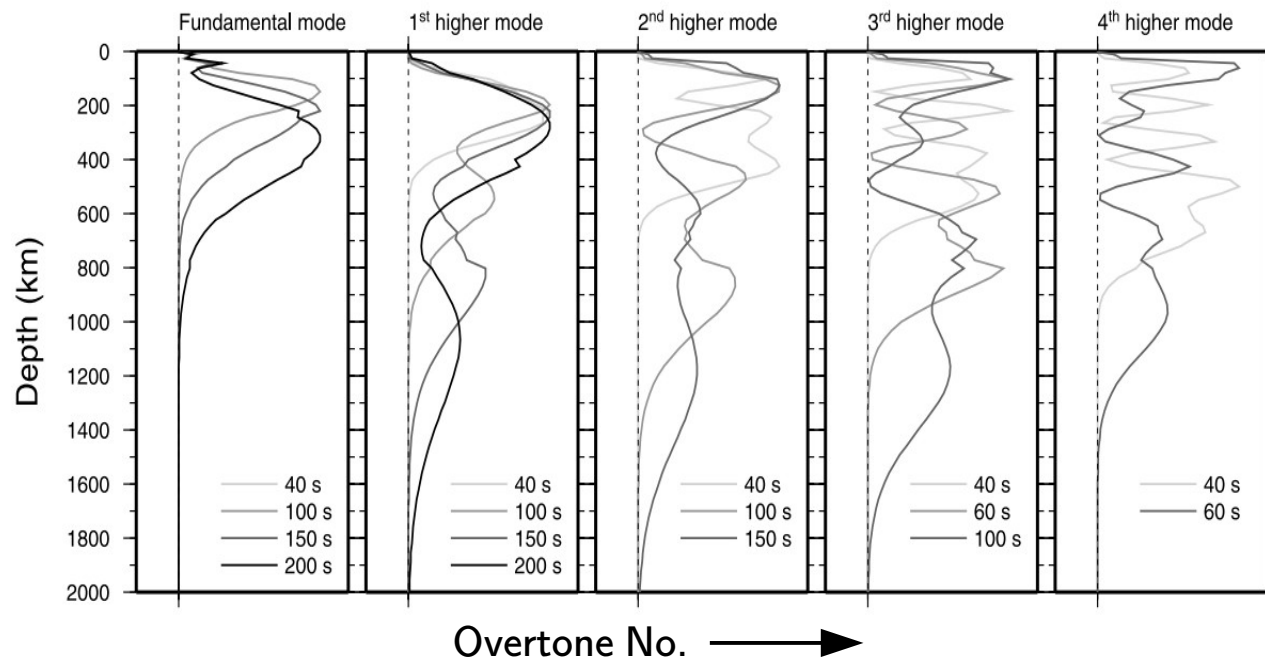
Surface  
disper



Body  
trav



### Surface-wave ray (JWKB) theory

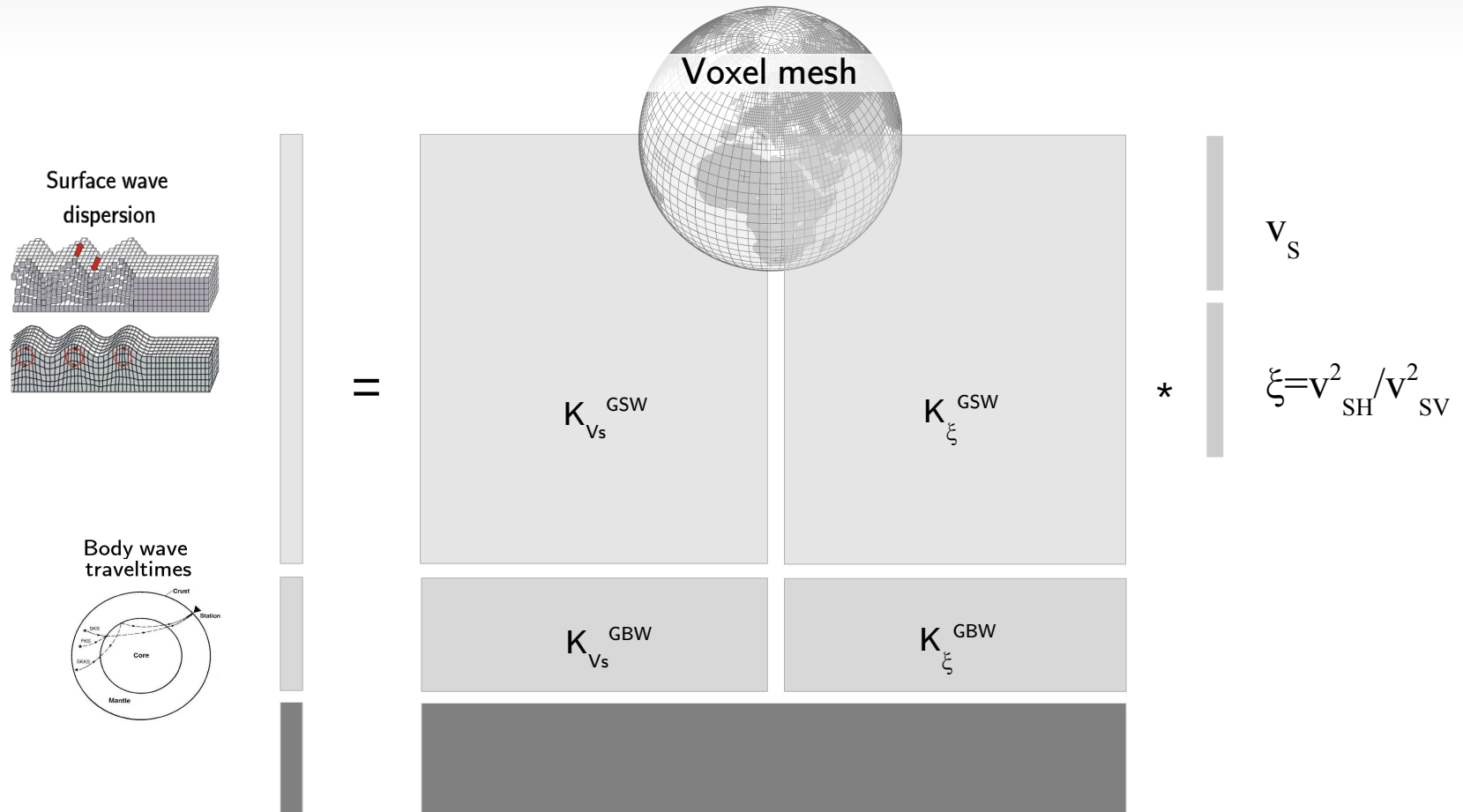


$\frac{c}{v^2}$   
SH SV



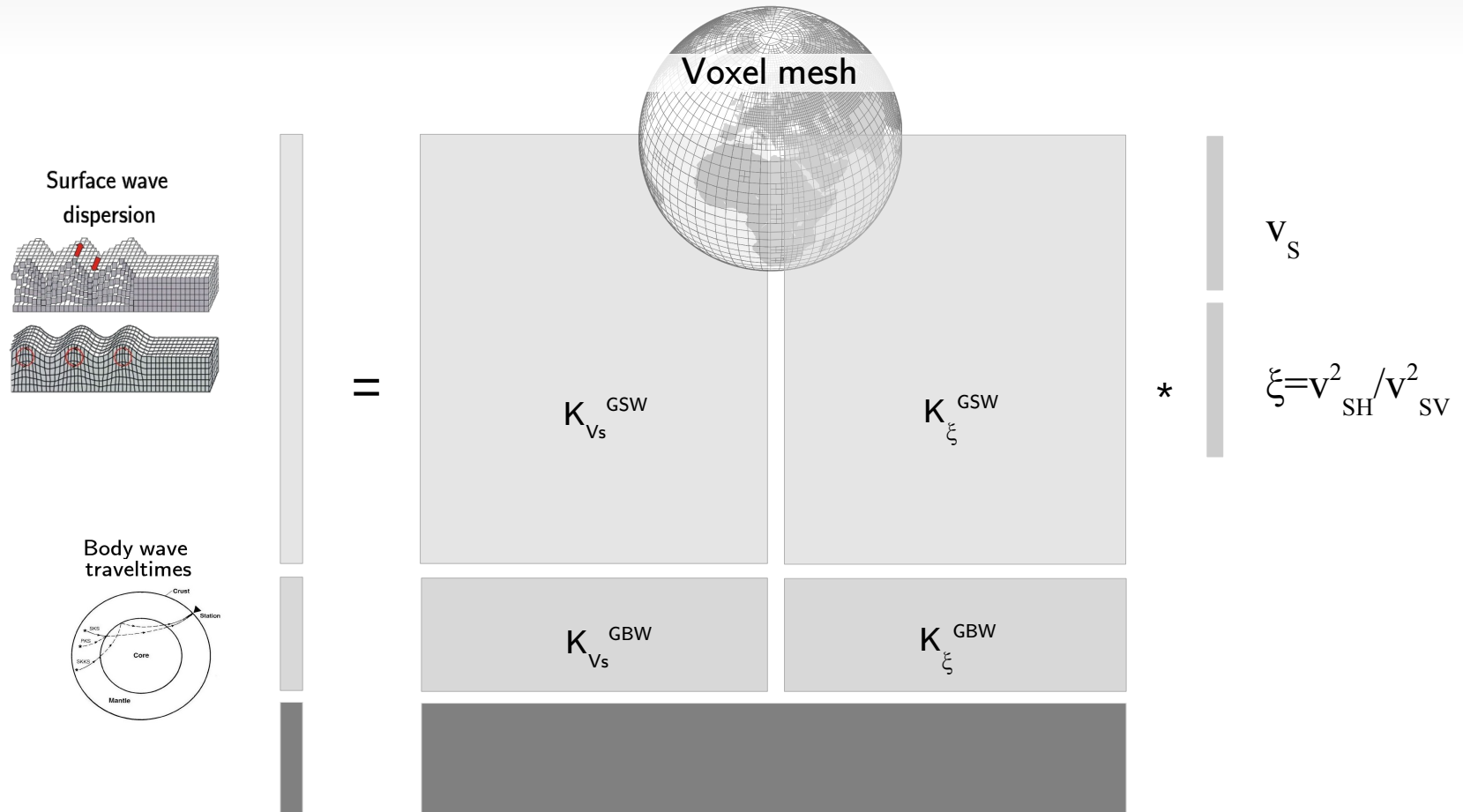
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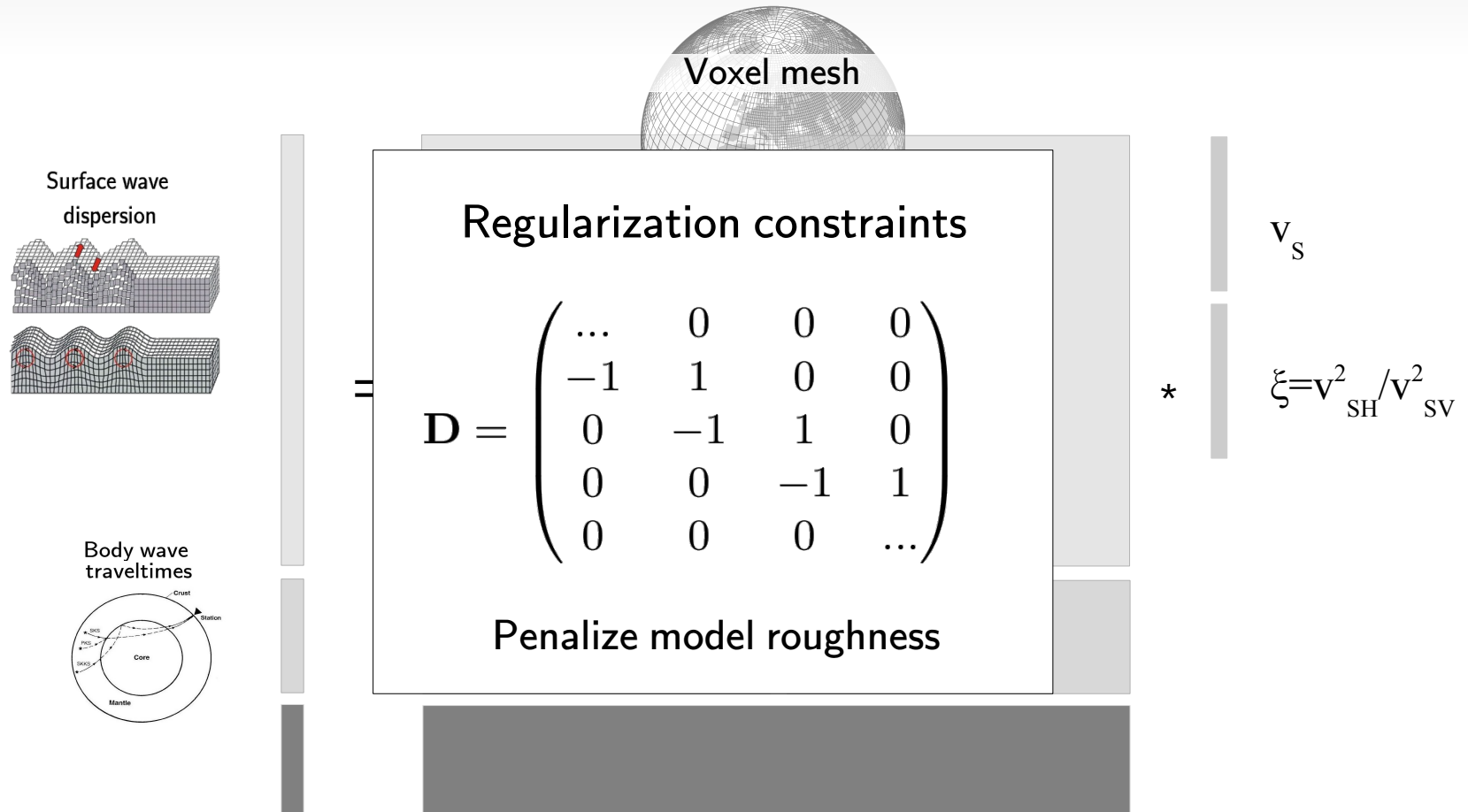
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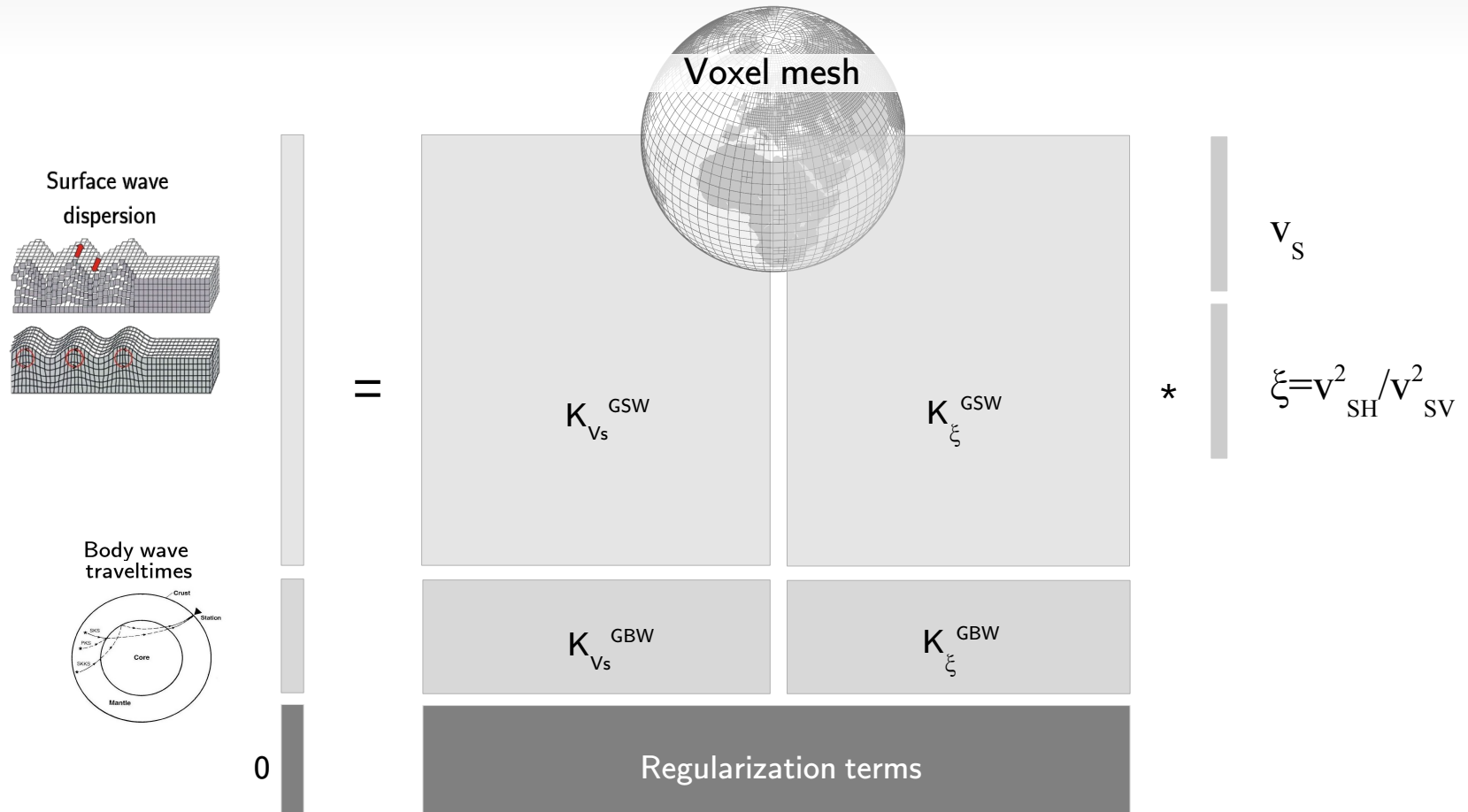
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## (2) Whole-mantle anisotropy

Solution of the inverse problem:  $\mathbf{d} = \mathbf{G}\mathbf{m}$

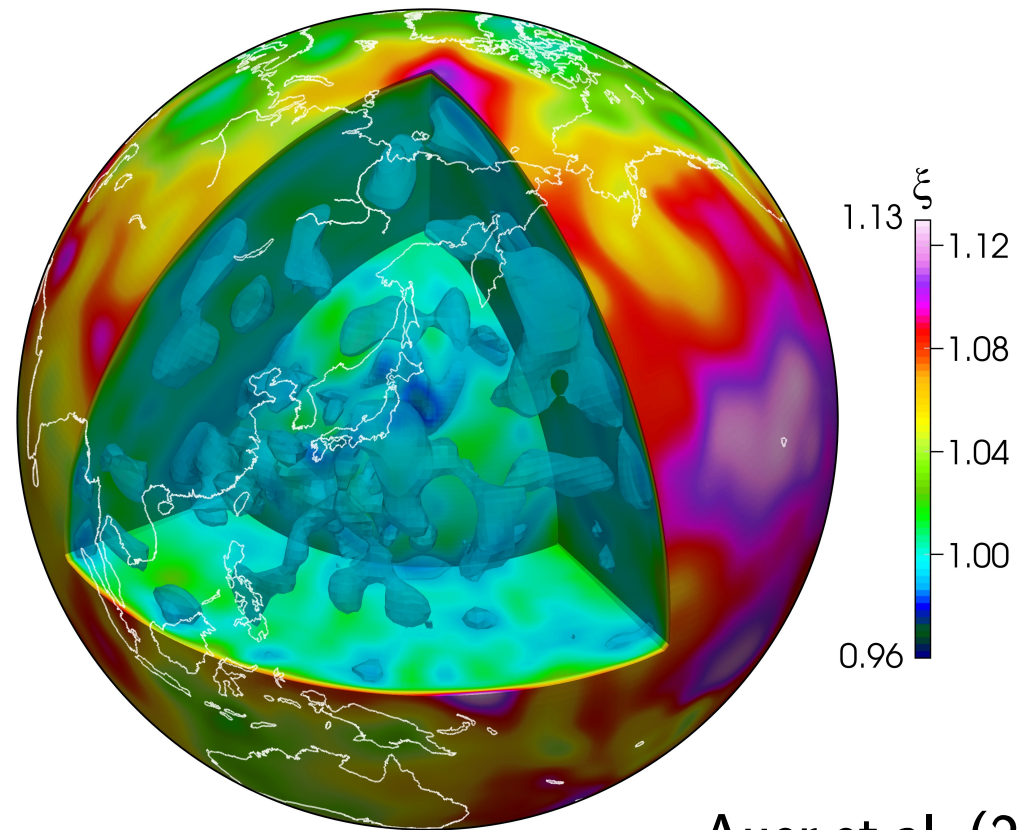
We solve the normal equations ...

$$\delta\mathbf{m} = [\mathbf{G}^T \mathbf{C}_d^{-1} \mathbf{G} + \mathbf{C}_m^{-1}] \mathbf{G}^T \mathbf{C}_d^{-1} \delta\mathbf{d}$$

Using a tailored parallel solver based on PETSc  
available @ <https://github.com/auerl/petscinv>

## (2) Whole-mantle anisotropy

Savani: a new, global, adaptive resolution model of radial shear-wave anisotropy



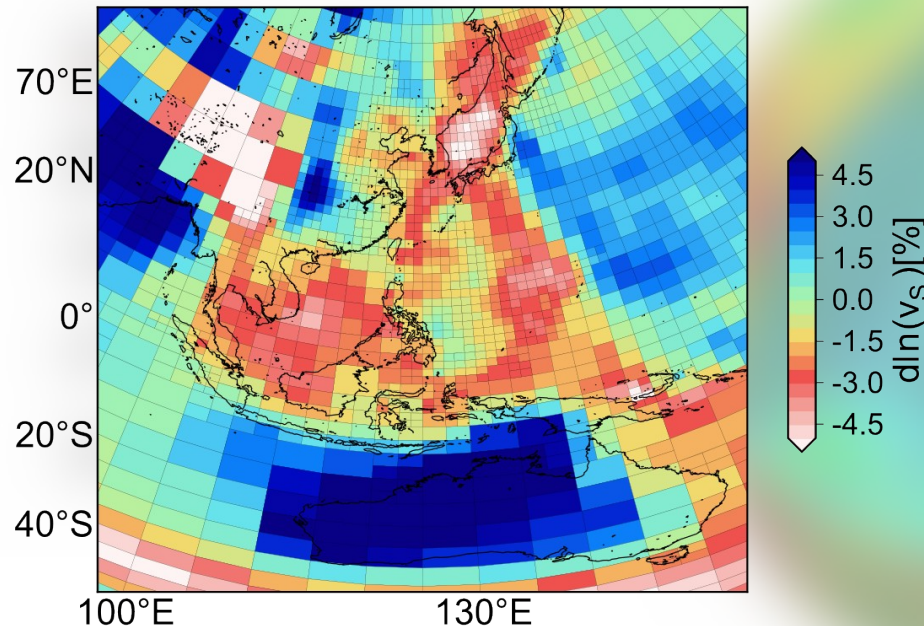
Auer et al. (2014)

## (2) Whole-mantle anisotropy

### Key features of savani

1) Variable, adaptive-resolution

Depth: 80 km



2) Radial anisotropy in the entire mantle

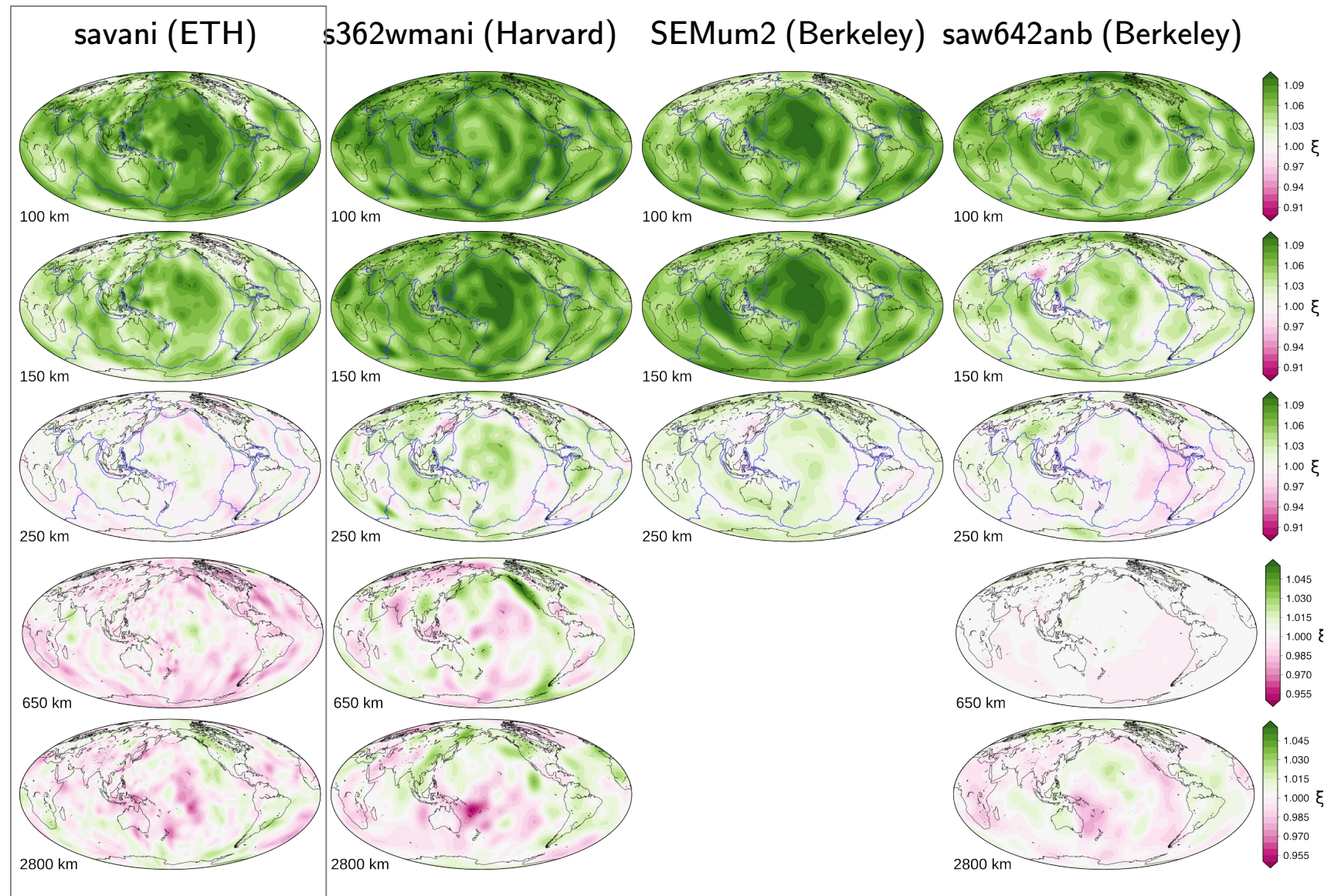
$$\xi = \left( \frac{v_{SH}}{v_{SV}} \right)^2$$

3) Extreme data diversity:  
Numerous previously  
published datasets

Download @ <http://n.ethz.ch/~auerl>

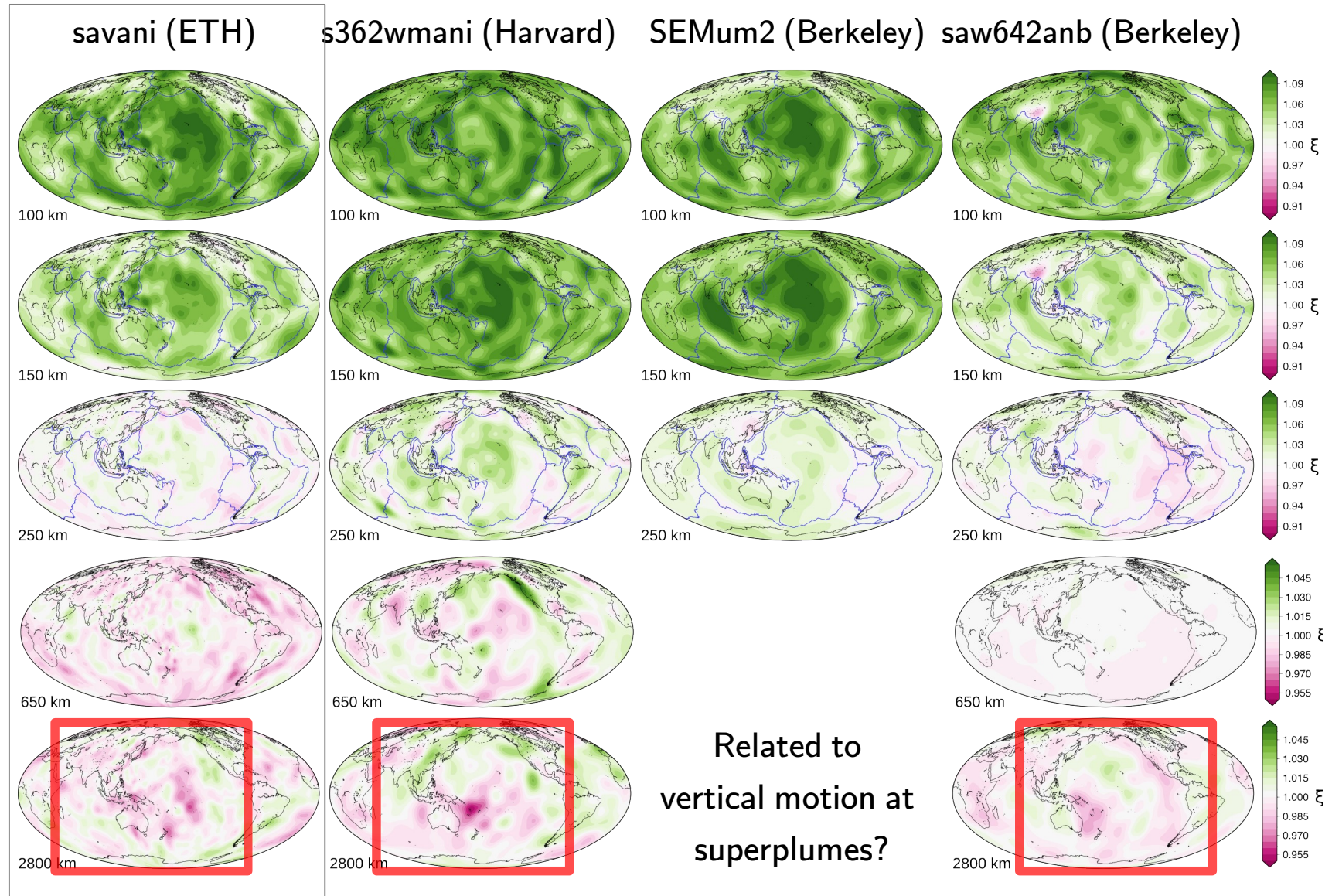


## (2) Model comparison, anisotropic



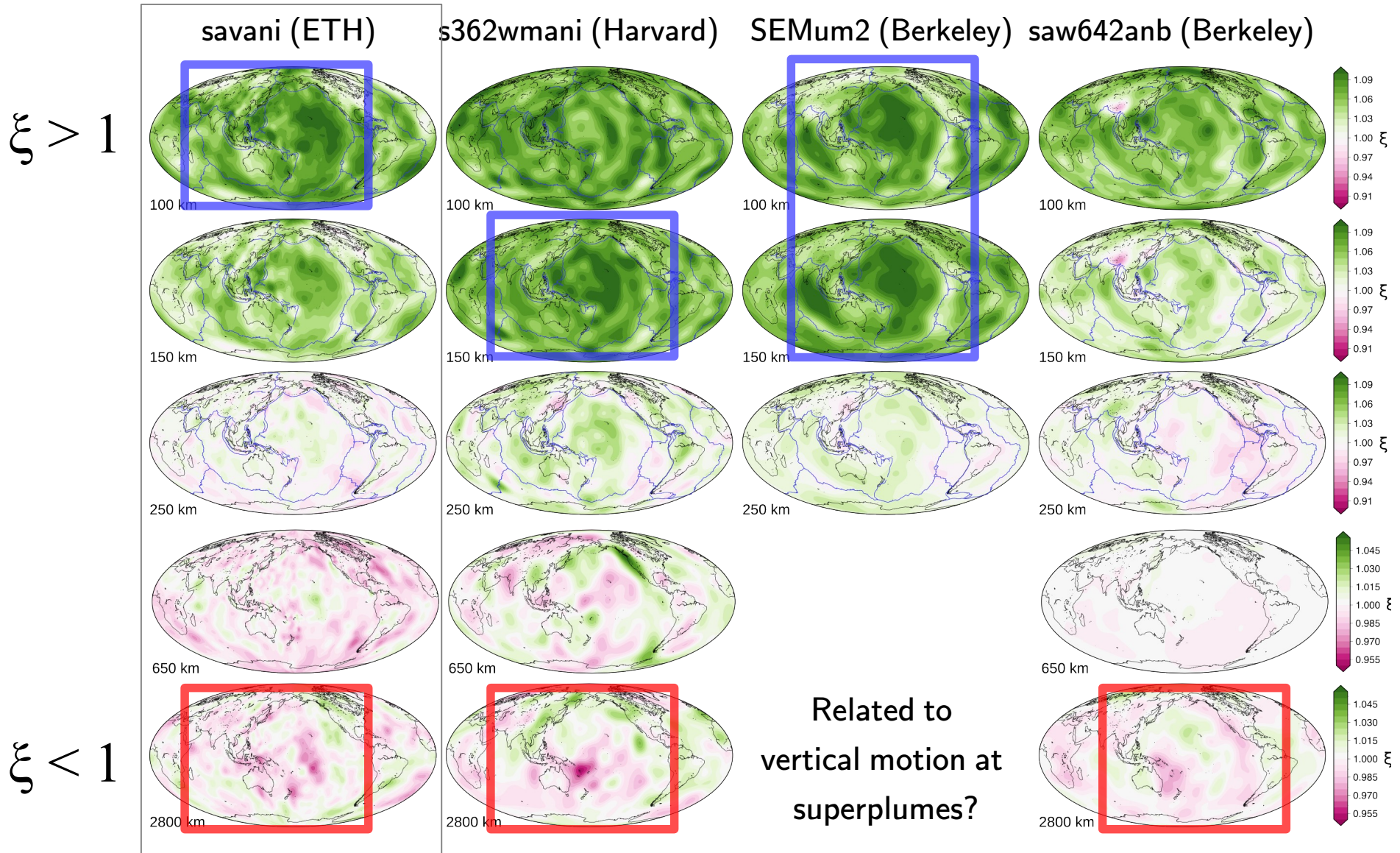


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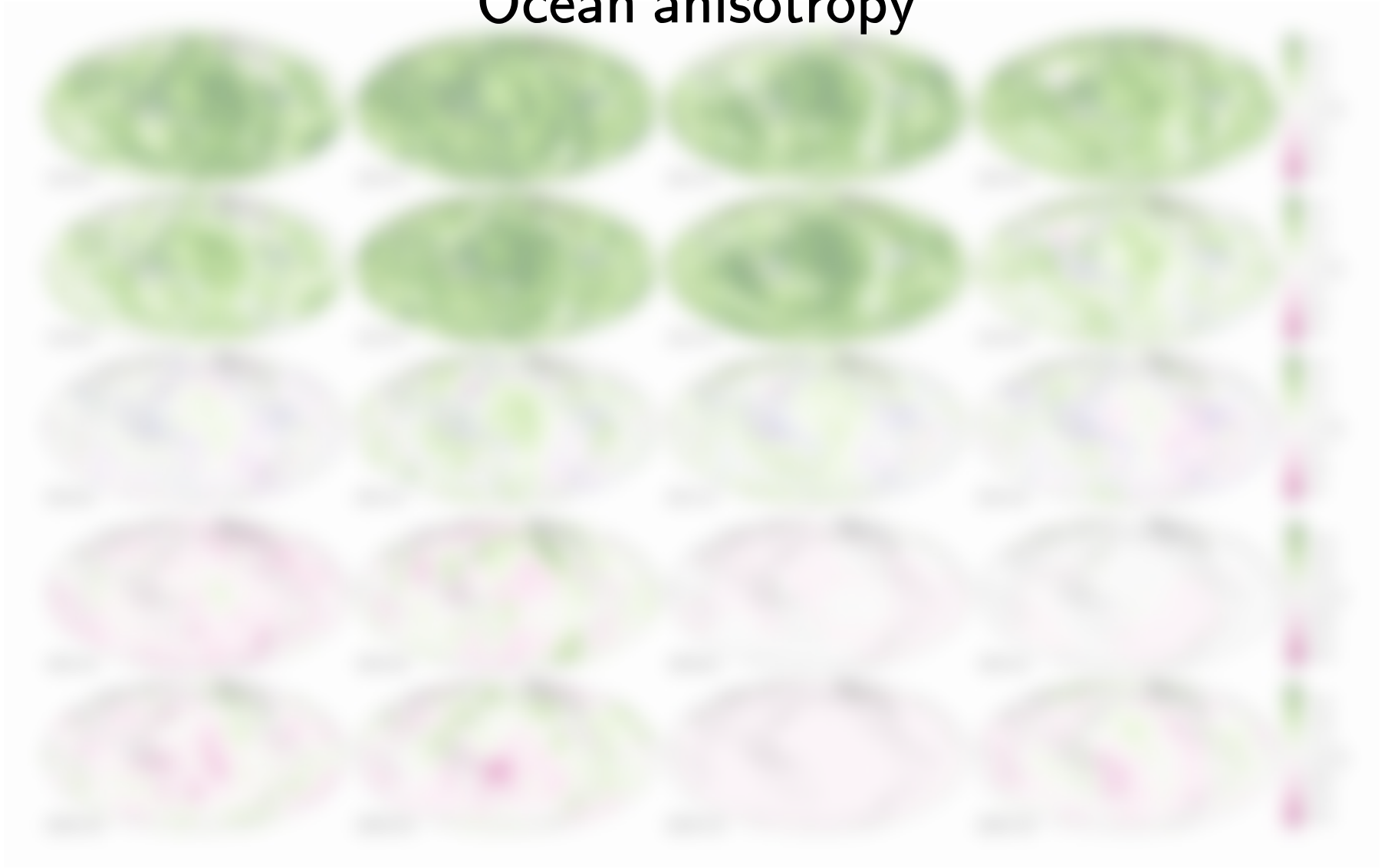


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### Ocean anisotropy

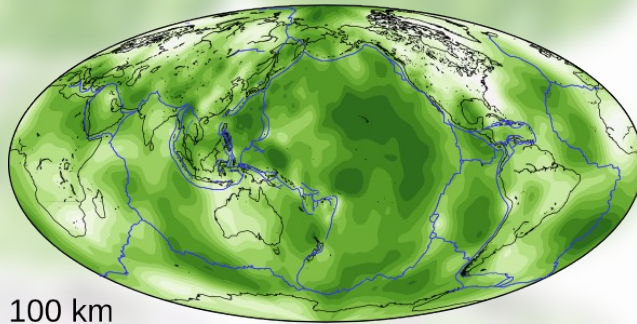




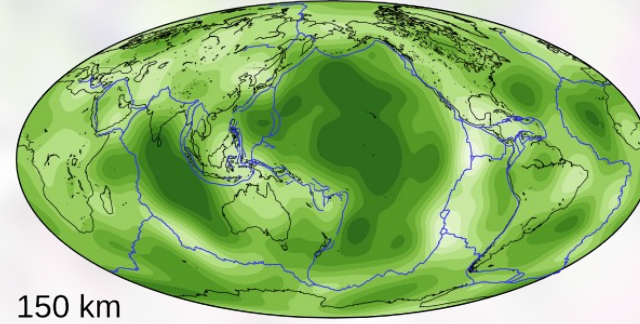
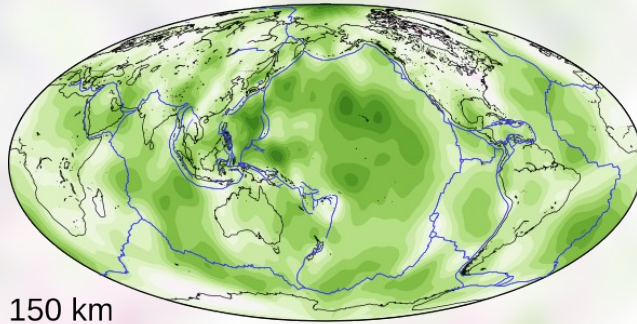
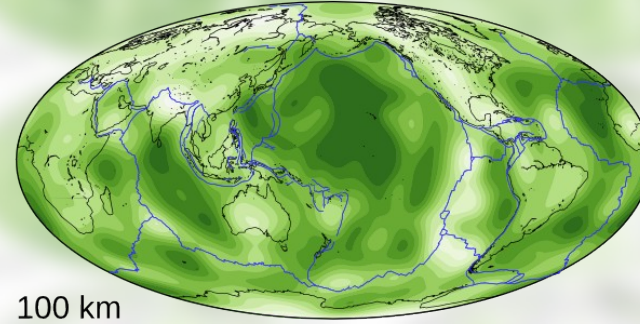
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savani



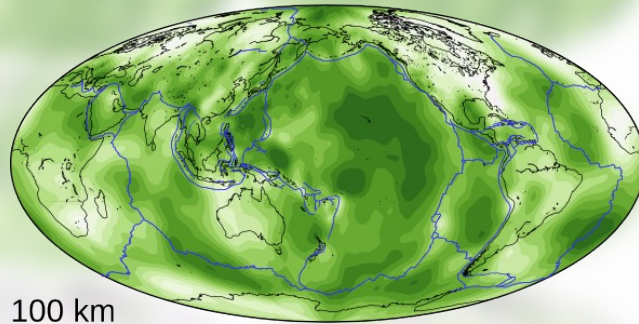
SEMum2



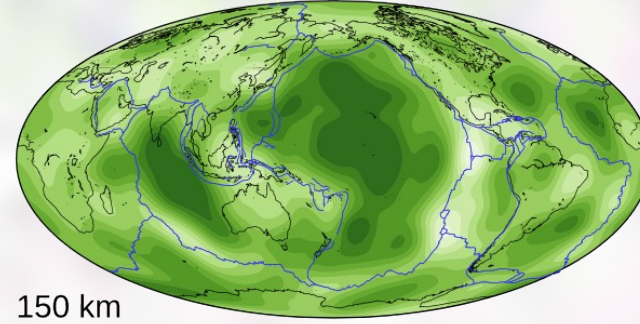
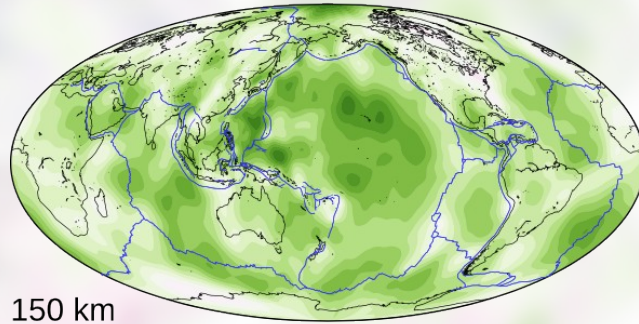
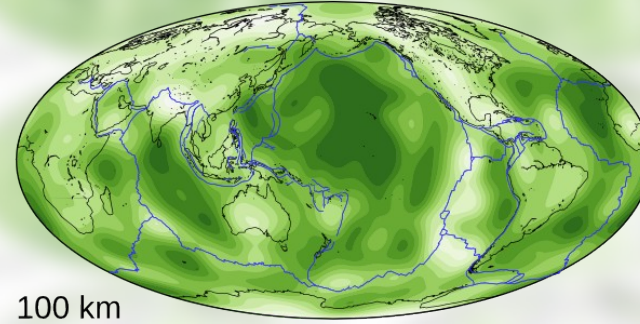
## (2) Model comparison, anisotropic

### Ocean anisotropy

savani



SEMum2



What is the  $\xi$  anomaly and how does it relate to upper-mantle dynamics?

## (2) Thesis outline

Chap.

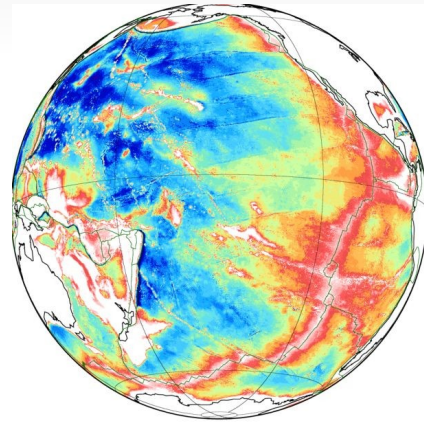
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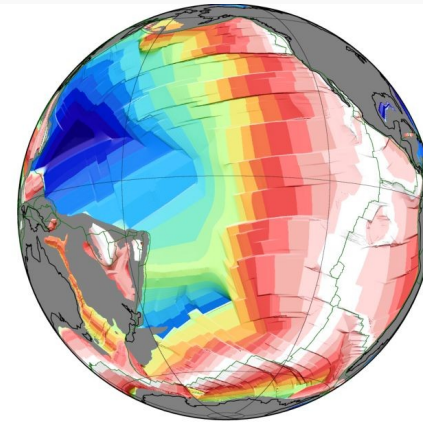
# (3) Oceanic upper-mantle structure

## Other geophysical observations

Bathymetry  $Z$

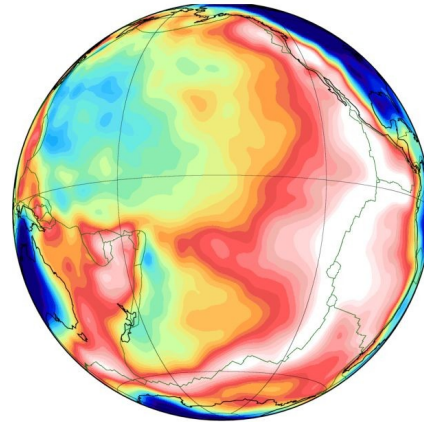


$\tau$  sea-floor ages

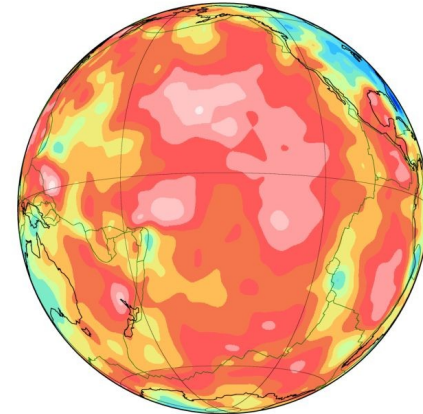


This study:

Wavespeeds  $V_S$



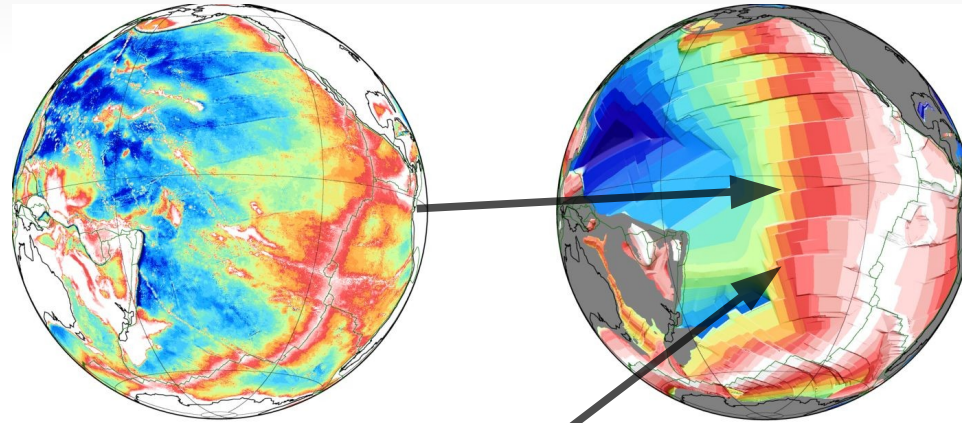
$\xi$  anisotropy



# (3) Oceanic upper-mantle structure

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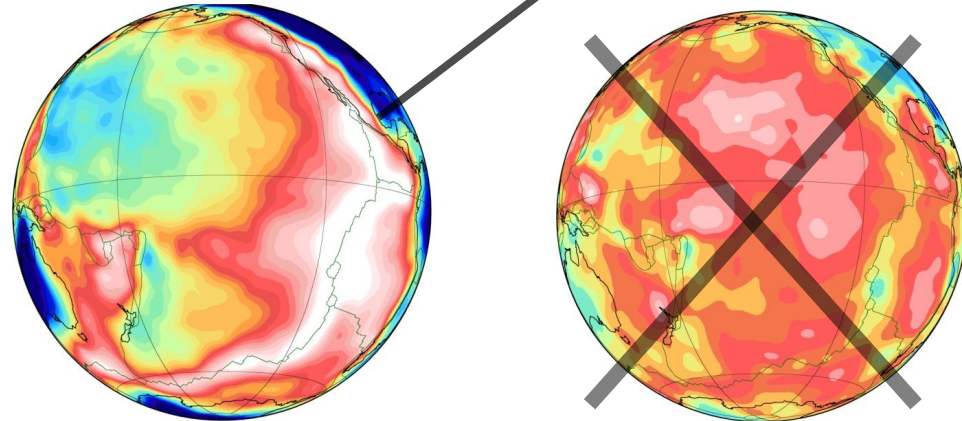


$\tau$  sea-floor ages

**Half-space  
Cooling (HSC)!**

This study:

Wavespeeds  $V_S$

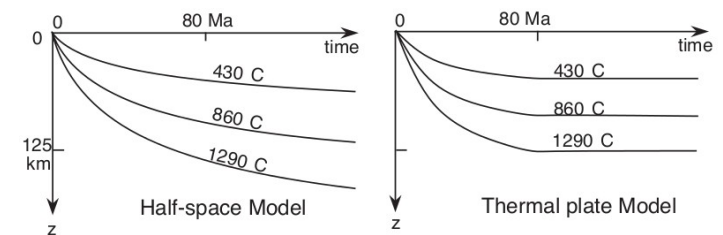
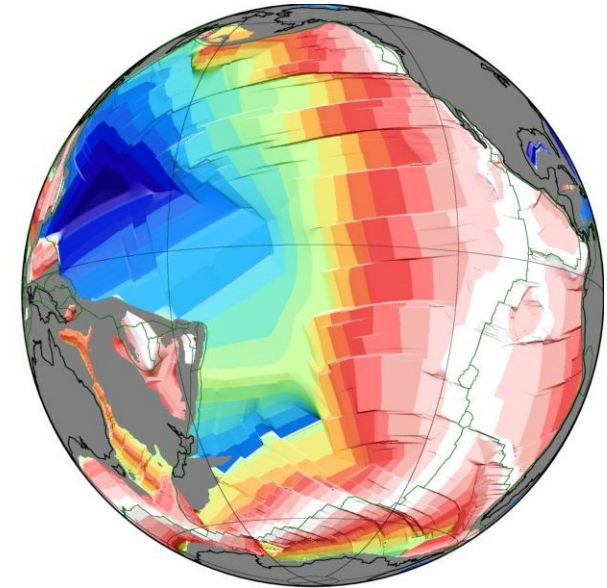


$\xi$  anisotropy

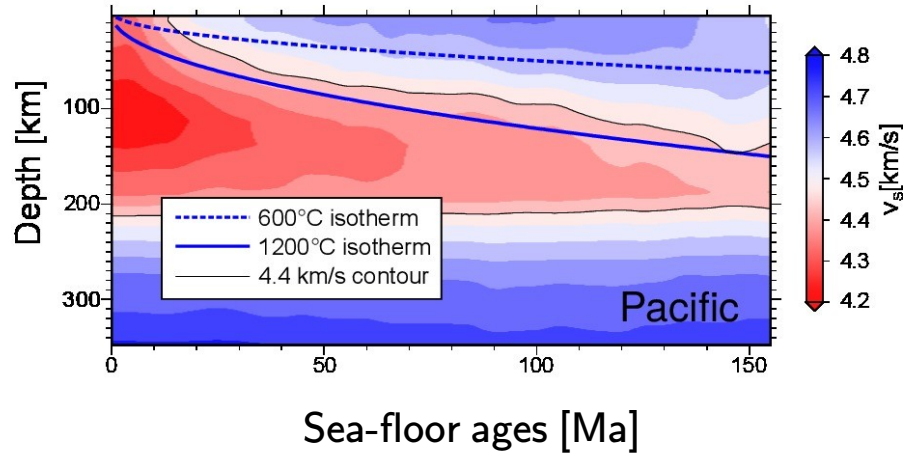


# (3) Age dependence, isotropic

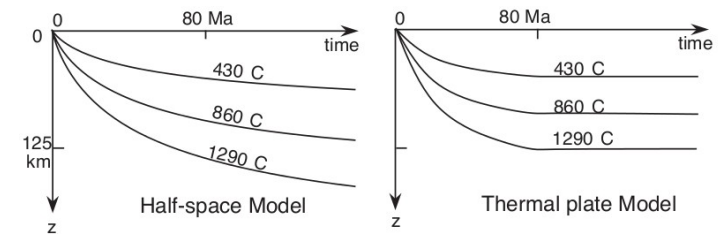
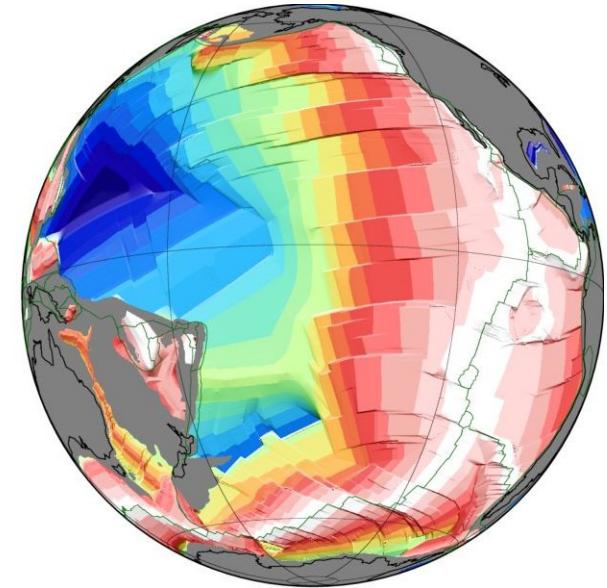
Reference  $v_s$  against sea-floor  
ages of the Pacific Ocean



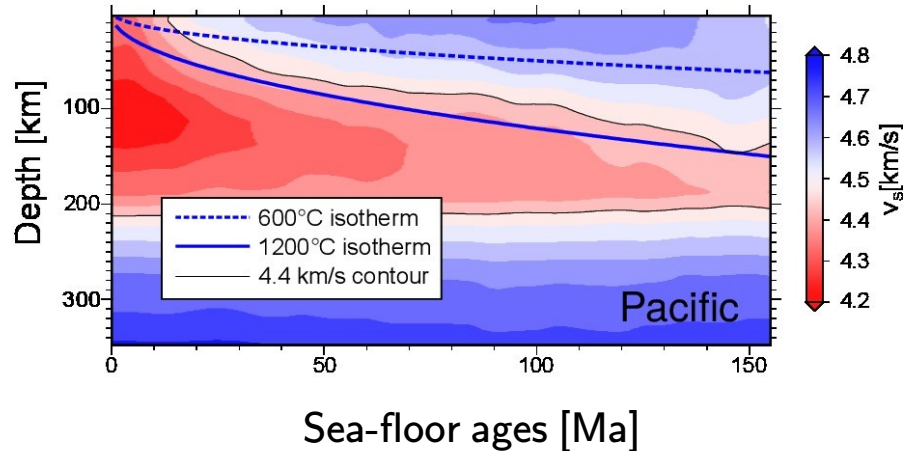
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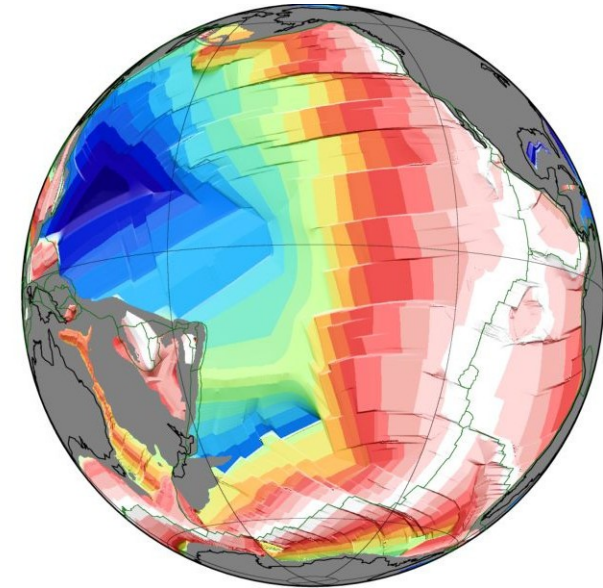
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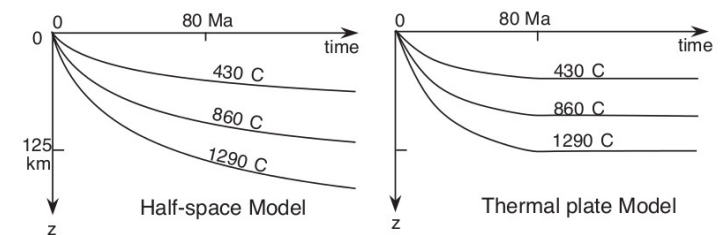
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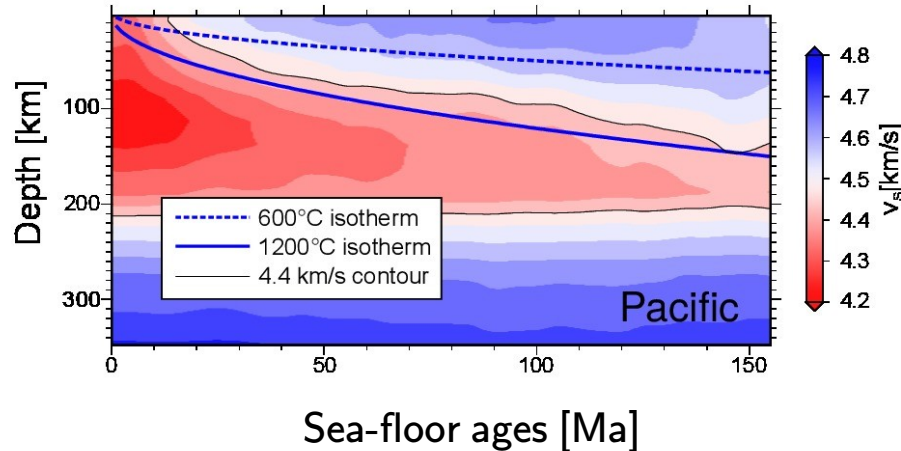
Reference  $v_s$  against sea-floor ages of the Pacific Ocean



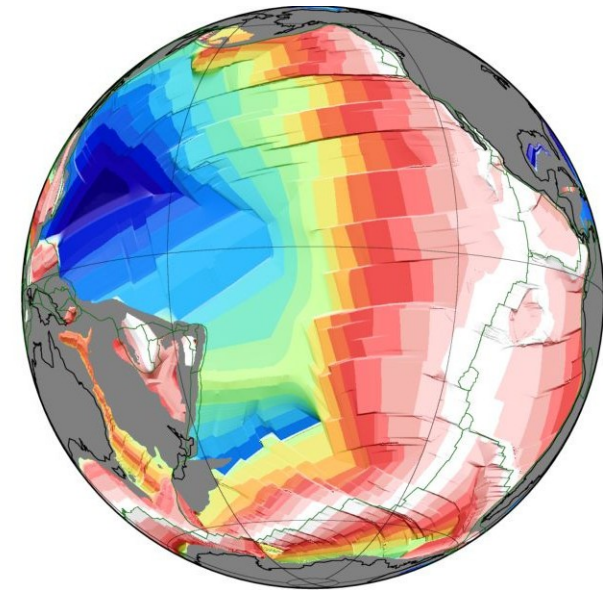
- Oceanic lithosphere agrees broadly with half-space cooling, LAB~1200°



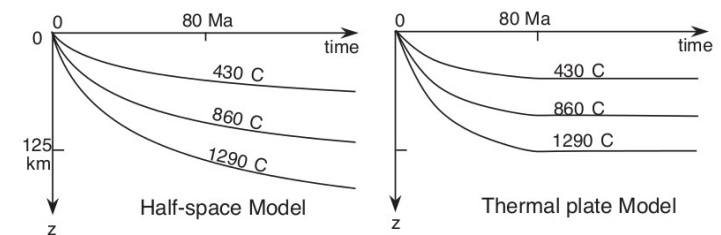
### (3) Age dependence, isotropic



Reference  $v_s$  against sea-floor ages of the Pacific Ocean

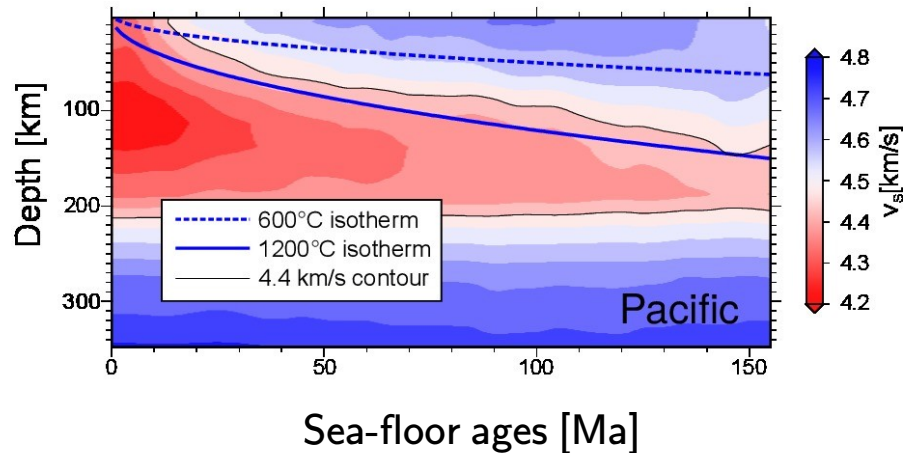


- Oceanic lithosphere agrees broadly with half-space cooling, LAB  $\sim 1200^\circ$
- **But:** anomaly in the Pacific at sea-floor ages of  $\sim 80$  million years

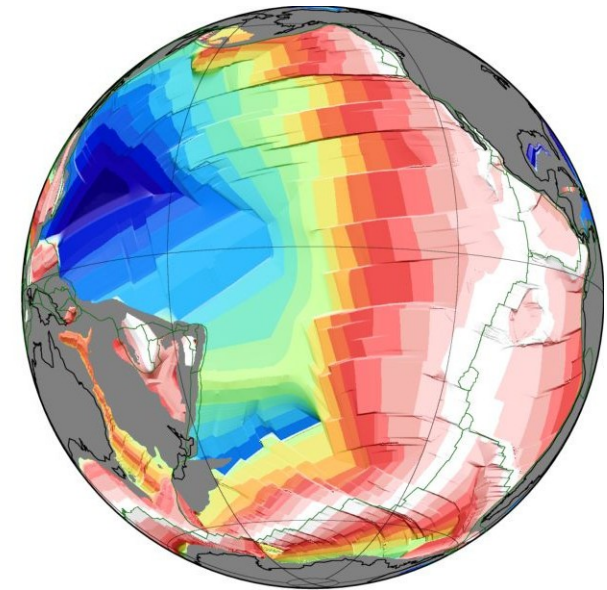




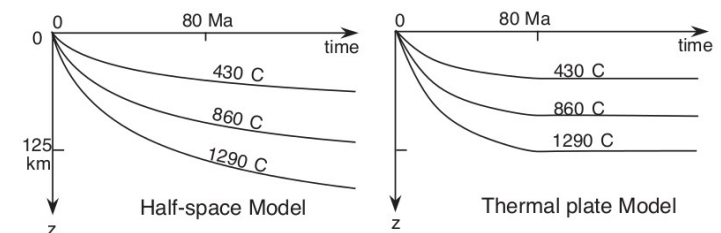
### (3) Age dependence, isotropic



Reference  $v_s$  against sea-floor ages of the Pacific Ocean



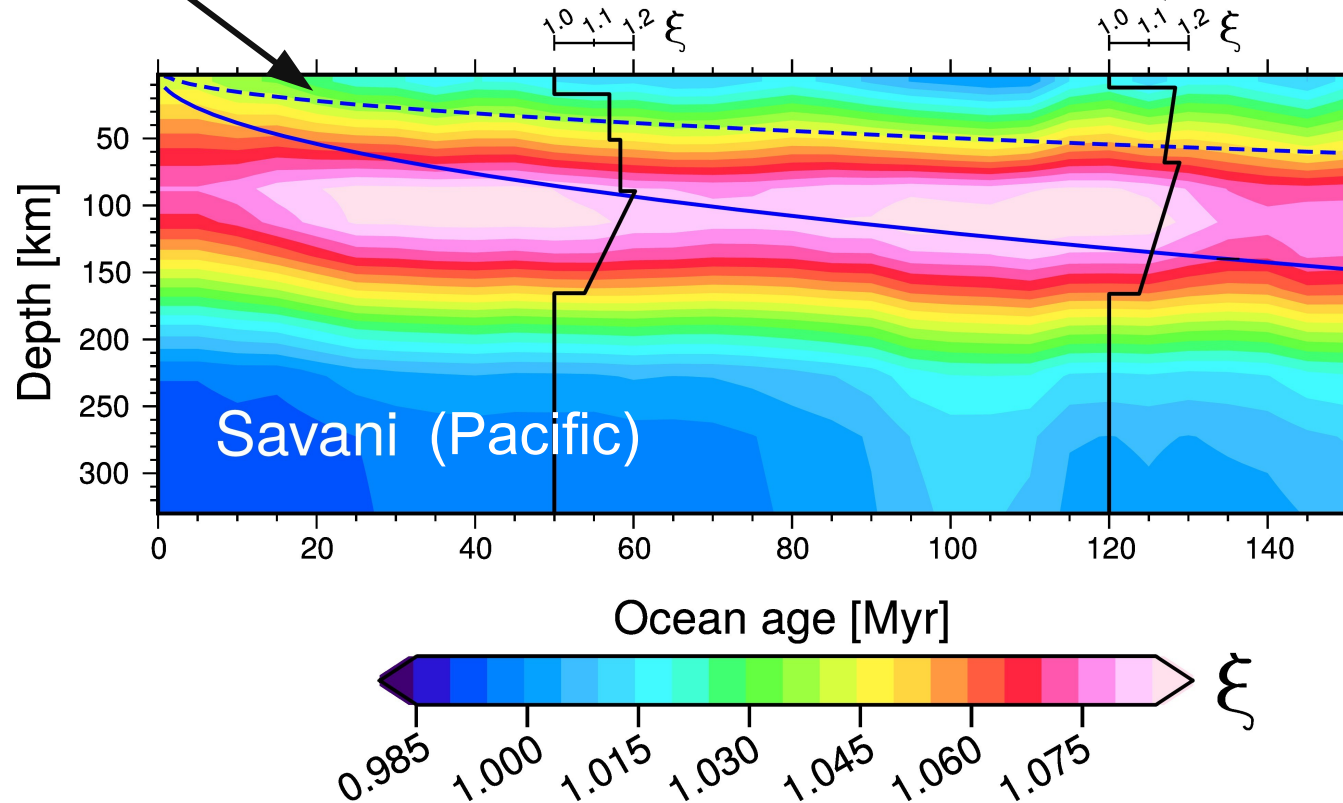
- Oceanic lithosphere agrees broadly with half-space cooling, LAB  $\sim 1200^\circ$
- **But:** anomaly in the Pacific at sea-floor ages of  $\sim 80$  million years
- Complexity beyond half-space cooling



# (3) Age dependence, anisotropic

600°C & 1200°C  
half-space cooling  
isotherms

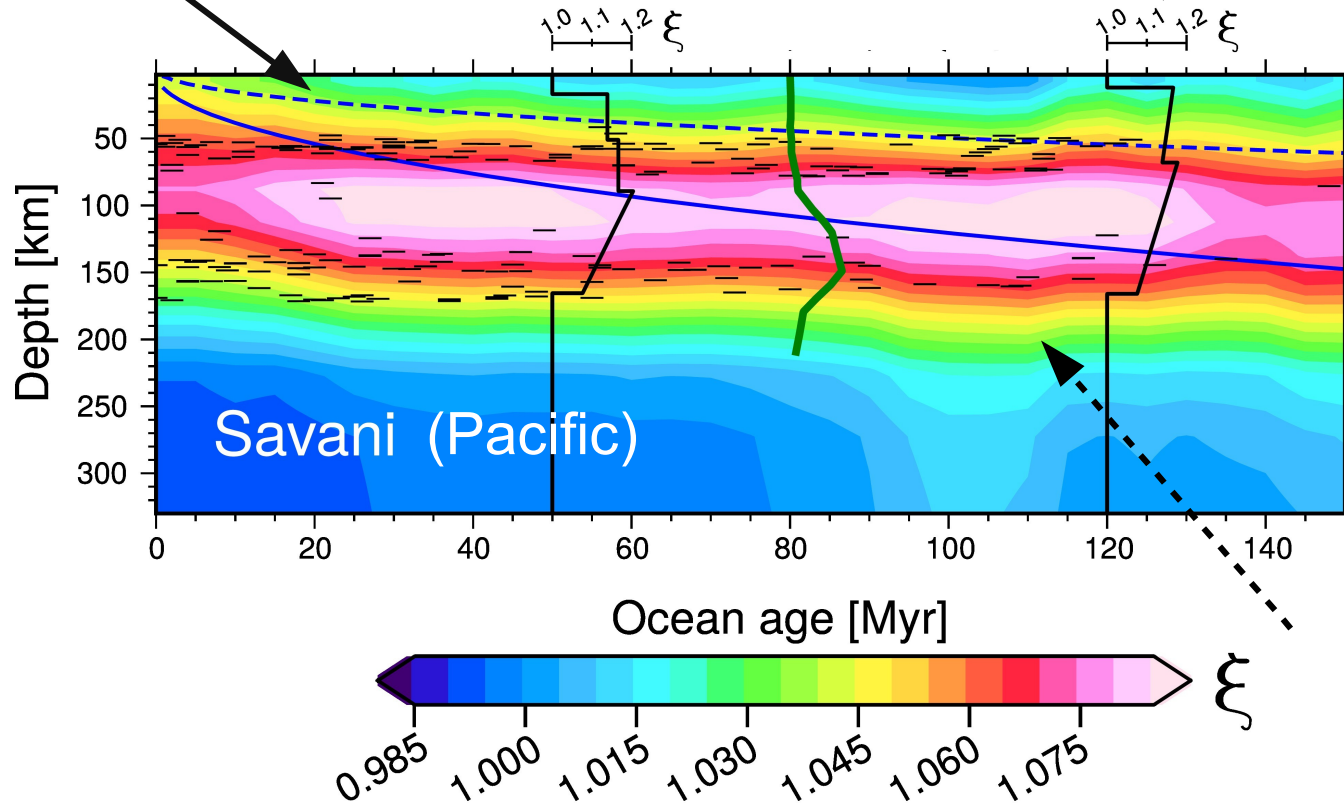
High-res. 1D models,  
(Gaherty 1996, 1999)



# (3) Age dependence, anisotropic

600°C & 1200°C  
half-space cooling  
isotherms

High-res. 1D models,  
(Gaherty 1996, 1999)



Savani (Pacific)

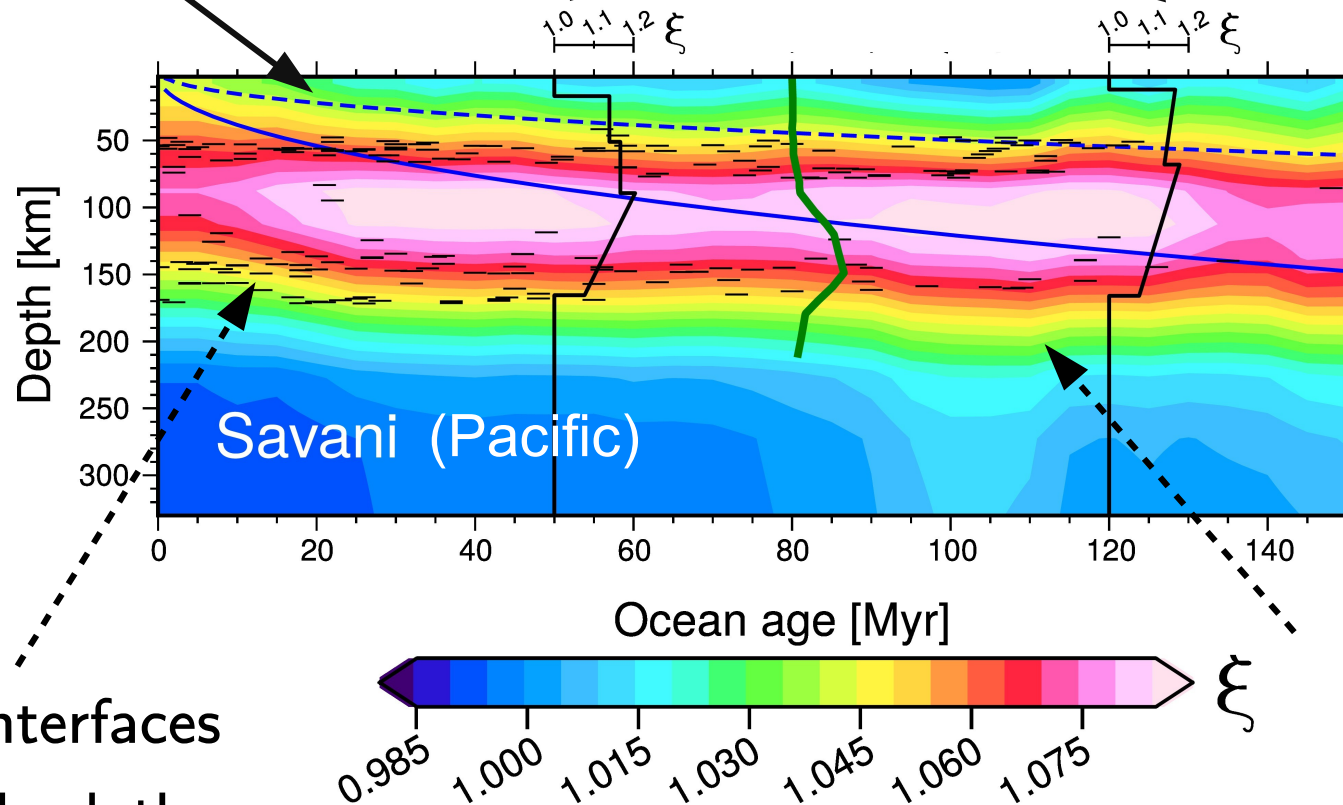
G = LAB?  
Oceanic L?

Receiver funct.  
& SS precursor  
depths, (Kawa-  
Katsu 2011 &  
Schmerr 2012)

# (3) Age dependence, anisotropic

600°C & 1200°C  
half-space cooling  
isotherms

High-res. 1D models,  
(Gaherty 1996, 1999)



G = LAB?

Oceanic L?

What are interfaces  
and  $\xi$ ? → Check the  
LPO hypothesis

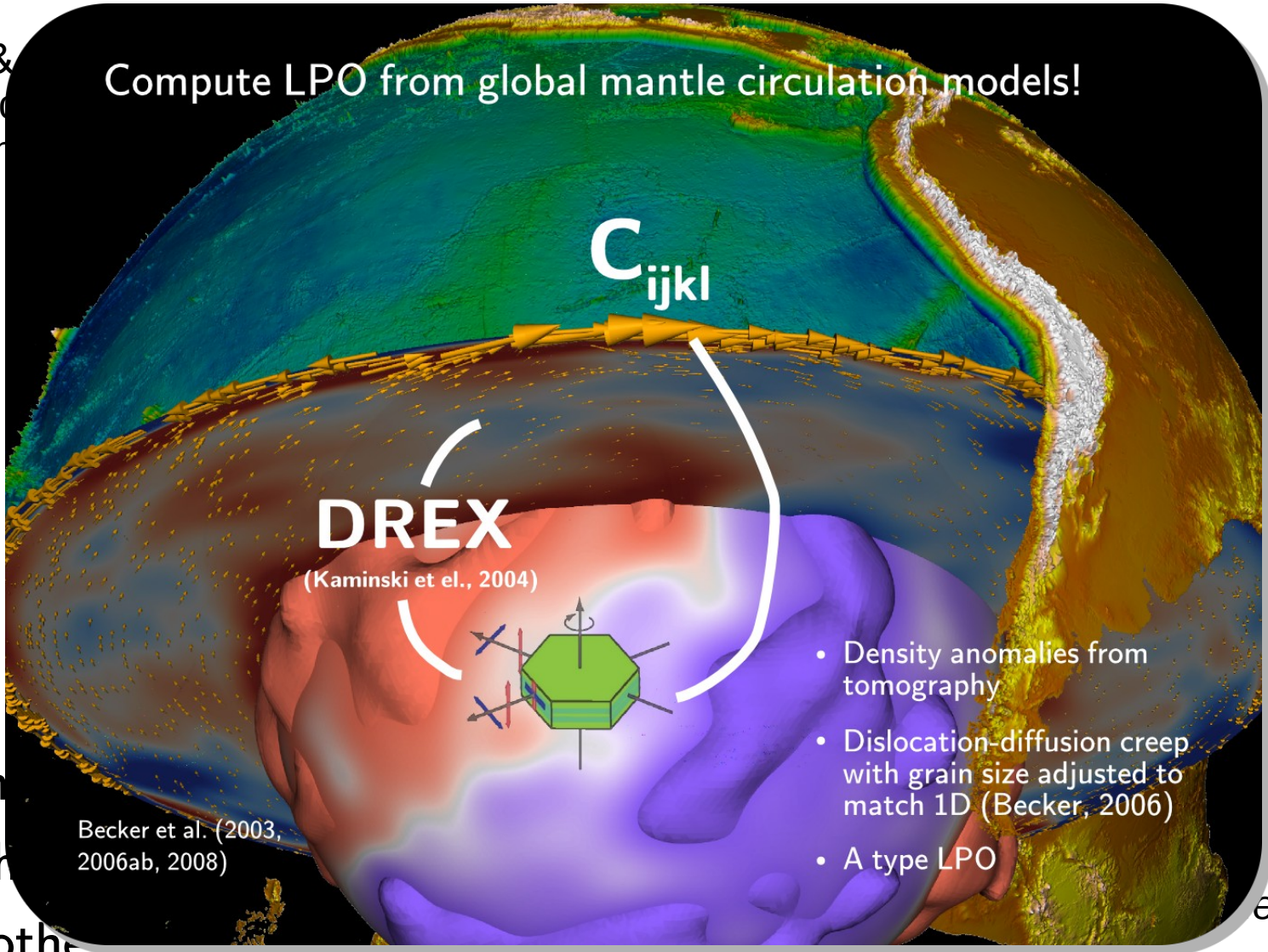
Receiver funct.  
& SS precursor  
depths, (Kawa-  
Katsu 2011 &  
Schmerr 2012)



# (3) Age dependence, anisotropic

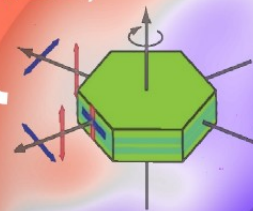
600°C &  
half-space  
isoth

Compute LPO from global mantle circulation models!



**DREX**

(Kaminski et al., 2004)



- Density anomalies from tomography
- Dislocation-diffusion creep with grain size adjusted to match 1D (Becker, 2006)
- A type LPO

Becker et al. (2003, 2006ab, 2008)

LAB?  
mechanic L?

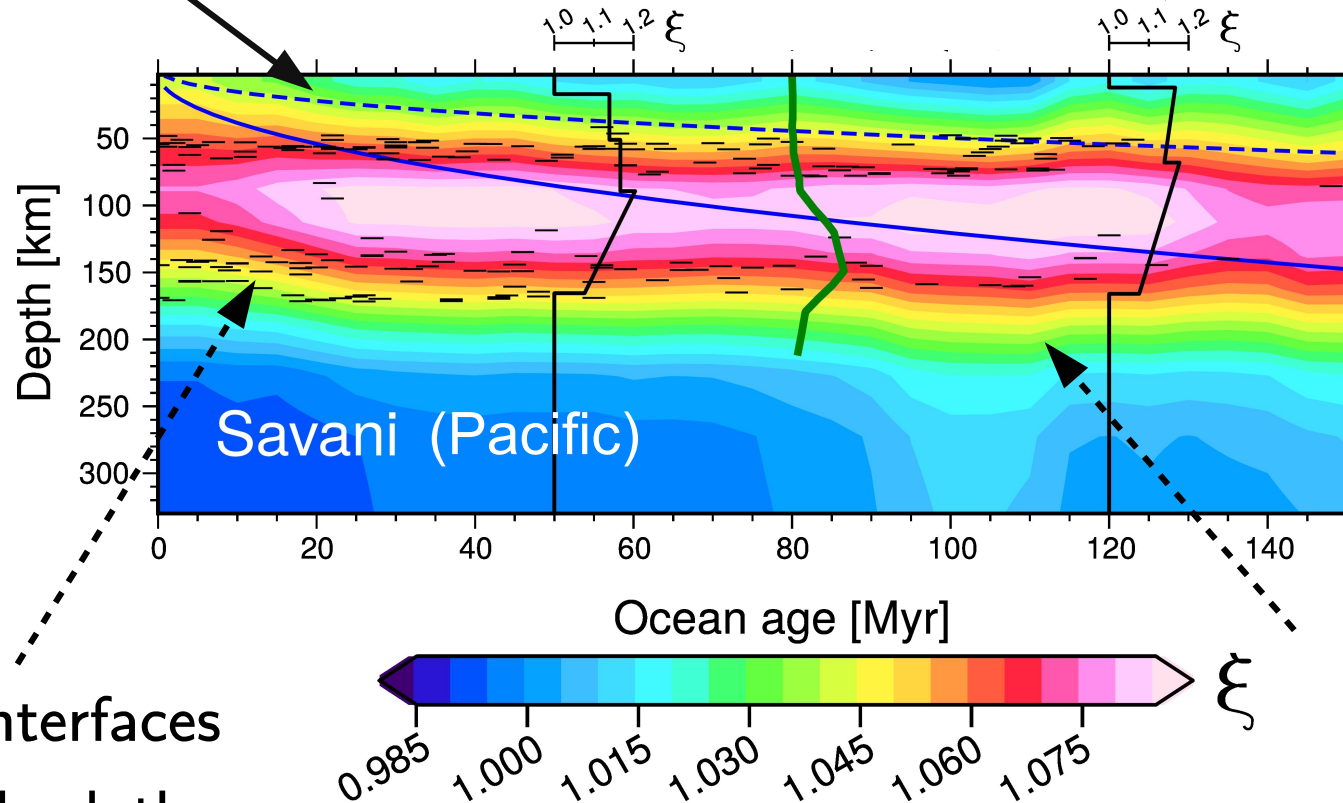
er funct.  
recursor  
, (Kawa-  
2011 &  
err 2012)

What are in  
and  $\xi$ ?  $\rightarrow$  Ch  
LPO hypothesis

# (3) Age dependence, anisotropic

600°C & 1200°C  
half-space cooling  
isotherms

High-res. 1D models,  
(Gaherty 1996, 1999)



G = LAB?

Oceanic L?

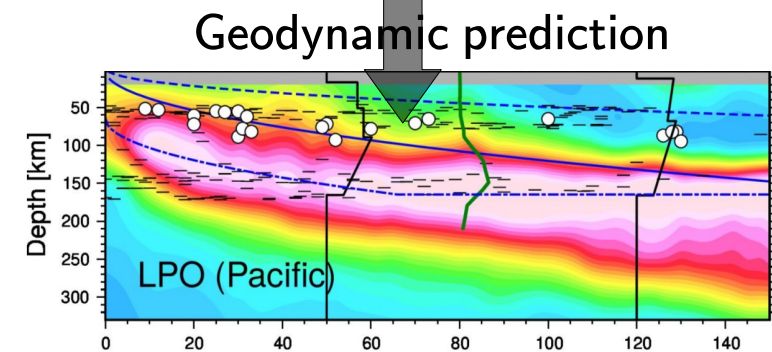
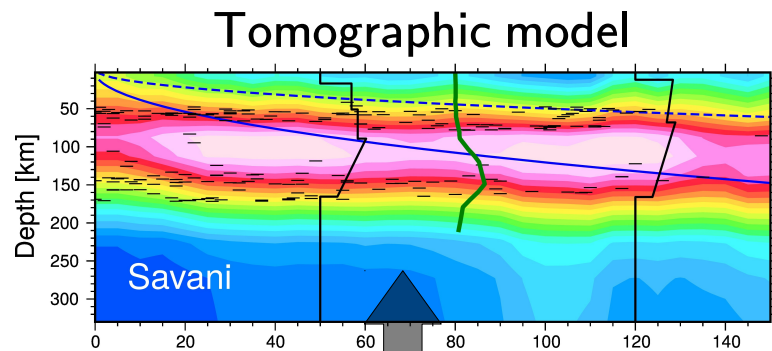
What are interfaces  
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Receiver funct.  
& SS precursor  
depths, (Kawa-  
Katsu 2011 &  
Schmerr 2012)

# (3) Age dependence, anisotropic

## Comparison with LPO-based prediction

### Radial anisotropy comparison (this study)

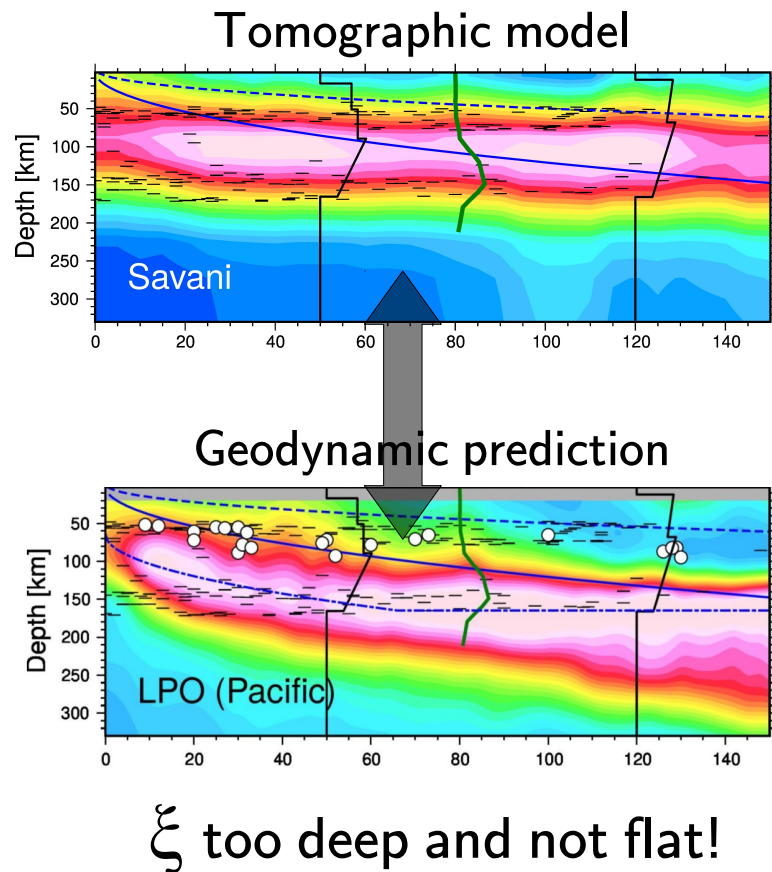


$\xi$  too deep and not flat!

# (3) Age dependence, anisotropic

## Comparison with LPO-based prediction

### Radial anisotropy comparison (this study)



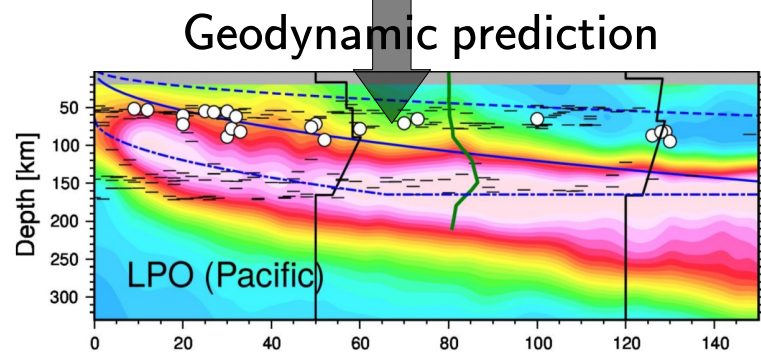
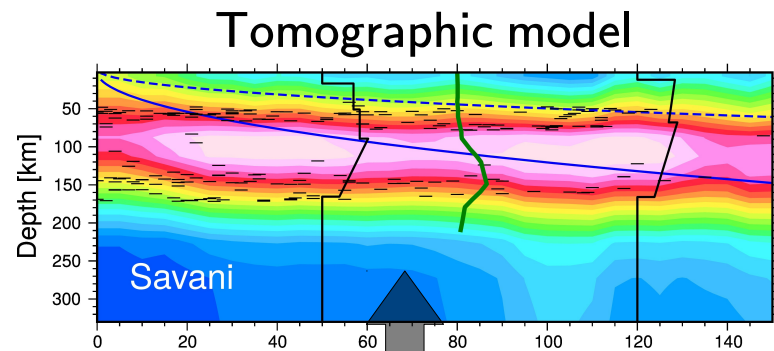
### Azimuthal anisotropy, Becker et al. (2015)



# (3) Age dependence, anisotropic

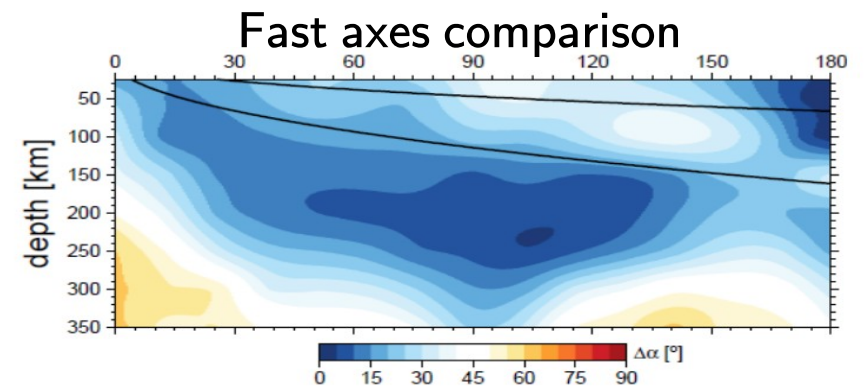
## Comparison with LPO-based prediction

### Radial anisotropy comparison (this study)



$\xi$  too deep and not flat!

### Azimuthal anisotropy, Becker et al. (2015)



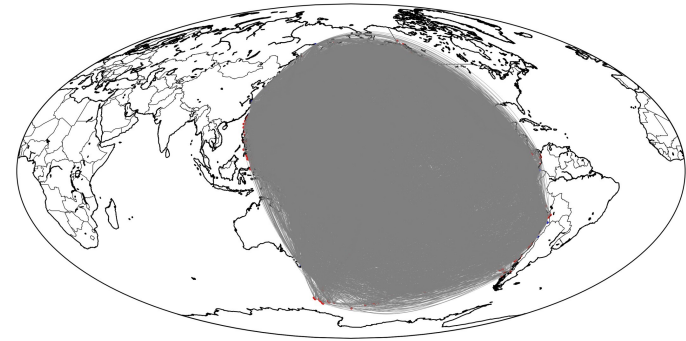
Tomographic and geodynamic  
fast axes directions **match** underneath  
the 1200°C isotherm!

## (3) Age dependence, anisotropic

- Flat  $\xi$  artifact or penalizing model roughness?
- **Idea:** Circumvent regularization by conducting “probabilistic” **hypothesis tests** → solve forward problem for conceptual geodynamic models of  $\xi$  and monitor data fit
- Are flat or age-dependent models preferred?

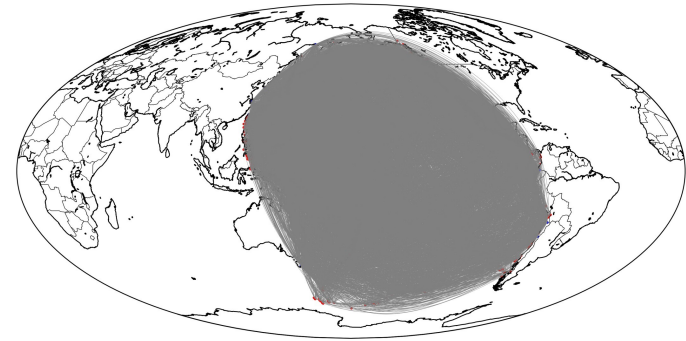
# (3) Hypothesis test: flat layer

Fit (VR) for  
Pacific dataset: 
$$VR = 1 - \frac{|\mathbf{G} \cdot \mathbf{m} - \mathbf{d}_{obs}|}{|\mathbf{d}_{obs}|}$$



# (3) Hypothesis test: flat layer

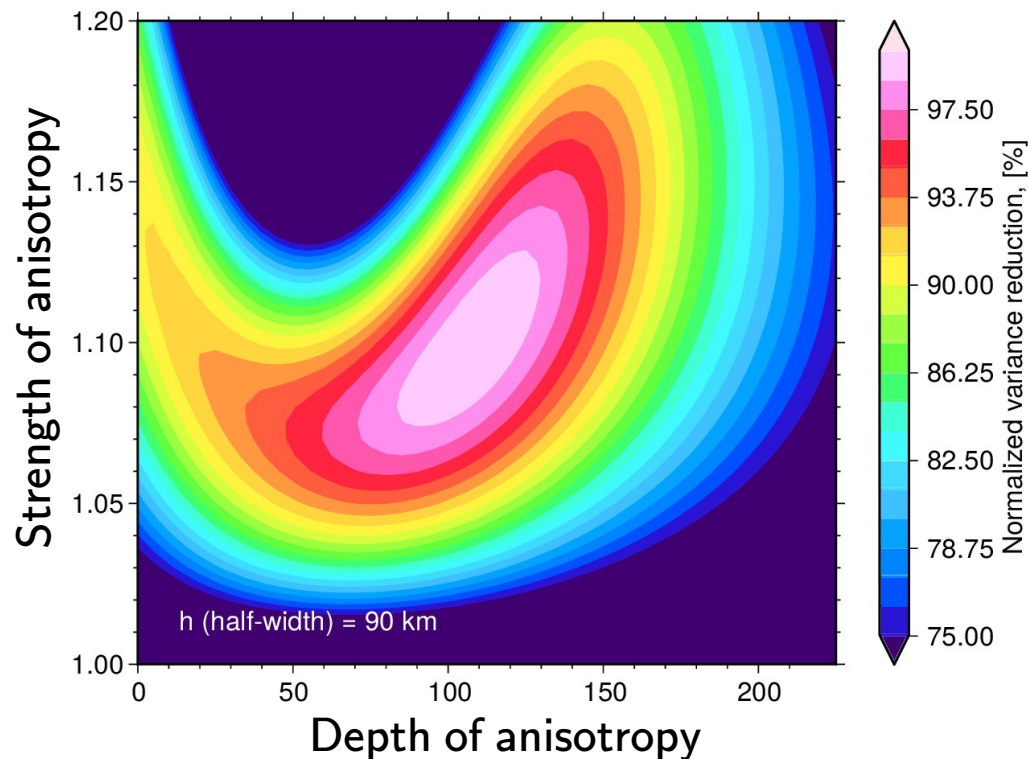
Fit (VR) for  
Pacific dataset: 
$$VR = 1 - \frac{|\mathbf{G} \cdot \mathbf{m} - \mathbf{d}_{obs}|}{|\mathbf{d}_{obs}|}$$





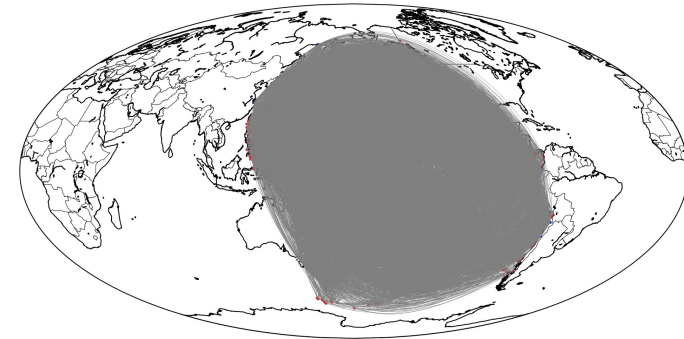
# (3) Hypothesis test: flat layer

Grid search over layer depth  $z_0$   
and the anisotropy strength  $\xi_{\max}$



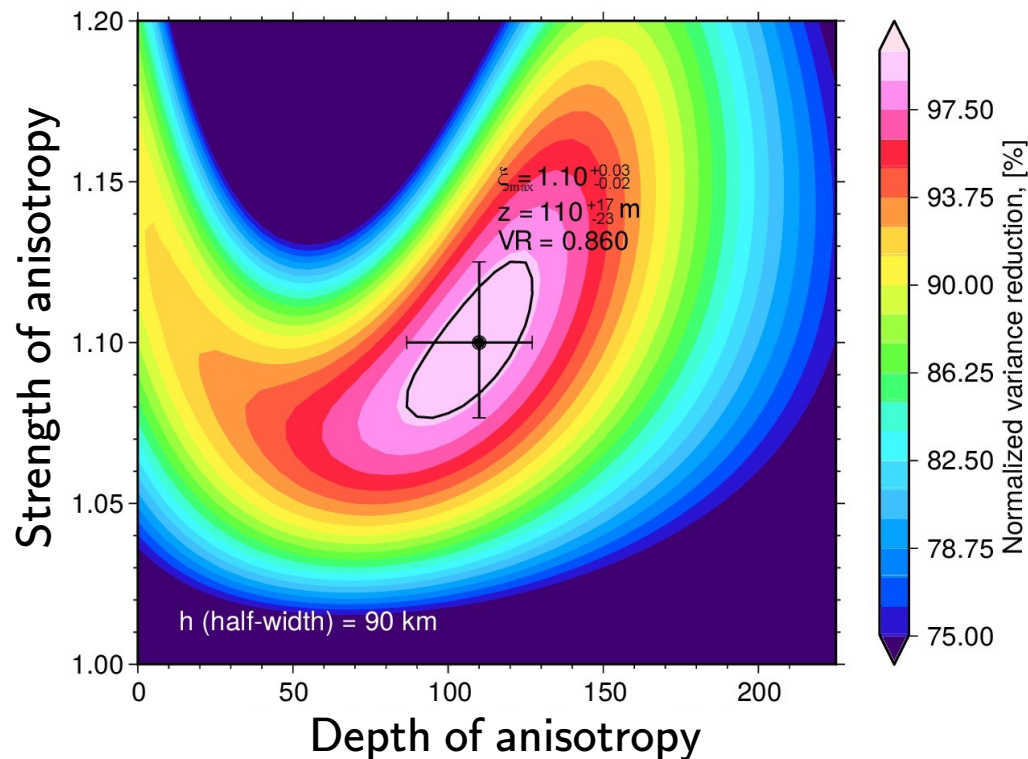
Fit (VR) for  
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$$VR = 1 - \frac{|\mathbf{G} \cdot \mathbf{m} - \mathbf{d}_{\text{obs}}|}{|\mathbf{d}_{\text{obs}}|}$$



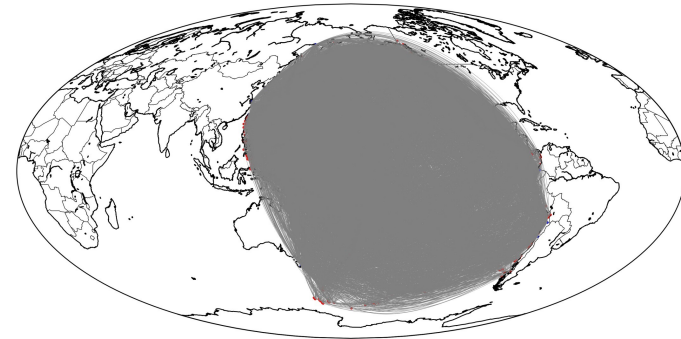
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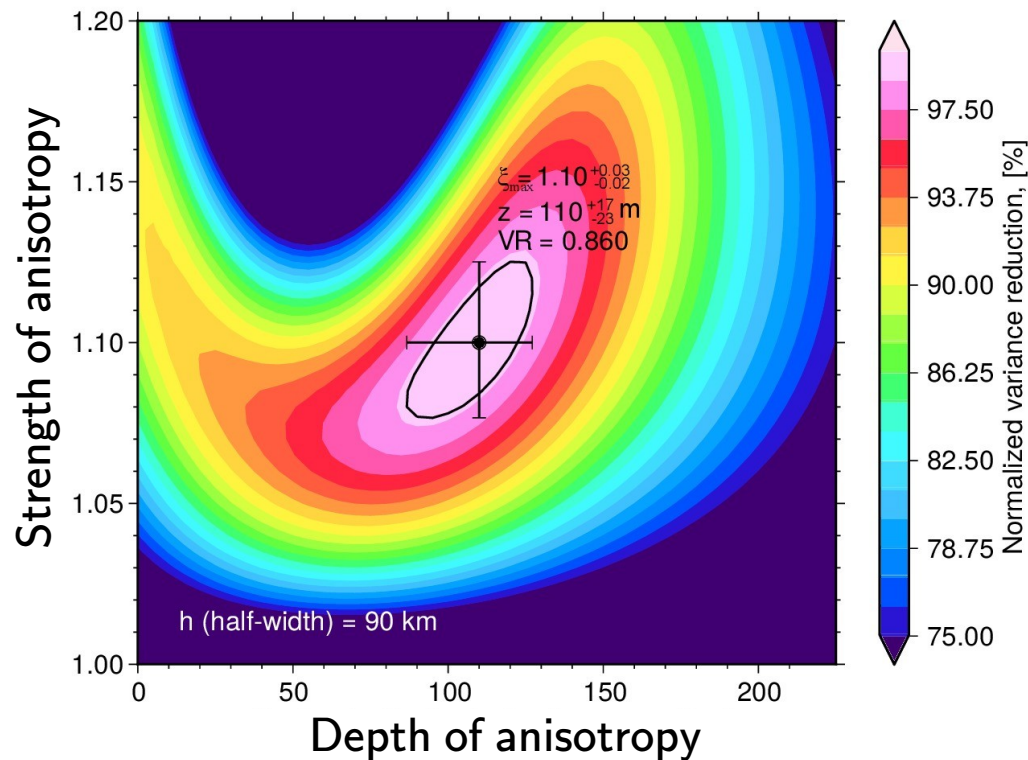
Fit (VR) for  
Pacific dataset:

$$VR = 1 - \frac{|\mathbf{G} \cdot \mathbf{m} - \mathbf{d}_{\text{obs}}|}{|\mathbf{d}_{\text{obs}}|}$$

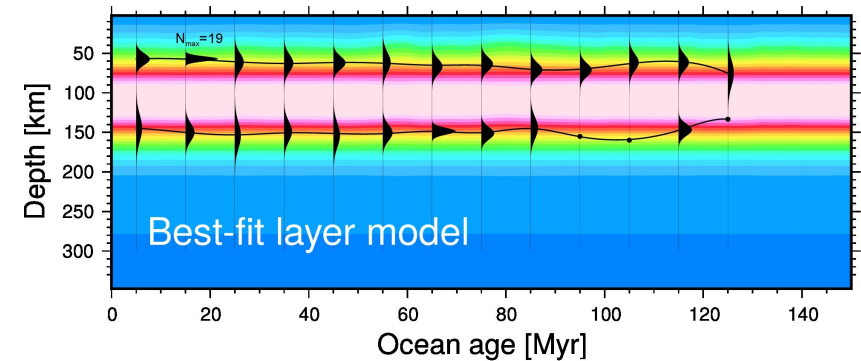
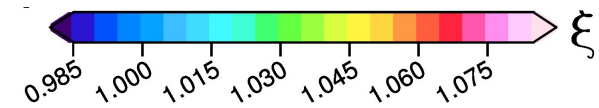
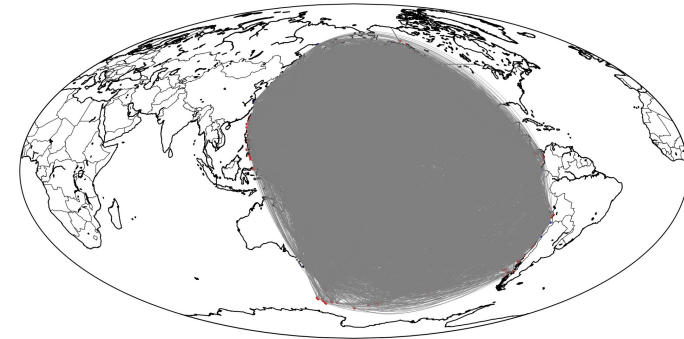


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Grid search over layer depth  $z_0$   
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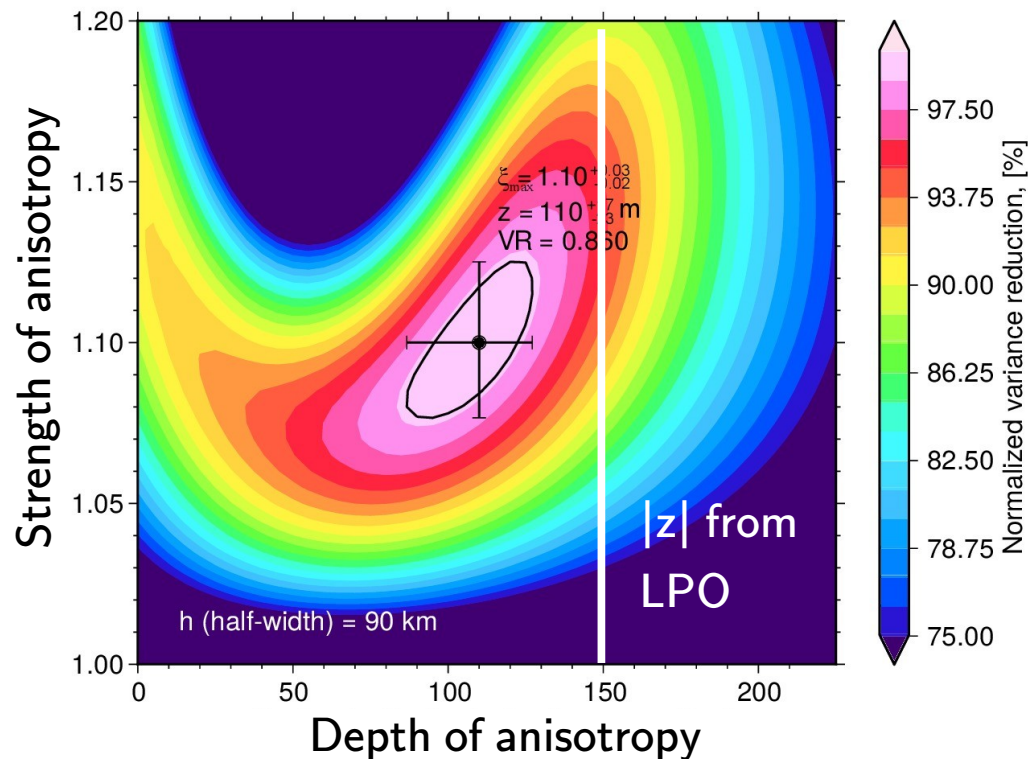


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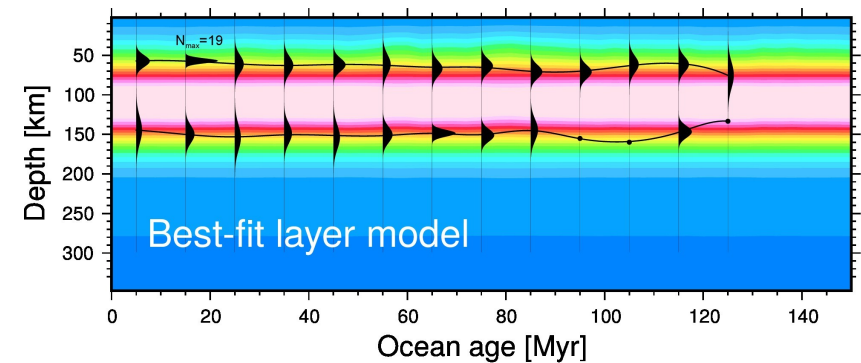
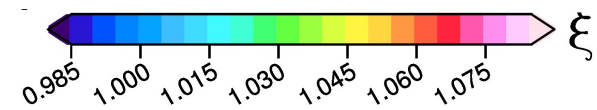
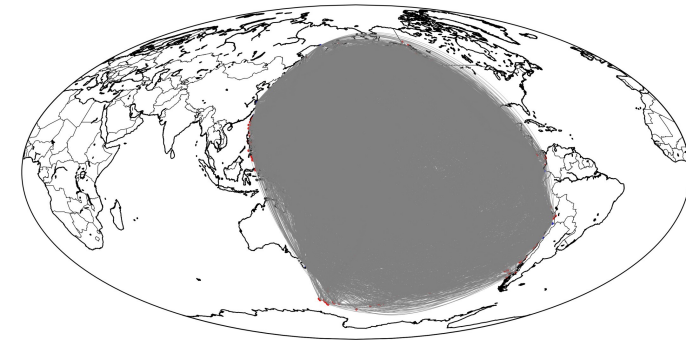
# (3) Hypothesis test: flat layer

Grid search over layer depth  $z_0$   
and the anisotropy strength  $\xi_{\max}$



**Again:** LPO anisotropy too deep

Fit (VR) for Pacific dataset:

$$VR = 1 - \frac{|\mathbf{G} \cdot \mathbf{m} - \mathbf{d}_{\text{obs}}|}{|\mathbf{d}_{\text{obs}}|}$$


### (3) Hypothesis test: $\sqrt{\tau}$ case

**Model:**

$$z_1 = a + c\sqrt{\kappa\hat{\tau}(\mathbf{x}, f)} \quad \hat{\tau}(\mathbf{x}, f) = \begin{cases} \tau(\mathbf{x}) & \text{for } \tau(\mathbf{x}) < f \\ f & \text{for } \tau(\mathbf{x}) \geq f \end{cases}$$

## (3) Hypothesis test: $\sqrt{\tau}$ case

**Model:**

$$z_1 = a + c\sqrt{\kappa\hat{\tau}(\mathbf{x}, f)} \quad \hat{\tau}(\mathbf{x}, f) = \begin{cases} \tau(\mathbf{x}) & \text{for } \tau(\mathbf{x}) < f \\ f & \text{for } \tau(\mathbf{x}) \geq f \end{cases}$$

Flattening age f

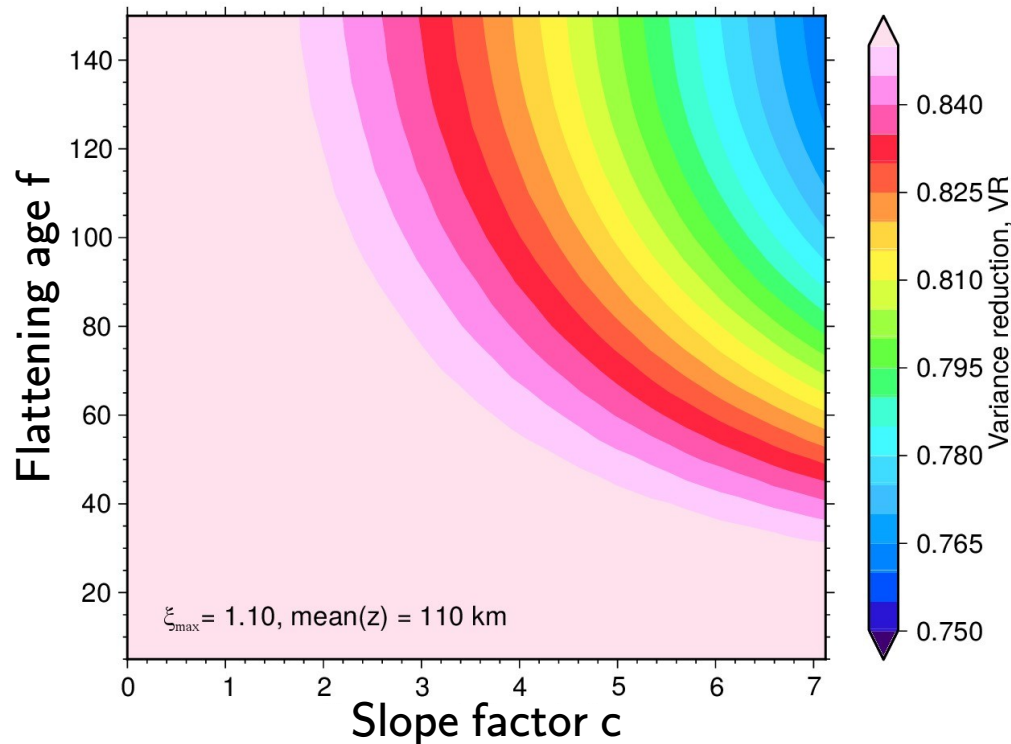
Slope factor c



# (3) Hypothesis test: $\sqrt{\tau}$ case

**Model:**

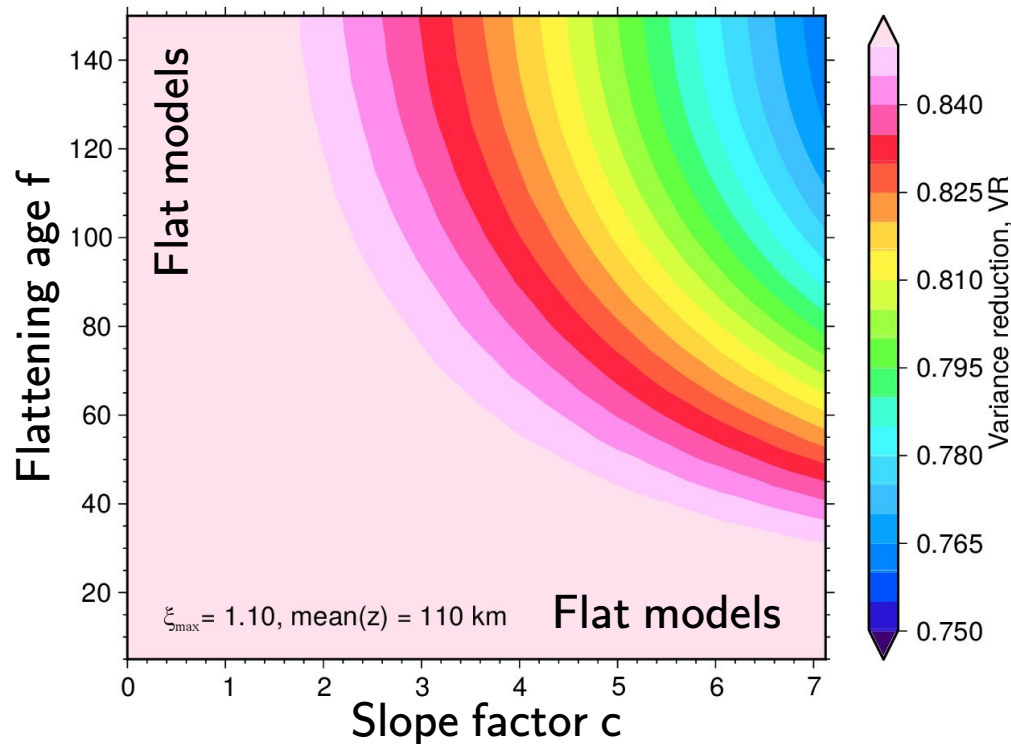
$$z_1 = a + c \sqrt{\kappa \hat{\tau}(\mathbf{x}, f)} \quad \hat{\tau}(\mathbf{x}, f) = \begin{cases} \tau(\mathbf{x}) & \text{for } \tau(\mathbf{x}) < f \\ f & \text{for } \tau(\mathbf{x}) \geq f \end{cases}$$



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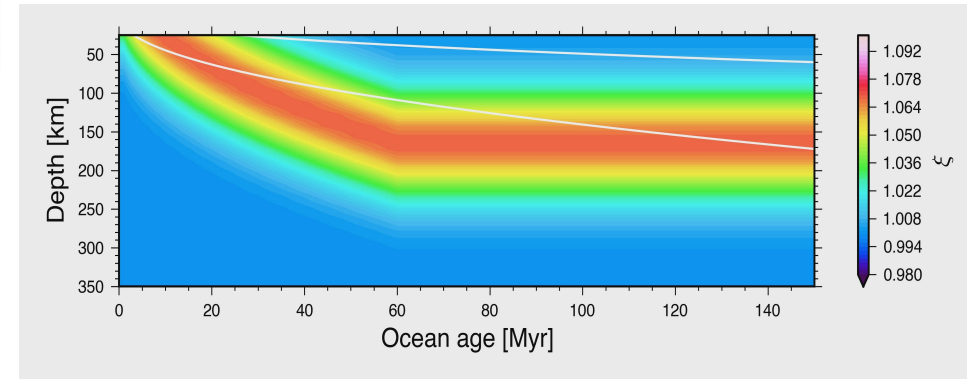
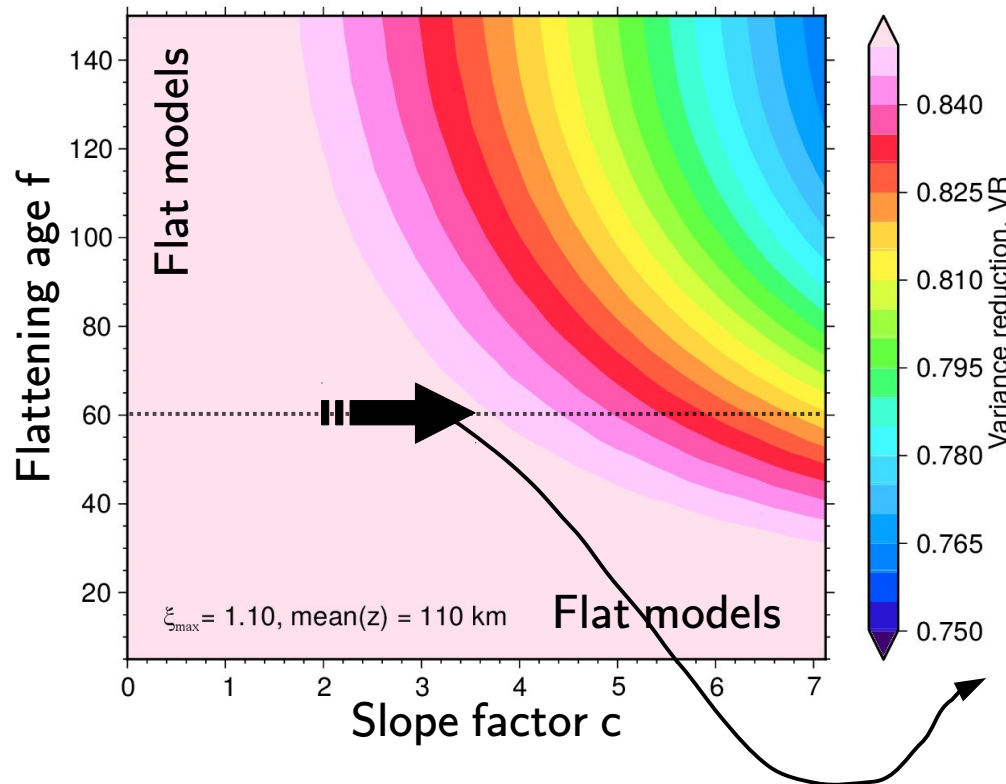




# (3) Hypothesis test: $\sqrt{\tau}$ case

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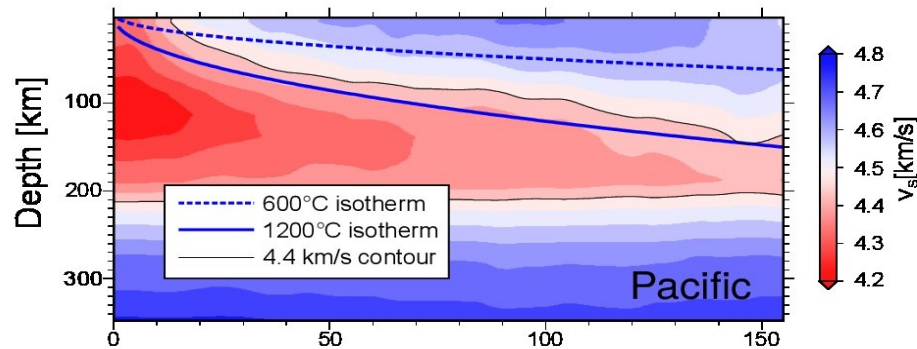


## (3) Hypothesis test: summary

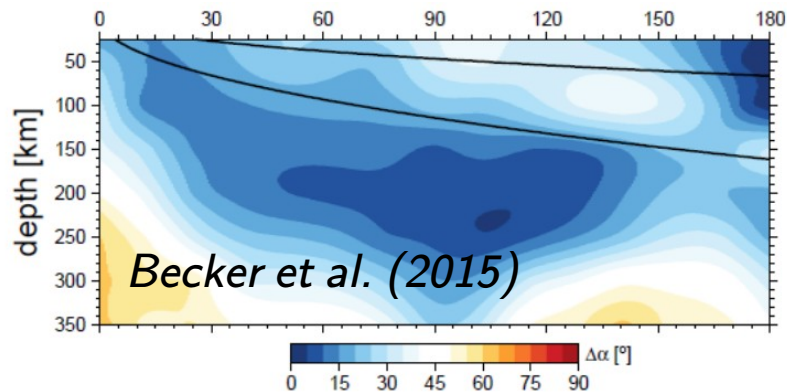
- Regionalized hypothesis tests confirm that depth of zone is shallower than LPO prediction.
- Some age-dependent models borderline compatible with data, but there is preference for flat models.

# (3) Summary of results

Isotropic structure broadly matches HSC

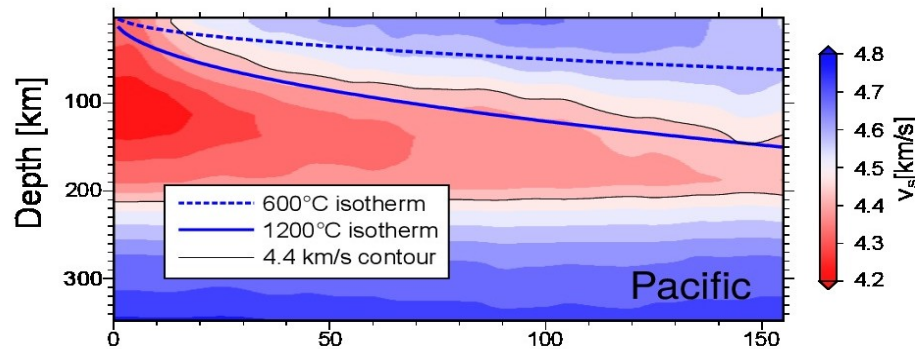


Azim. aniso. from LPO vs. sl2013sv match

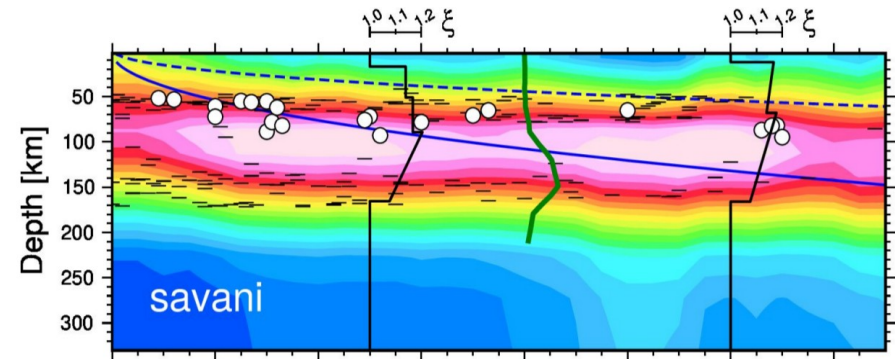


# (3) Summary of results

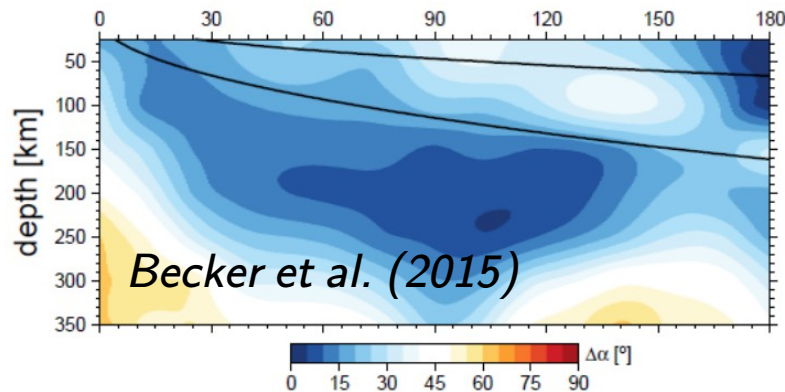
Isotropic structure broadly matches HSC



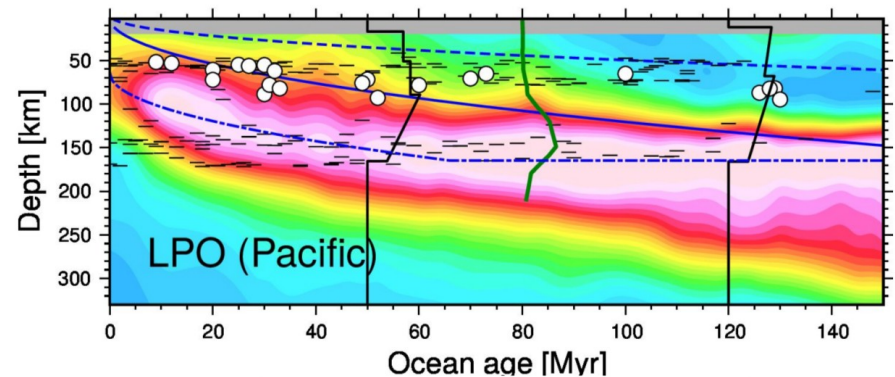
Tests indicate that flatness is **no artifact**



Azim. aniso. from LPO vs. sl2013sv **match**

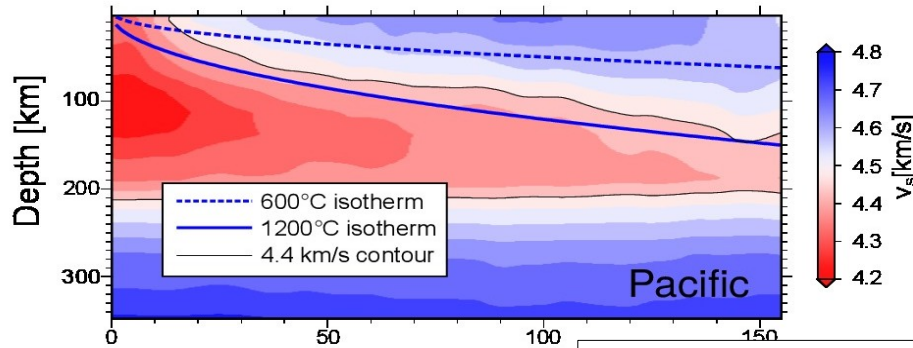


Pure LPO model doesn't work, **too deep**

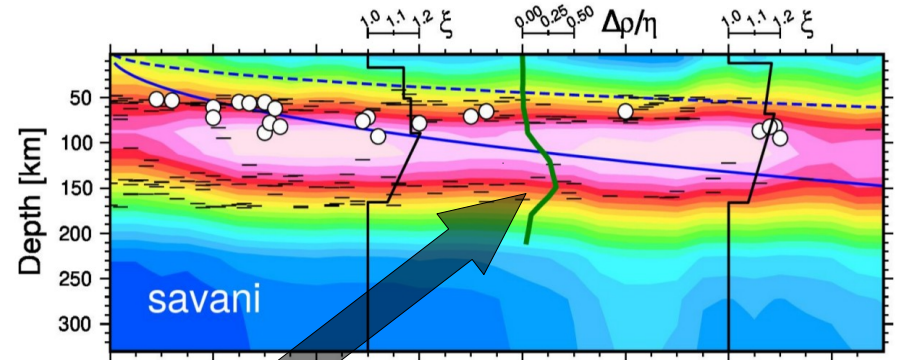


# (3) Summary of results

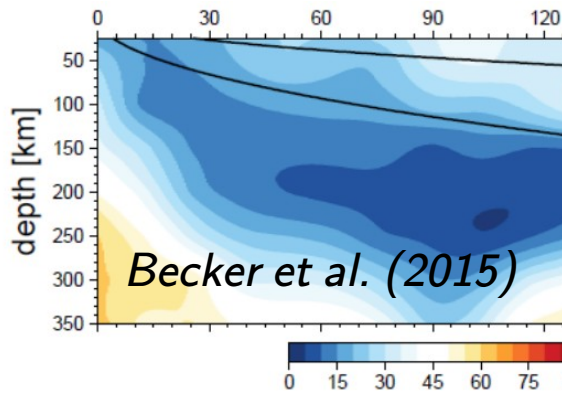
Isotropic structure broadly matches HSC



Tests indicate that flatness is no artifact

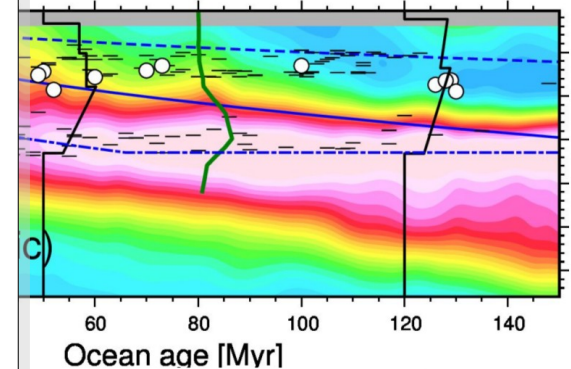


Azim. aniso. from LPO vs. sl20



Increased partial melt mobility  
(Sakamaki et al. 2013)

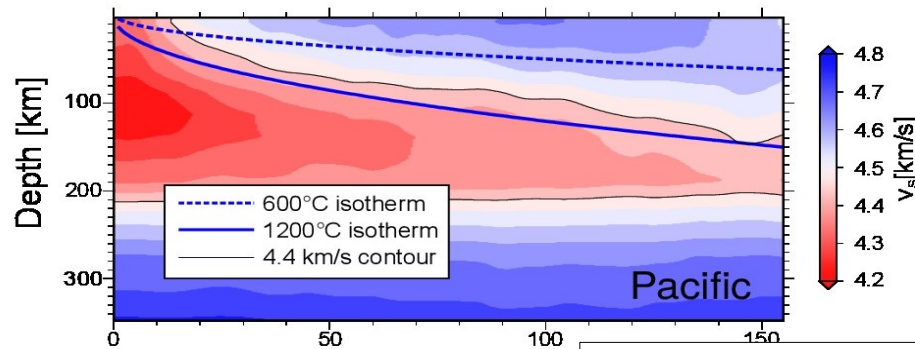
del doesn't work, too deep



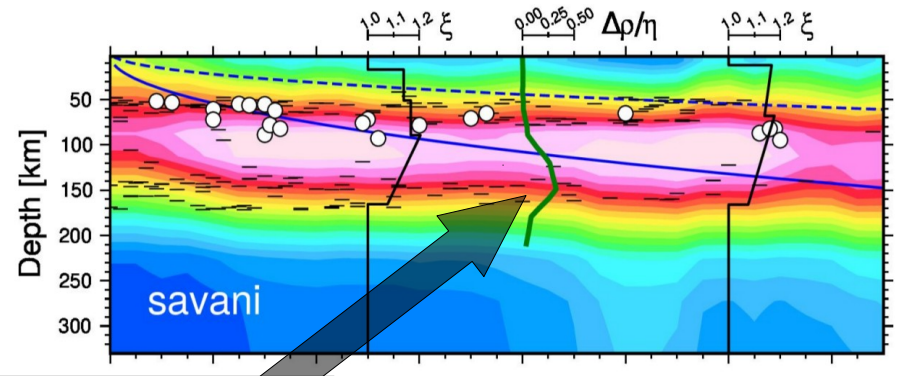


# (3) Summary of results

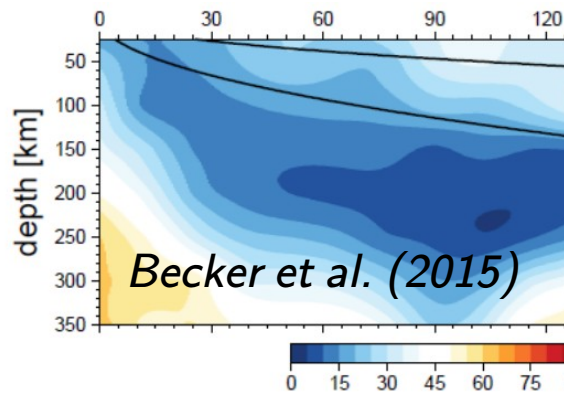
Isotropic structure broadly matches HSC



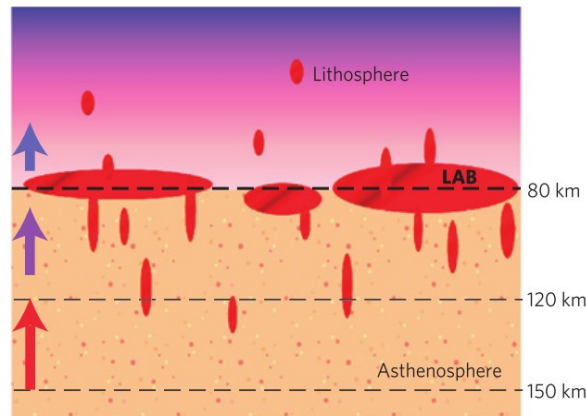
Tests indicate that flatness is **no artifact**



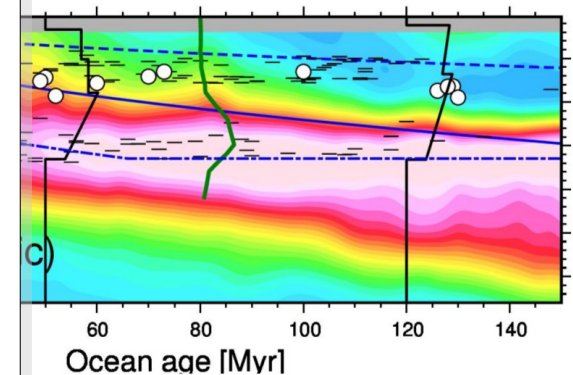
Azim. aniso. from LPO vs. sl20



Increased partial melt mobility  
(Sakamaki et al. 2013)



del doesn't work, **too deep**





# (3) Summary of results

## Unified conceptual model of the upper mantle

Spreading center

Hotspot

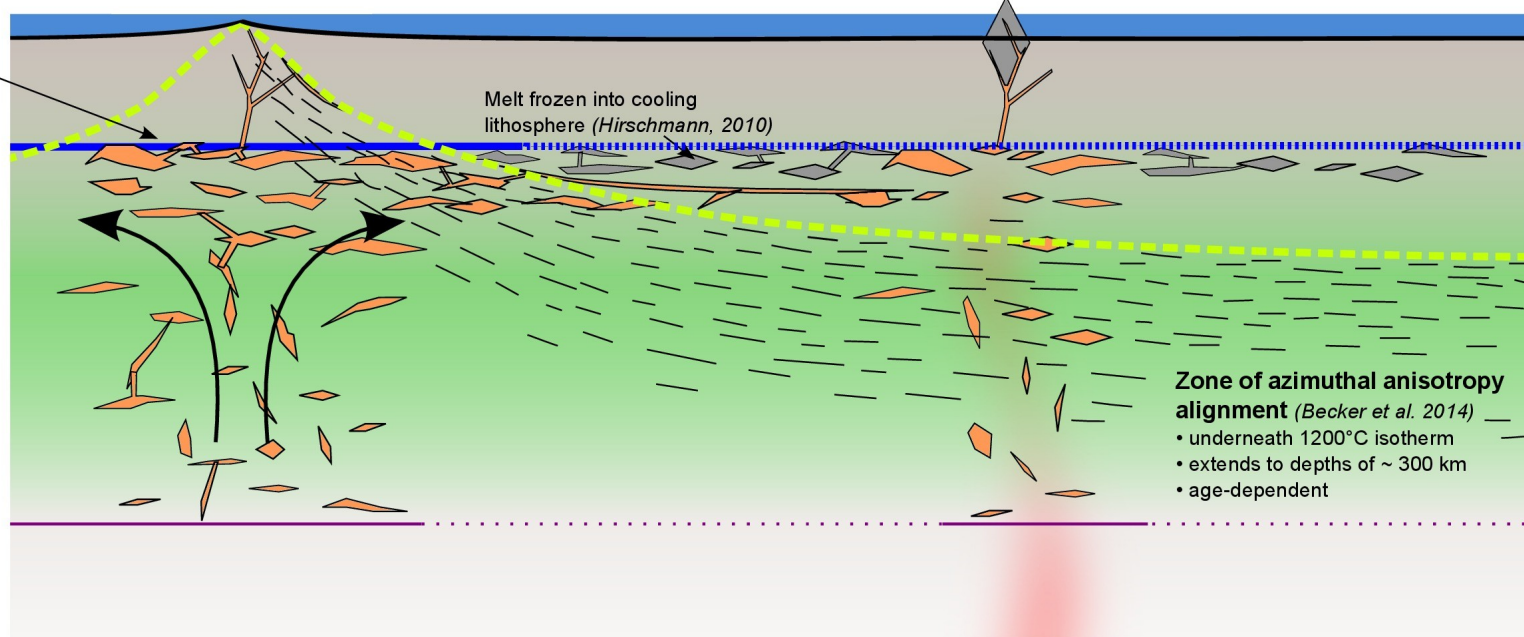
Auer et al. (2015)

High partial melt fraction due to reduced melt mobility (Sakamaki et al. 2013, Hirschmann 2010)

~ 60 km

High radial anisotropy (this study, Beghein et al. 2014, Burgos et al. 2014), age-independent, ~ 60 - 150 km, high melt mobility between 120 and 150 km (Sakamaki et al. 2013)

~ 150 km



**----- Mechanical base of lithosphere**  
 (e.g. Becker et al. 2014, Beghein et al. 2014, Debayle and Ricard 2013, Burgos et al. 2014)  
 •  $v_s$  and azimuthal anisotropy, ~ 1200°C isotherm  
 • age-dependent, ~ 0 - 200 km

**----- Oceanic L Discontinuity**  
 (e.g. Gaherty et al. 1996, 1999, Schmerr et al. 2012)  
 • SS precursors, mainly at MOR  
 • age-independent ~ 150 km

**----- Oceanic G discontinuity**  
 (e.g. Kawakatsu et al. 2009, Schmerr 2012)  
 • RF, SS precursors and  $\xi$   
 • age-independent, ~ 60 km

## (3) Summary of results

### Unified conceptual model of the upper mantle

- G is probably not the LAB, but rather a mid-lithospheric discontinuity, as seen under continents
- Radial anisotropy sees frozen-in (100Ma-1Ga) structure (petrofabrics, melt lamellae) + contributions from asthenospheric LPO
- Azimuthal anisotropy dominated LPO from recent (0-100 Ma) mantle flow, marks the “mechanical LAB”

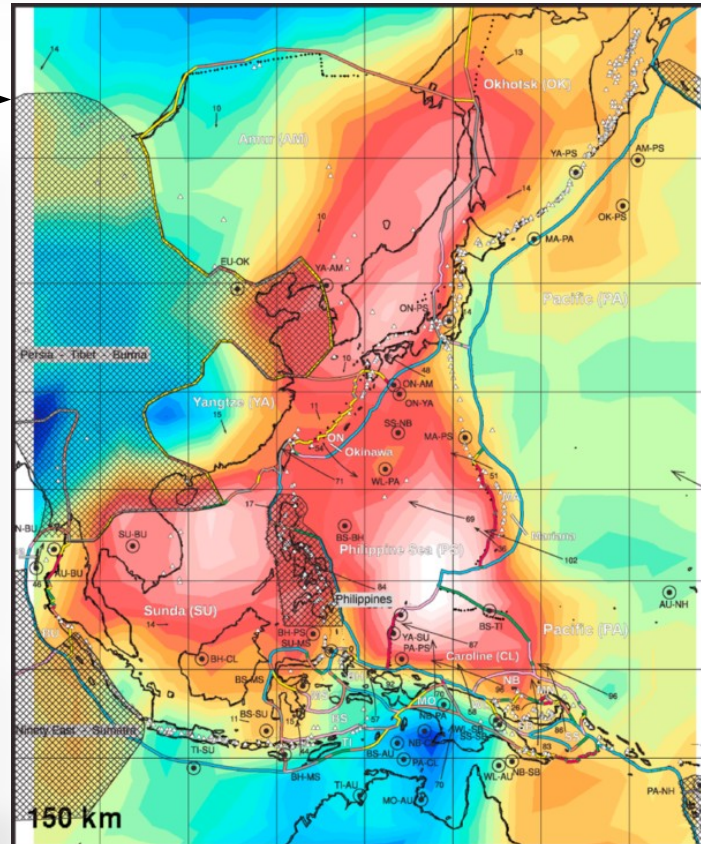
# (4) Compositional tomography

## Summary of results

$$\frac{(v_p - v_{p\_exp})}{v_{p\_exp}}$$

Scaling:

$$\gamma \frac{\delta \ln V_{SV}}{\delta \ln V_{PV}} = \gamma \frac{\delta \ln V_{SH}}{\delta \ln V_{PH}} = \gamma R_{S/P}$$



### Key points

- *SPani*: joint model of  $v_p$  and  $v_s$  variations in the Earth's mantle
- Anomalous  $v_p/v_s$  ratios at marginal basins
- Might be related to hydration

# (5) Hybrid waveform tomography

Motivation: Using regional waves?

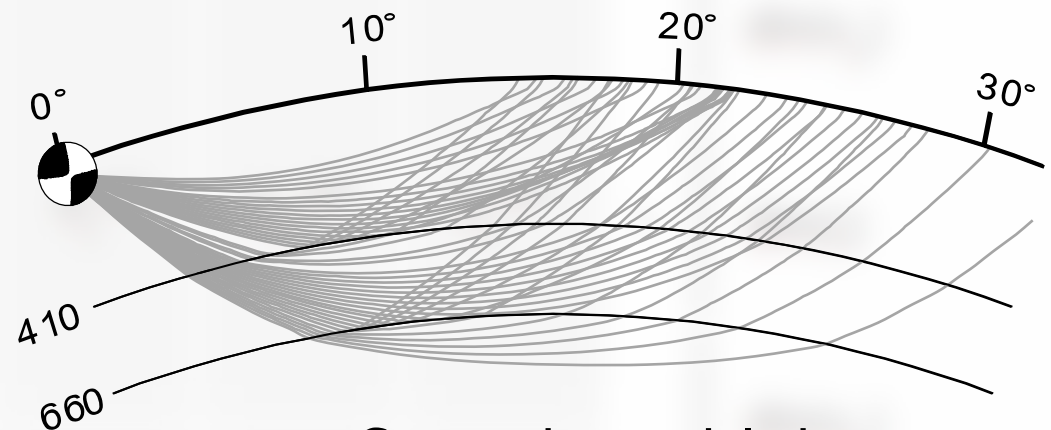
- Resolution of tele-seismic P-waves is limited, especially in the transition zone
- Can we use new datatypes such as regional body waves to overcome this problem?

# (5) Hybrid waveform tomography

Motivation: Using regional waves?

- Resolution of tele-seismic P-waves is limited, especially in the transition zone
- Can we use new datatypes such as regional body waves to overcome this problem?

... triplicated P-waves?



- Cannot be modeled with ray theory
- $\Delta = 14 - 32^\circ$

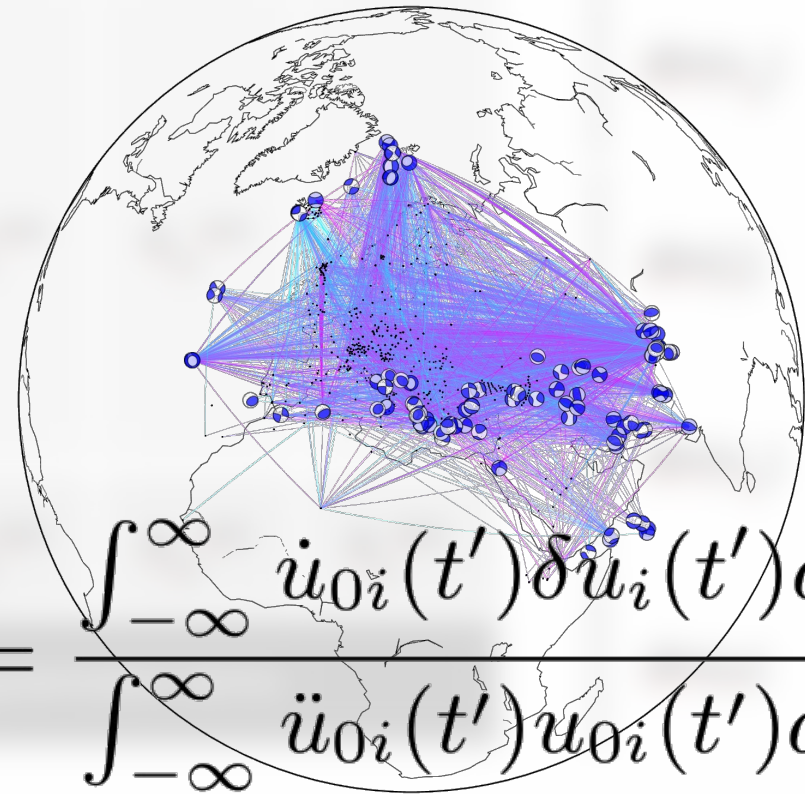


# (5) Hybrid waveform tomography

Motivation: Using regional waves?

- Resolution of tele-seismic P-waves is limited, especially in the transition zone
- Can we use new datatypes such as regional body waves to overcome this problem?

Yes! Cf. *Stähler et al. (2012)*

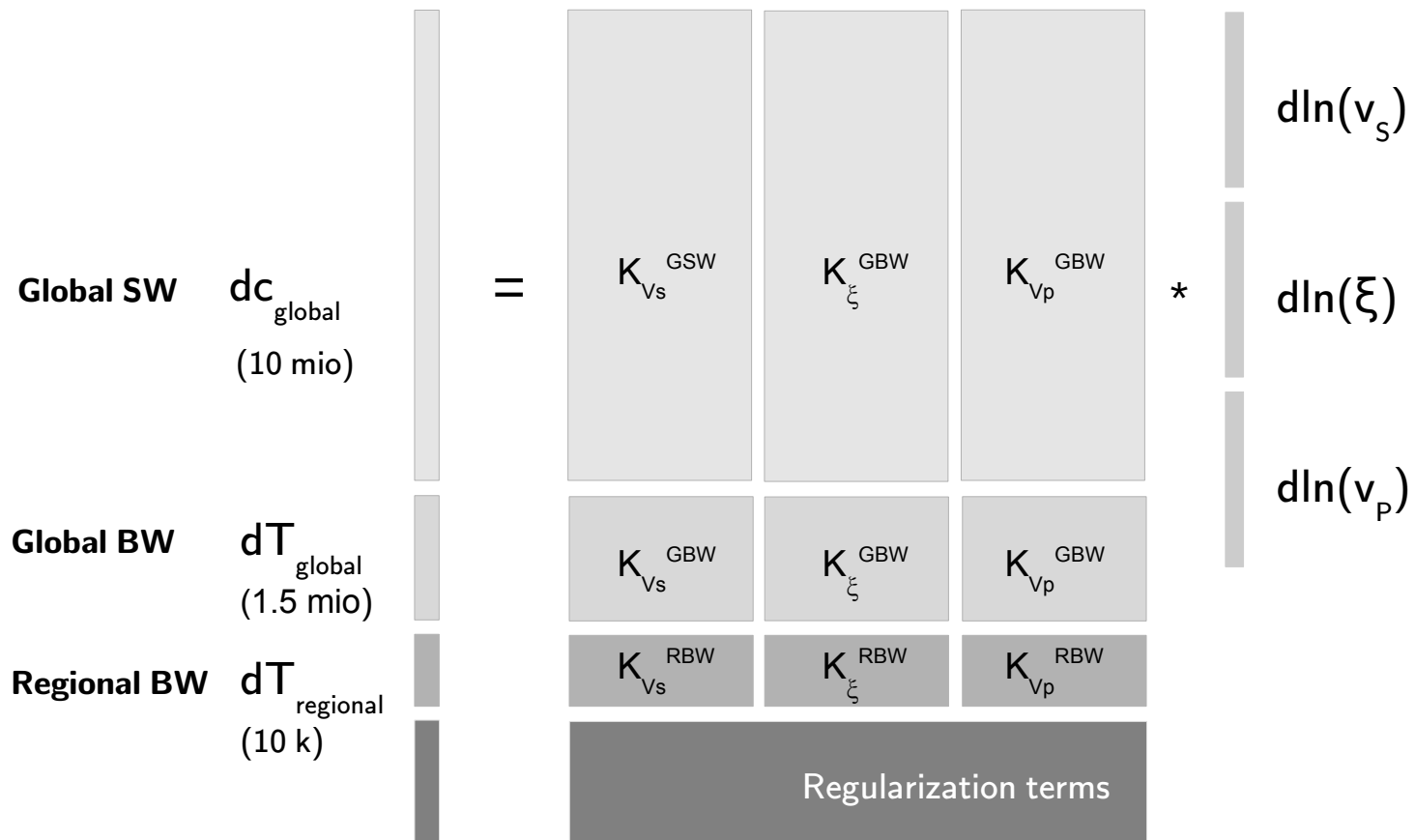


$$\delta T_r = \frac{\int_{-\infty}^{\infty} \dot{u}_{oi}(t') \delta \dot{u}_i(t') dt'}{\int_{-\infty}^{\infty} \ddot{u}_{oi}(t') u_{oi}(t') dt'}$$



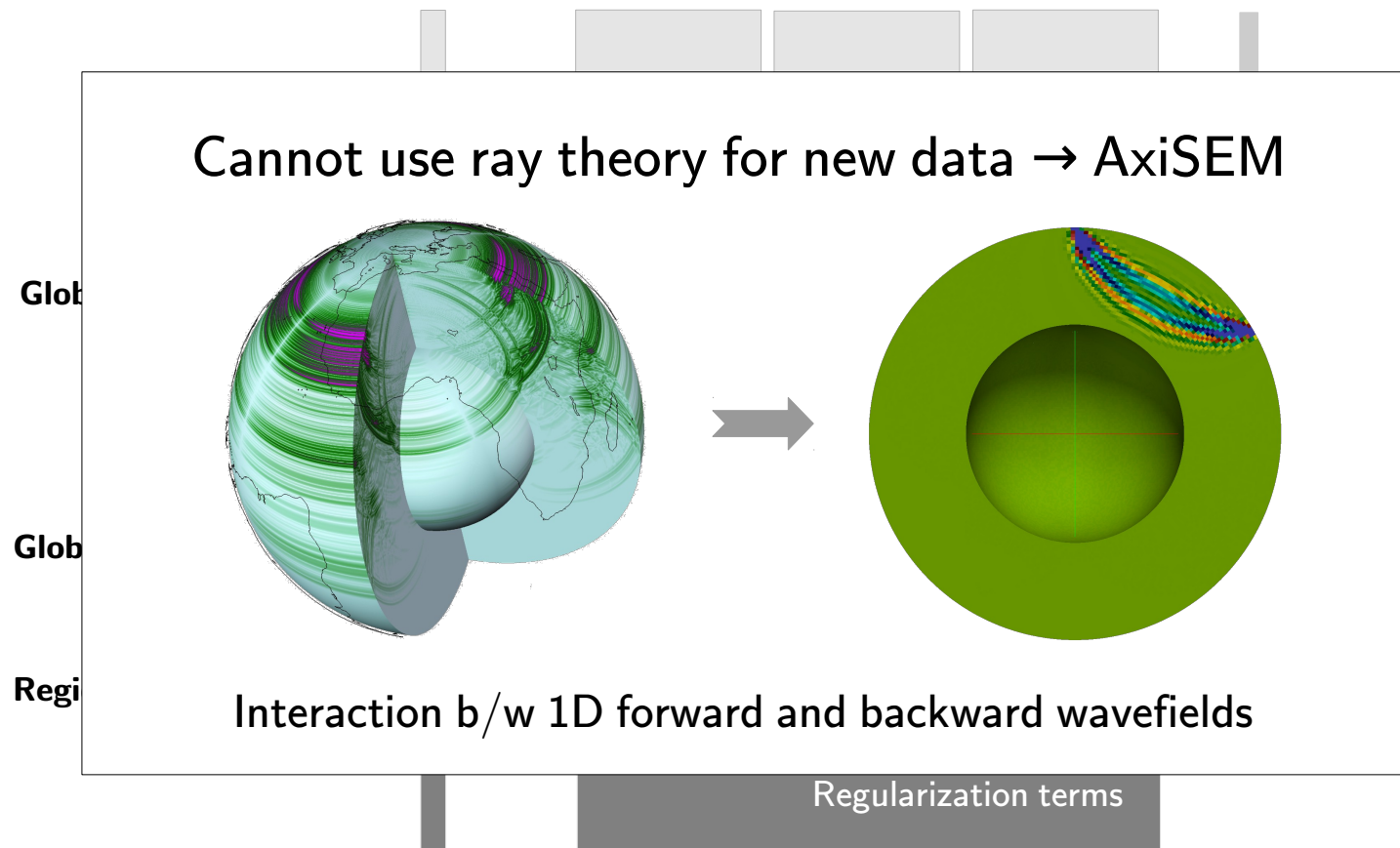
# (5) Hybrid waveform tomography

## Summary of the inverse problem



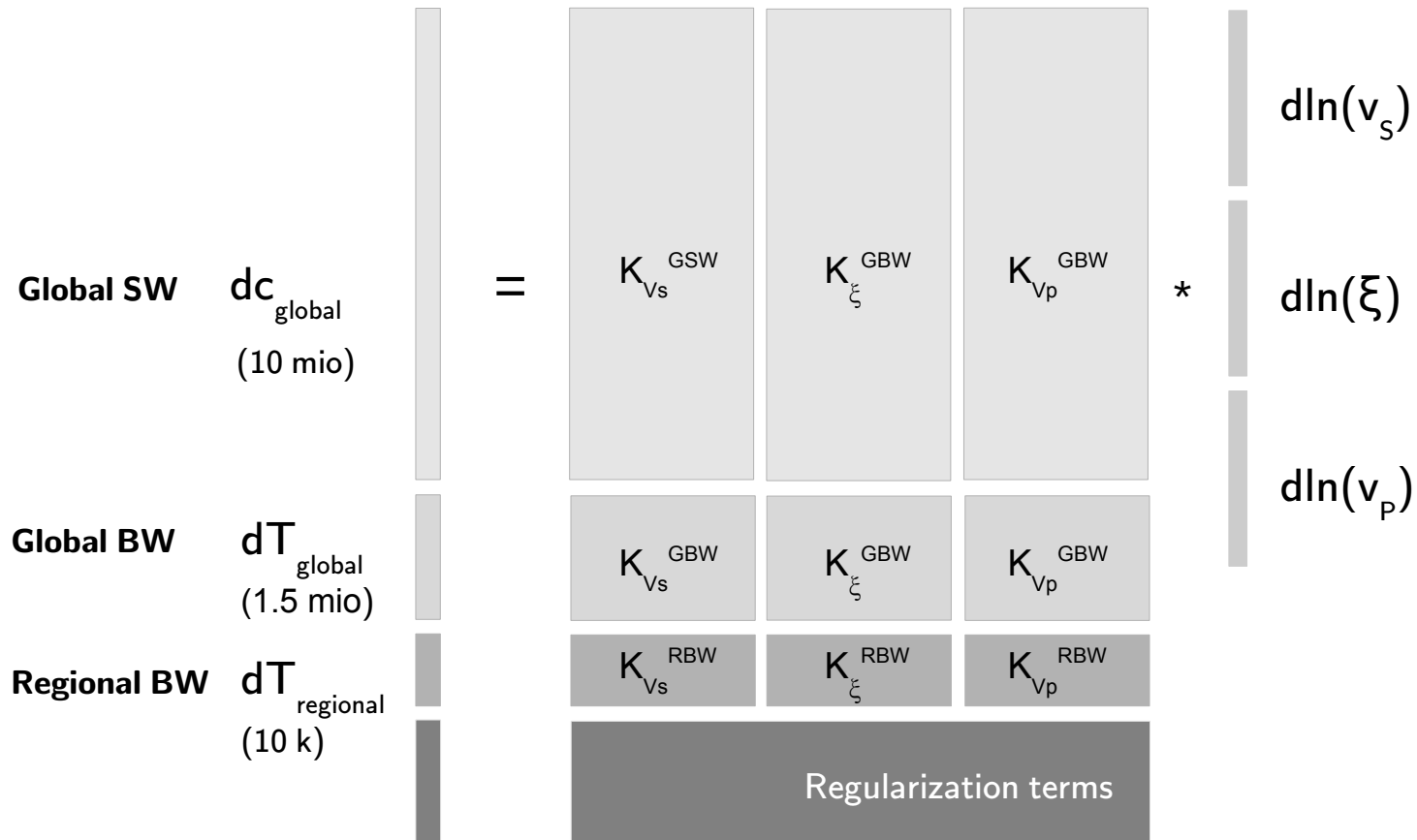
# (5) Hybrid waveform tomography

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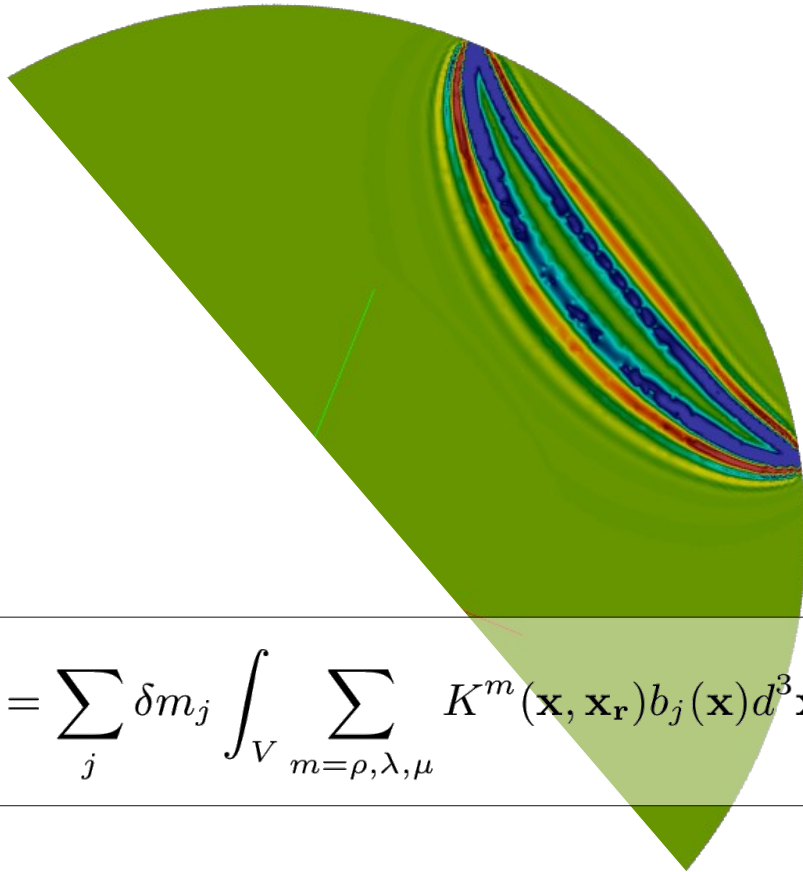
# (5) Hybrid waveform tomography

## Summary of the inverse problem



# (5) Hybrid waveform tomography

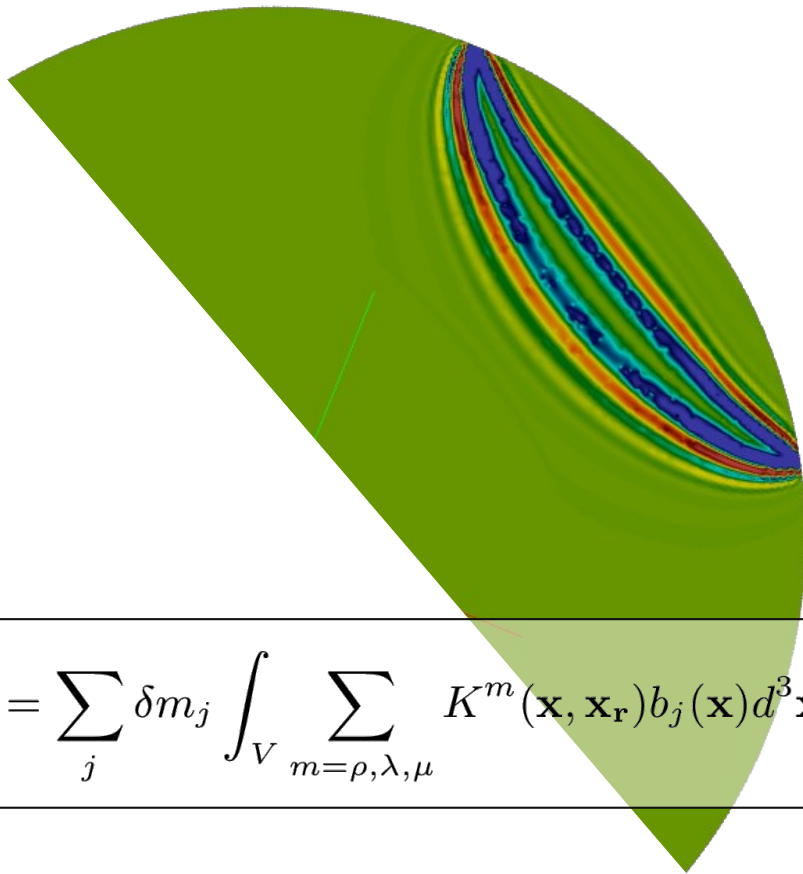
MCkernel: Kernel quadrature



$$\delta T_r = \sum_j \delta m_j \int_V \sum_{m=\rho,\lambda,\mu} K^m(\mathbf{x}, \mathbf{x}_r) b_j(\mathbf{x}) d^3 \mathbf{x}$$

# (5) Hybrid waveform tomography

MCkernel: Kernel quadrature

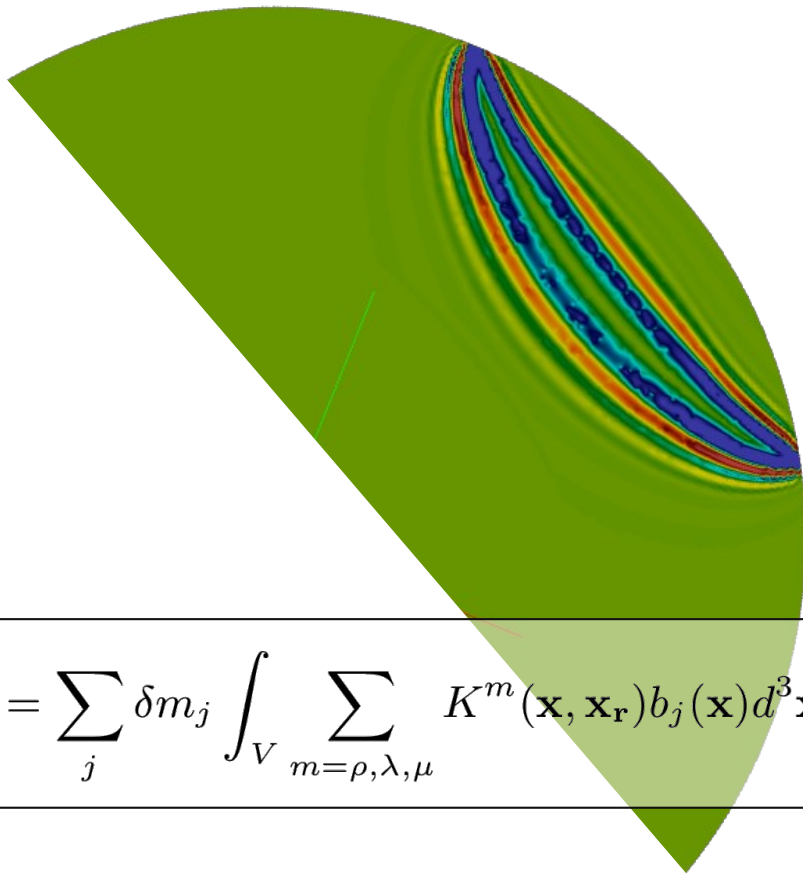


- Parallel Monte-Carlo integration approach

$$\delta T_r = \sum_j \delta m_j \int_V \sum_{m=\rho,\lambda,\mu} K^m(\mathbf{x}, \mathbf{x}_r) b_j(\mathbf{x}) d^3 \mathbf{x}$$

# (5) Hybrid waveform tomography

MCkernel: Kernel quadrature



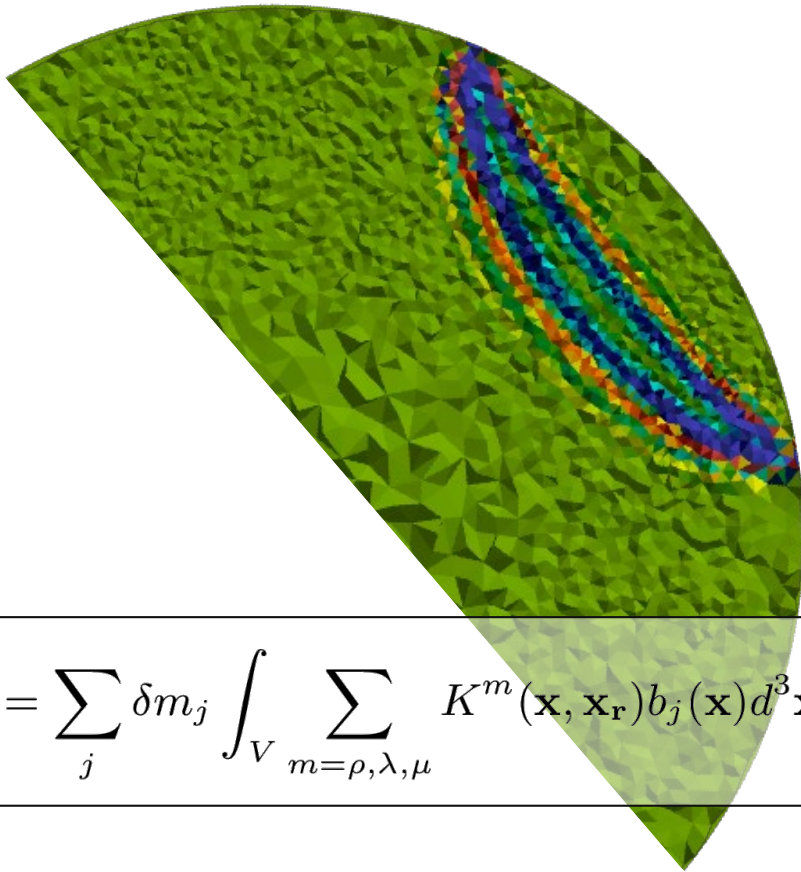
- Parallel Monte-Carlo integration approach
- Supports tetrahedral and voxel meshes

$$\delta T_r = \sum_j \delta m_j \int_V \sum_{m=\rho,\lambda,\mu} K^m(\mathbf{x}, \mathbf{x}_r) b_j(\mathbf{x}) d^3 \mathbf{x}$$



# (5) Hybrid waveform tomography

MCkernel: Kernel quadrature

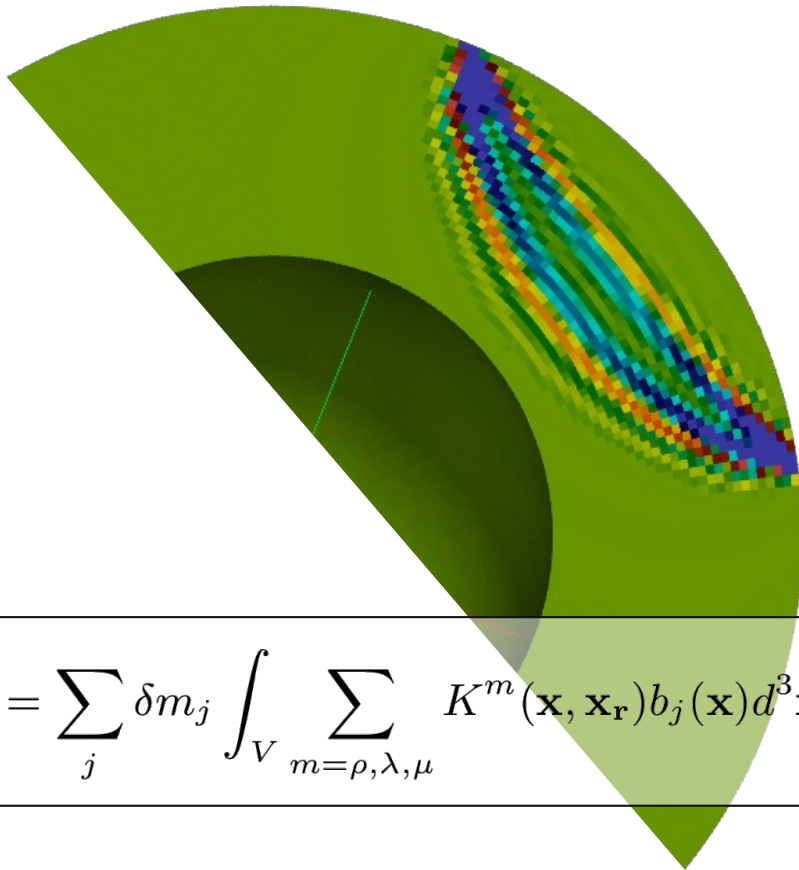


- Parallel Monte-Carlo integration approach
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$$\delta T_r = \sum_j \delta m_j \int_V \sum_{m=\rho,\lambda,\mu} K^m(\mathbf{x}, \mathbf{x}_r) b_j(\mathbf{x}) d^3 \mathbf{x}$$

# (5) Hybrid waveform tomography

MCkernel: Kernel quadrature

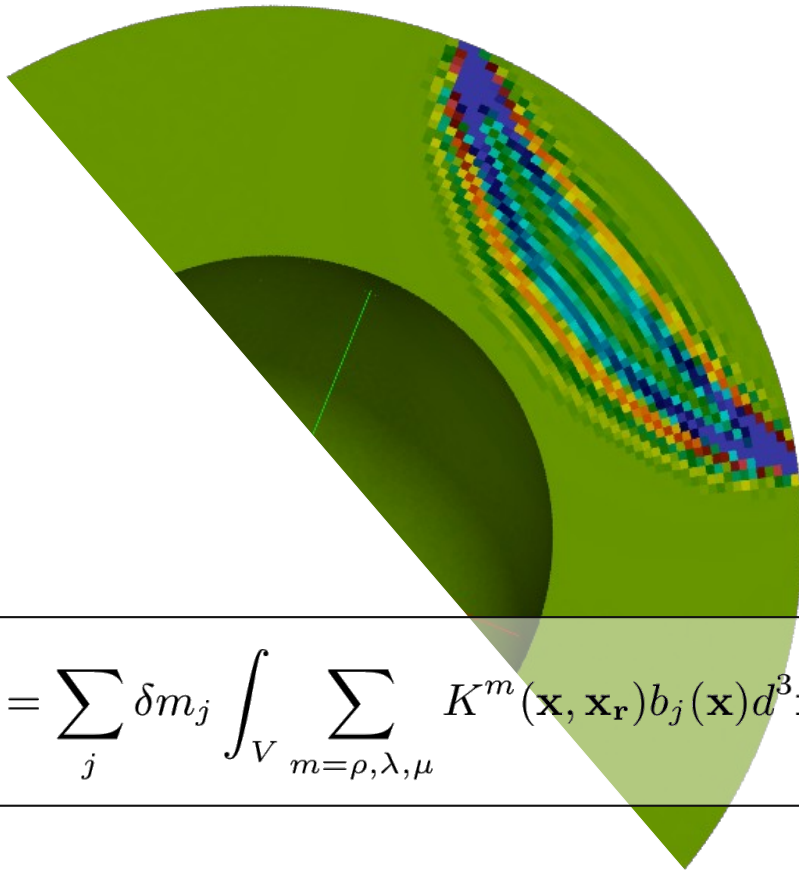


- Parallel Monte-Carlo integration approach
- Supports tetrahedral and voxel meshes

$$\delta T_r = \sum_j \delta m_j \int_V \sum_{m=\rho,\lambda,\mu} K^m(\mathbf{x}, \mathbf{x}_r) b_j(\mathbf{x}) d^3 \mathbf{x}$$

# (5) Hybrid waveform tomography

MCkernel: Kernel quadrature

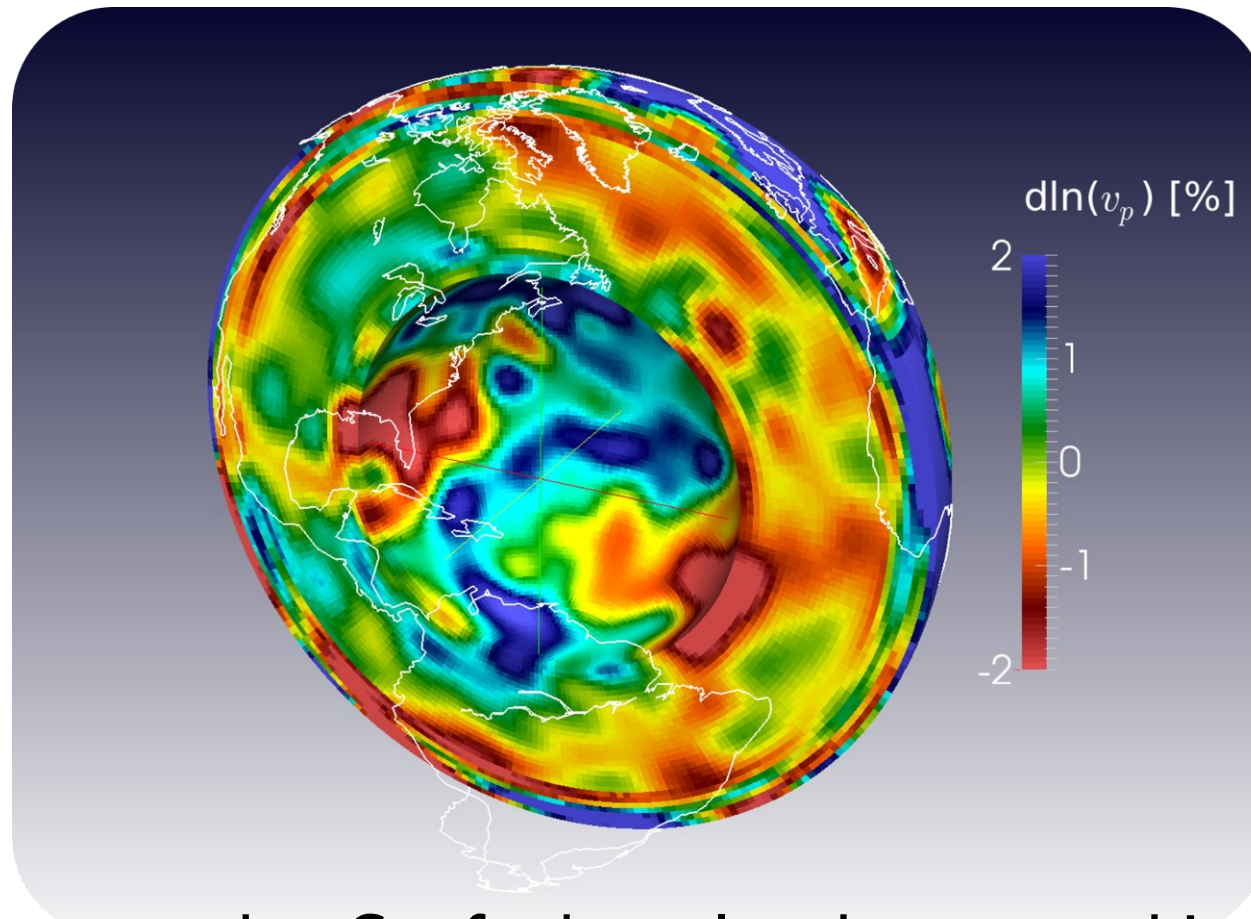


- Parallel Monte-Carlo integration approach
- Supports tetrahedral and voxel meshes
- Available for download @ <https://github.com/sstaehler/mckernel>

$$\delta T_r = \sum_j \delta m_j \int_V \sum_{m=\rho,\lambda,\mu} K^m(\mathbf{x}, \mathbf{x}_r) b_j(\mathbf{x}) d^3 \mathbf{x}$$

# (5) Hybrid waveform tomography

Prediction quality of kernels



Compute  $d = Gm$  for kernel and tomographic model

# (5) Hybrid waveform tomography

Prediction quality of kernels



# (5) Hybrid waveform tomography

Prediction quality of kernels





# **(5) Hybrid waveform tomography**

## Prediction quality of kernels

- Compute SPECFEM synthetics for tomographic model

# **(5) Hybrid waveform tomography**

## Prediction quality of kernels

- Compute SPECFEM synthetics for tomographic model
- Measure cross-correlation traveltimes for synthetics

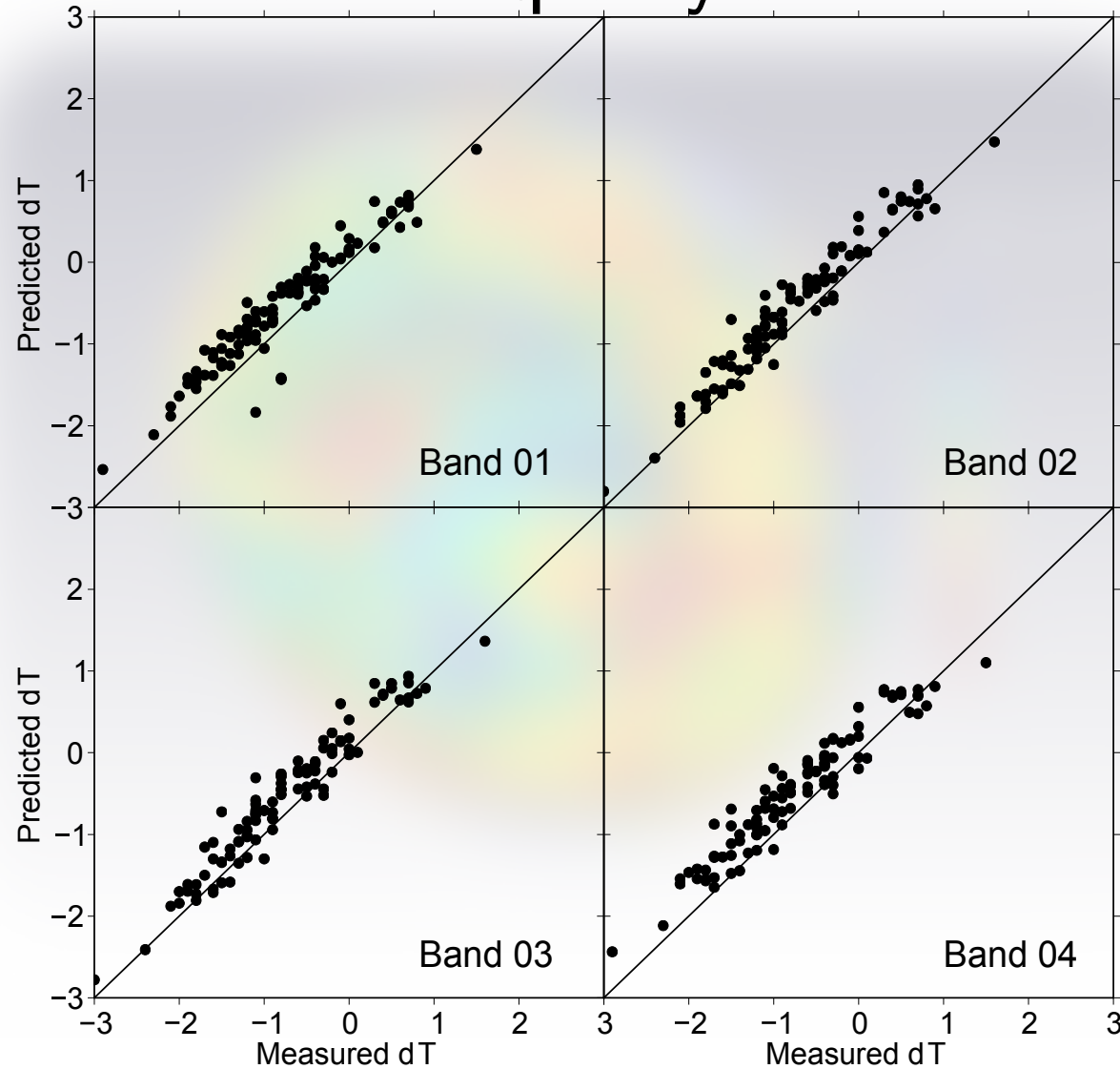
# (5) Hybrid waveform tomography

## Prediction quality of kernels

- Compute SPECFEM synthetics for tomographic model
- Measure cross-correlation traveltimes for synthetics
- Compare predicted and measured traveltimes

# (5) Hybrid waveform tomography

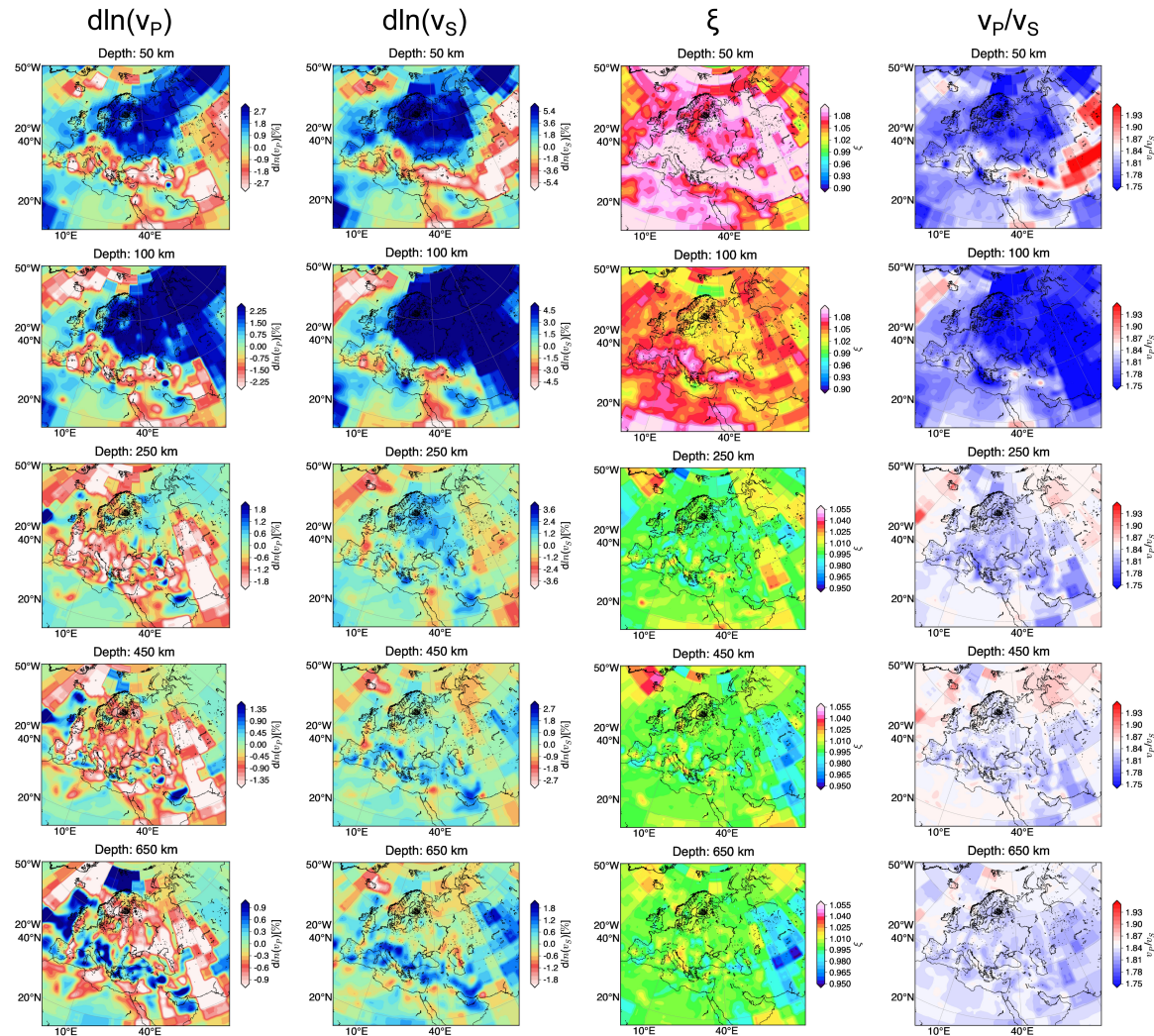
## Prediction quality of kernels



# (5) Hybrid waveform tomography

## Preliminary model Spani+EU (WIP)

With regional data



# Conclusions



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- *Savani*, a model of global whole-mantle radial anisotropy available for download

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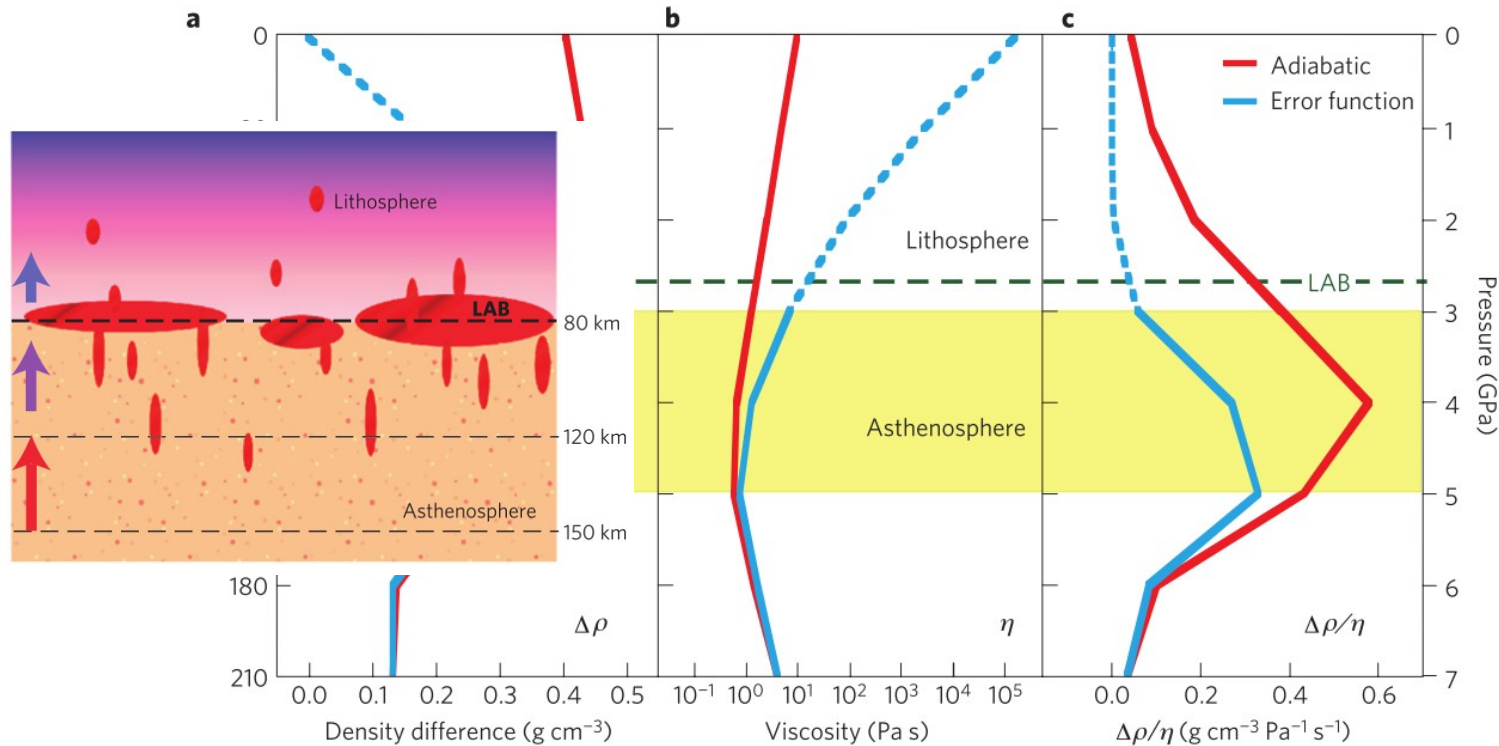
# Thank you!

(Apero at 19:00 in J Floor)

**Backup slides**



# Melt mobility



**Figure 2 | Magma properties and mobility as a function of depth.** **a**, Density difference between basaltic magma and olivine ( $\Delta\rho$ ) as a function of depth (and pressure). The red and blue lines are based on an adiabatic temperature gradient, and a realistic error-function temperature profile for mature oceanic lithosphere, respectively. The potential temperature is 1,623 K. Yellow shading highlights the depth range with anomalous physical properties of basaltic melts due to the structural transition in Al coordination. **b**, Viscosity ( $\eta$ ) of basaltic magma along the same two profiles. Hypothetical subsolidus density and viscosity profiles are dashed. **c**, Mobility  $\Delta\rho/\eta$  of basaltic magma along the same two profiles, respectively.

# LPO (1)

- LPO = Lattice preferred orientation
- Mainly due to dislocation glide (slip) along certain preferred glide planes and associated internal rotation/alignment of mineral grains
- To a first order, seismic anisotropy closely linked to the direction of shear or kinematic deformation

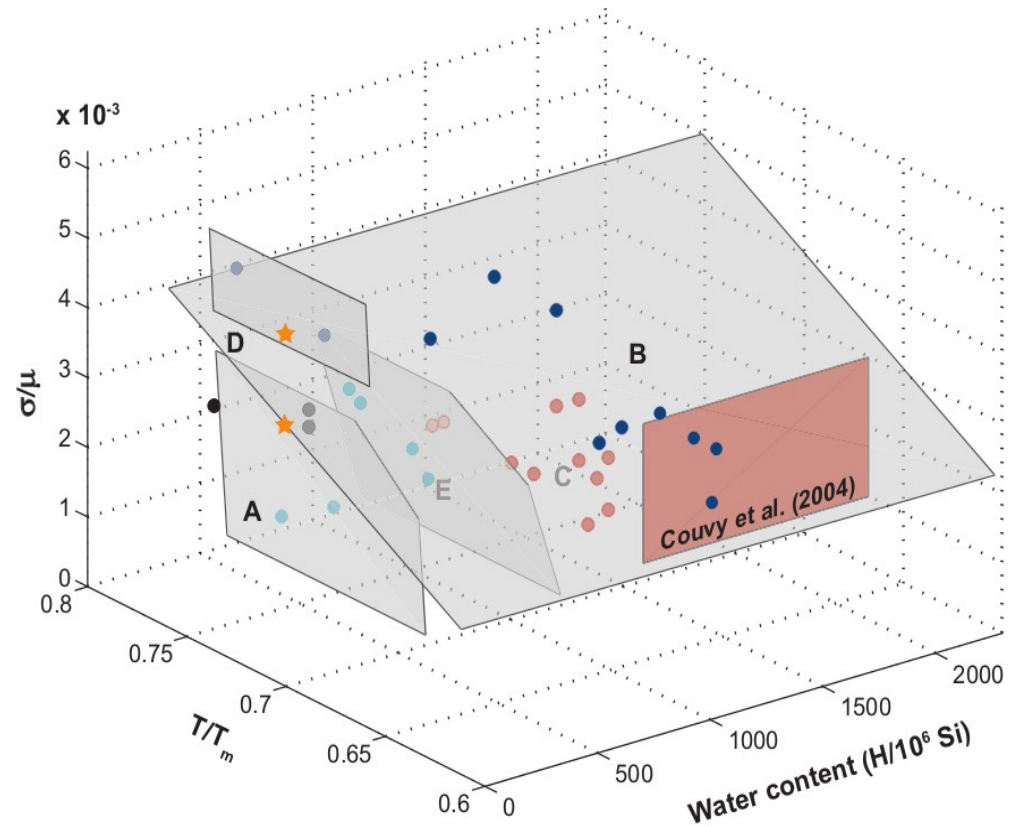
# LPO (2)

- Olivine shows a strong single-crystal elastic anisotropy. Fastest P- and S-wave velocities parallel to [100] crystallographic axis
- 70ties: Experimental studies on olivine indicated [100] axes of crystals to be nearly || to flow direction and (010) plane || to flow plane.
- Analysis of naturally deformed peridotite → [100](010) crystallographic axis of grains often coincide with rock lineation
- MCM + VPSC modelling + simple model of LPO development. **Works well** in case of large length scales and simple dynamics (e.g. Tommasi 1998, Becker 2006)

$\text{Xi} > 1 = \text{horizontal flow}$        $\text{Xi} < 1 = \text{vertical flow}$

# LPO (3)

- New findings in past ~10 years suggest that it's more difficult
- LPO type is sensitive to water, stress, temperature, pressure, melting (→ Carmen's work)
- Way of interpretation of seismic anisotropy needs major modifications!



A-type: Lithosphere, low water content

E-type: Asthenosphere, higher water content

# LPO (4)

Shear wave splitting (direction of the polarization of the faster, vertically traveling shear wave)

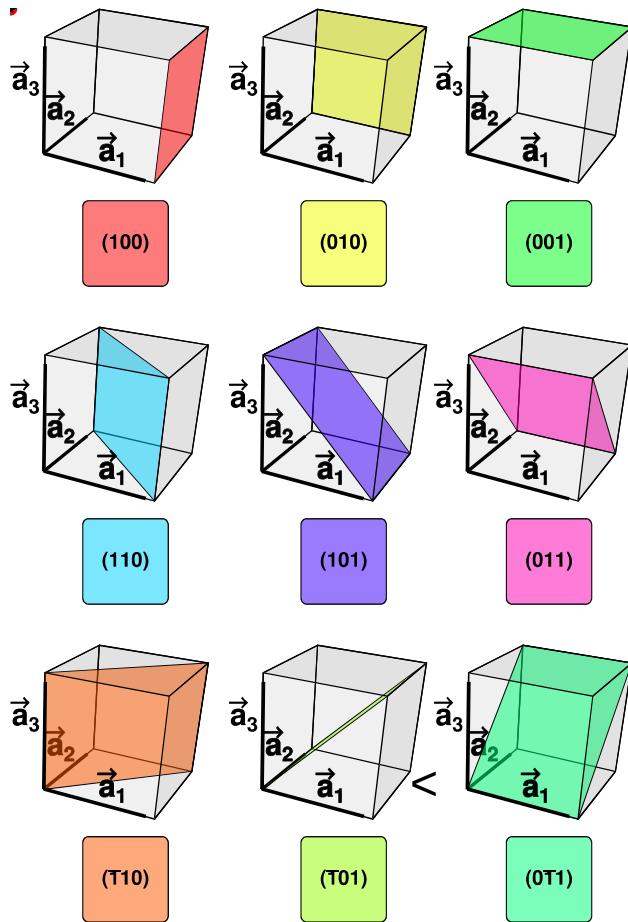
Fabric	Horizontal flow	Vertical planar flow
A-type	Parallel to flow	Small splitting
B-type	Normal to flow	Parallel to the plane
C-type	Parallel to flow	Normal to the plane
D-type	Parallel to flow	Small splitting
E-type	Parallel to flow	Small splitting

*V<sub>SH</sub>* / *V<sub>SV</sub>* anisotropy

Fabric	Horizontal flow	Vertical cylindrical flow
A-type	$V_{SH}/V_{SV} > 1$	$V_{SH}/V_{SV} < 1$
B-type	$V_{SH}/V_{SV} > 1$	$V_{SH}/V_{SV} > 1$ (weak)
C-type	$V_{SH}/V_{SV} < 1$	$V_{SH}/V_{SV} > 1$ (weak)
D-type	$V_{SH}/V_{SV} > 1$	$V_{SH}/V_{SV} < 1$
E-type	$V_{SH}/V_{SV} > 1$ (weak)	$V_{SH}/V_{SV} < 1$

A & E type most abundant

# LPO (5)



Dominant slip systems in  
Various types of olivine LPO

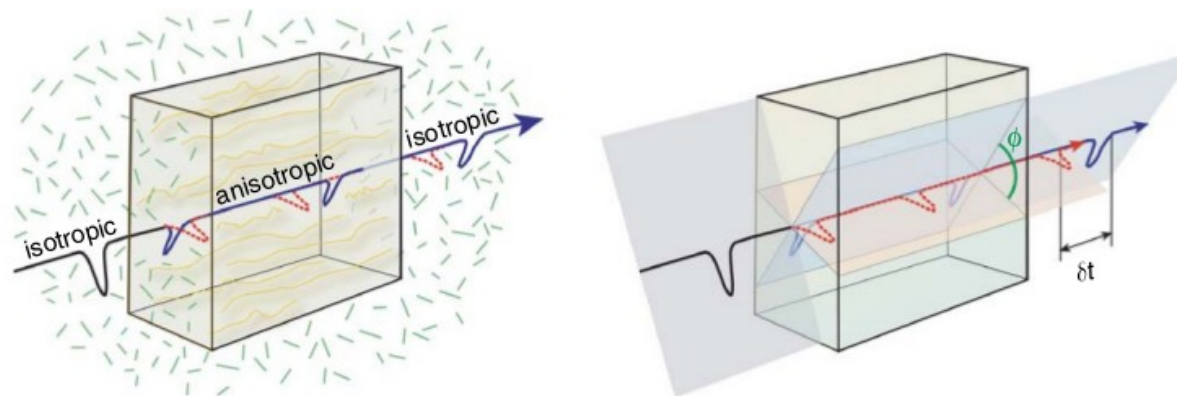
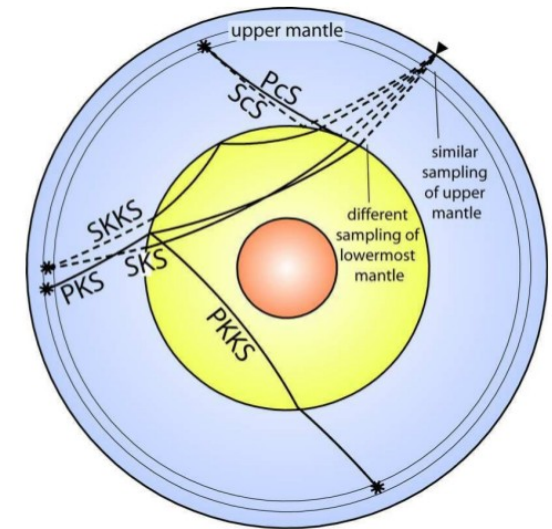
A-type	B-type	C-type
$[100] (010)$	$[001] (010)$	$[001] (100)$

D-type	E-type
$[100] \{0kl\}$	$[100] (001)$



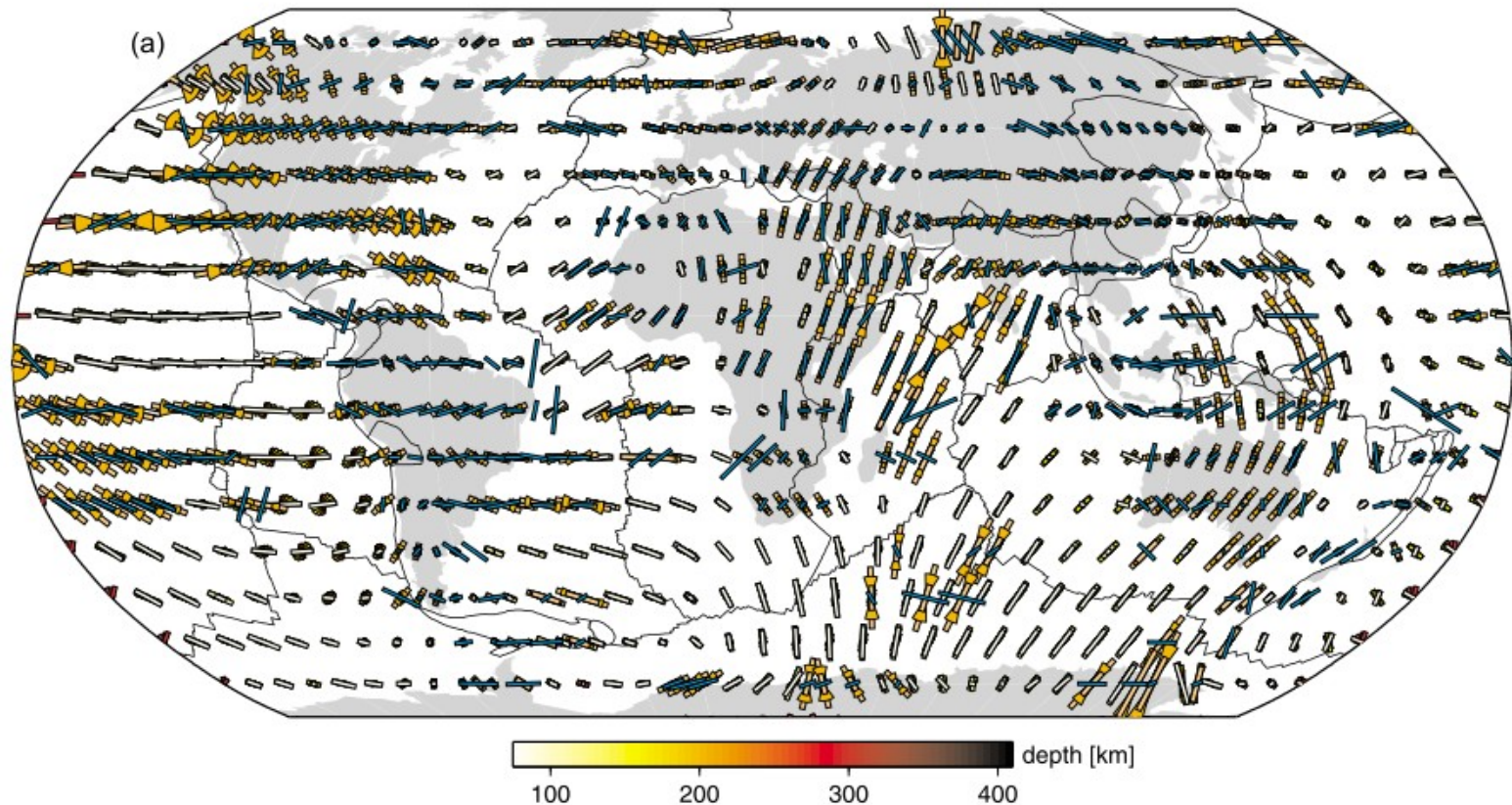
# SKS splitting (1)

- Vertically incident s-wave, transversing an anisotropic medium
- Polarization removed by transversing the liquid core
- Initial polarization of SKS is in plane of propagation (vertical)
- Splits in fast and slow polarized shear wave
- One can measure the fast axis, the delay time and the splitting angle
- Can be done on S, ScS, SS, PKS, but most popular are SKS and SKKS.
- Differential splitting of SKS-SKKS to infer lowermost mantle anisotropy



# SKS splitting (2)

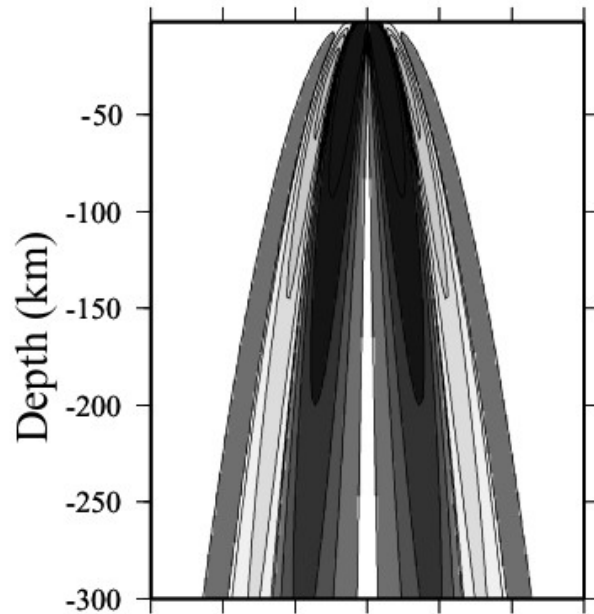
$2\psi$  azimuthal anisotropy fast axes well reconciled by spatially averaged global shear wave splitting



Becker (2011)

# SKS splitting (2)

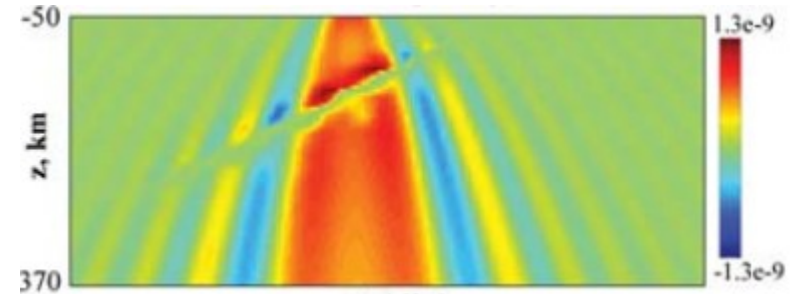
Favier & Chevrot (2003)



Analytical kernels,  
following the formalism  
of Dahlen (2000)

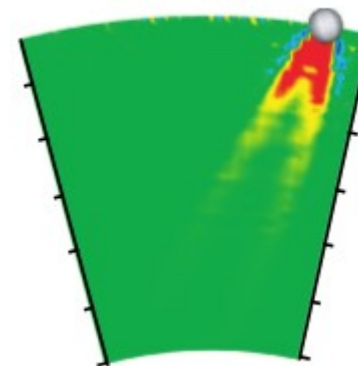
In the finite-frequency sense  
splitting has also vertical resolution

Long (2008)



2D kernels in heterogeneous  
models, adjoint method

Sieminski (2008)  
 $G_s/\rho$

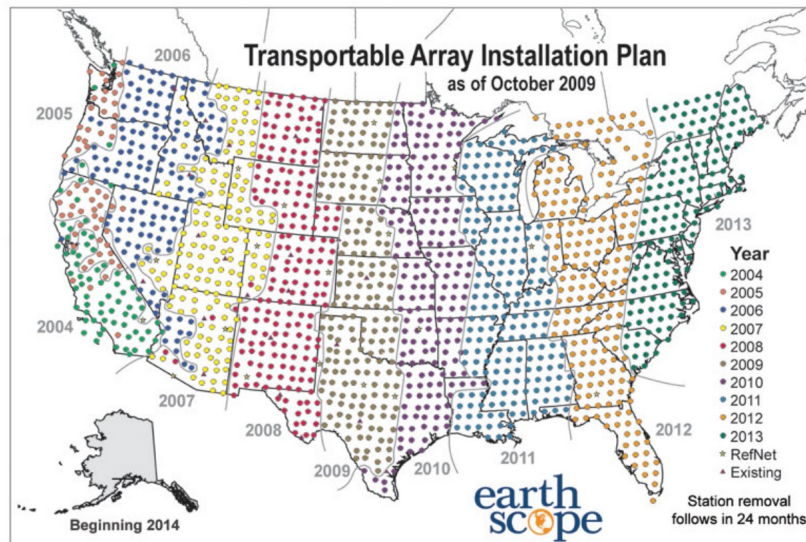


3D kernels in 3D  
models, Specfem,  
adjoint method

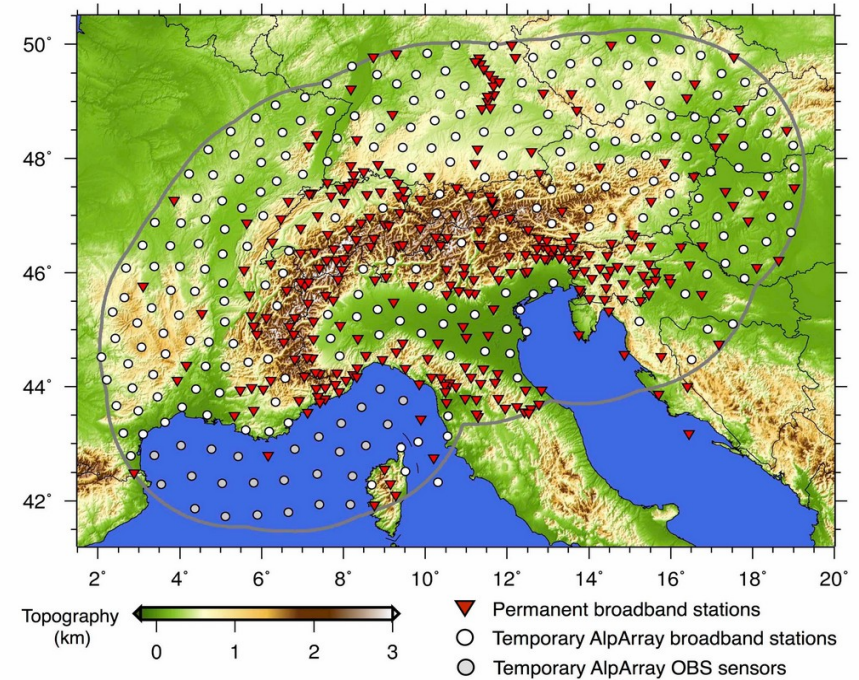


# Regional datasets / Mermaids

## USarray

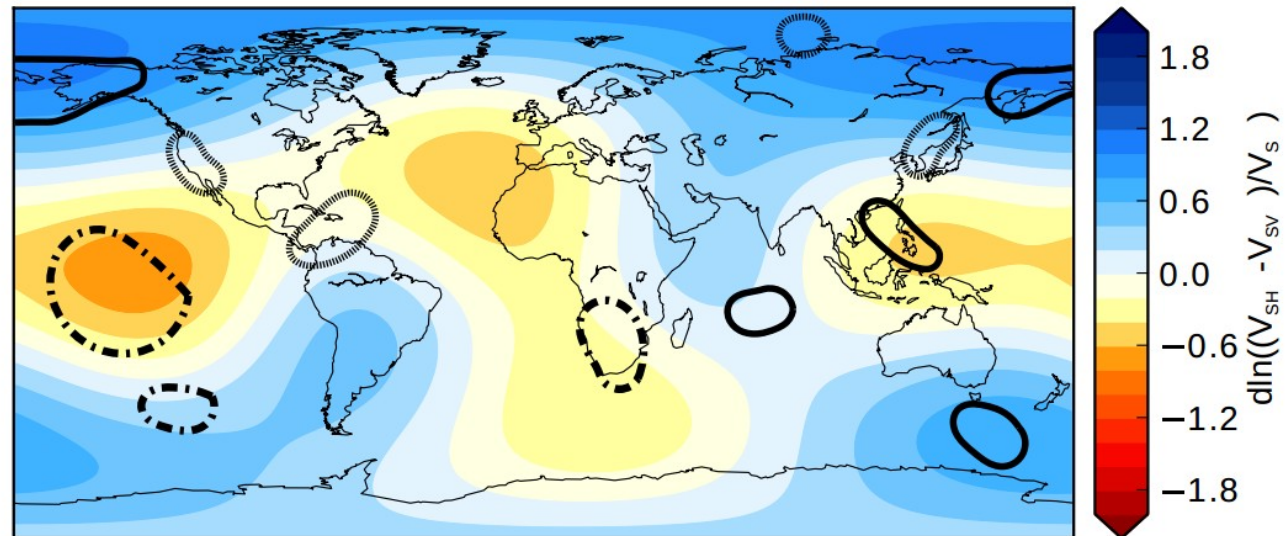



## Alparray





# Lower-mantle anisotropy (1)

Anisotropy at 2800 km depth, degrees 2 and 3 only



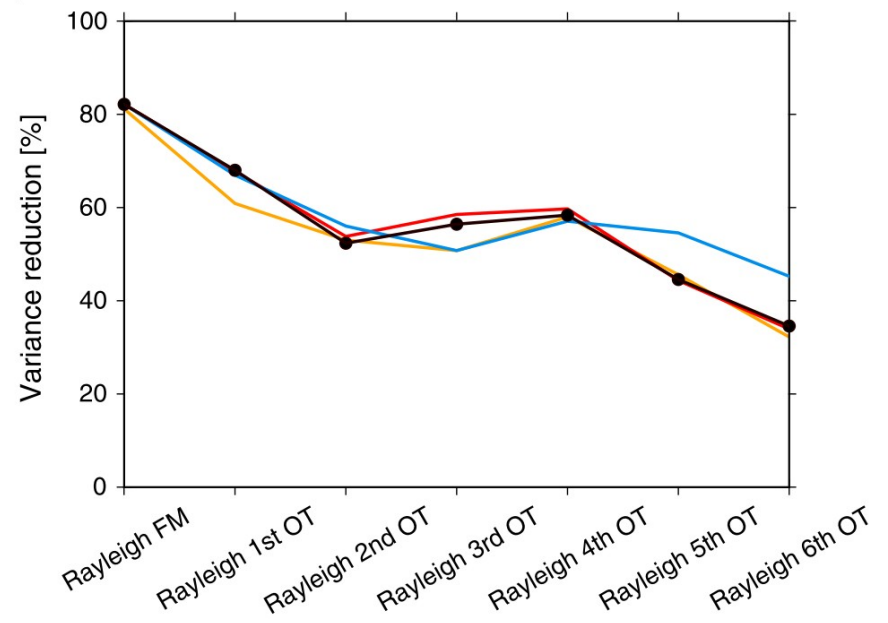
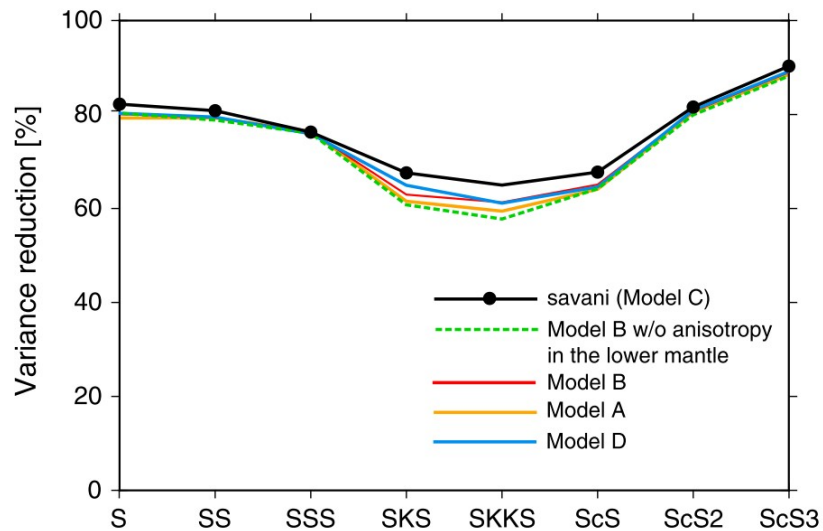
  $V_{SH} > V_{SV}$   
horizontal flow

  $V_{SH} > V_{SV}$   
or  $V_{SV} > V_{SH}$   
horizontal or  
vertical flow

 Tilted (TTI)  
anisotropy

# Lower-mantle anisotropy (2)

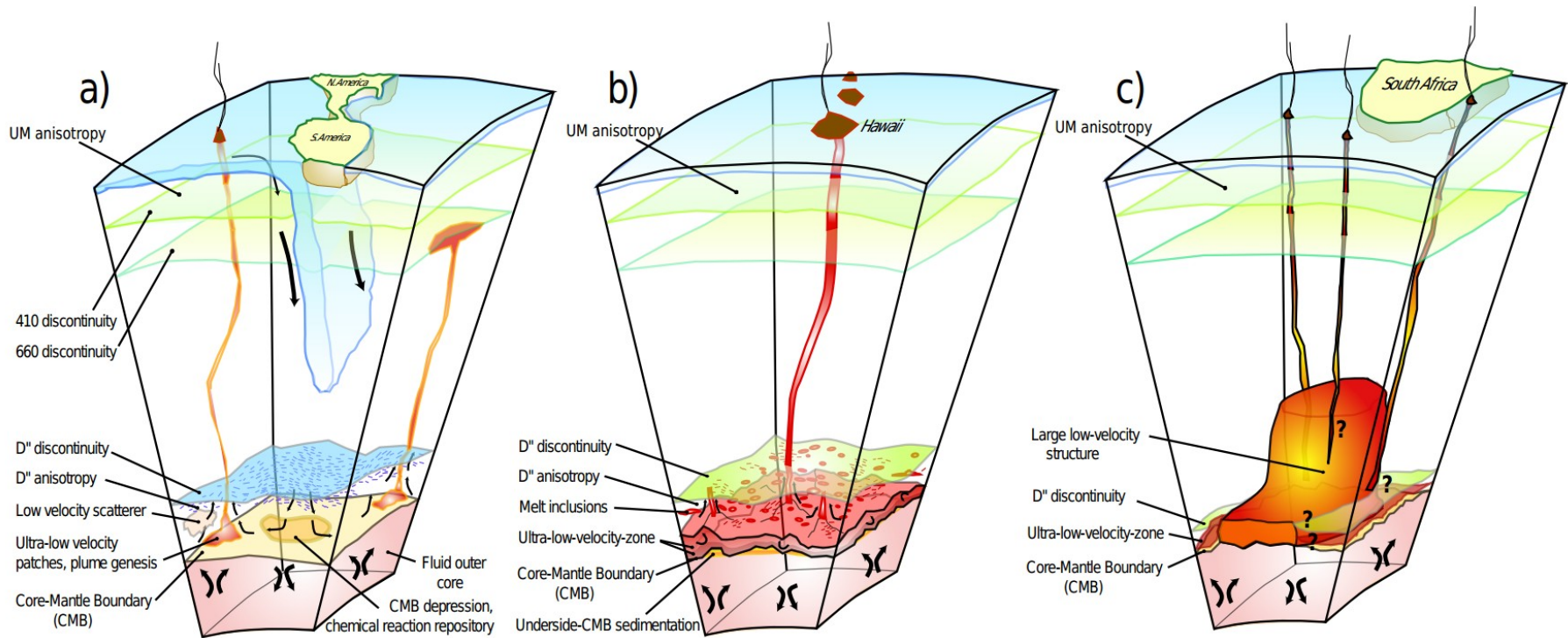
Required by the data?





# Lower-mantle anisotropy (3)

## Summary of LM observations



from Garnero (2004)

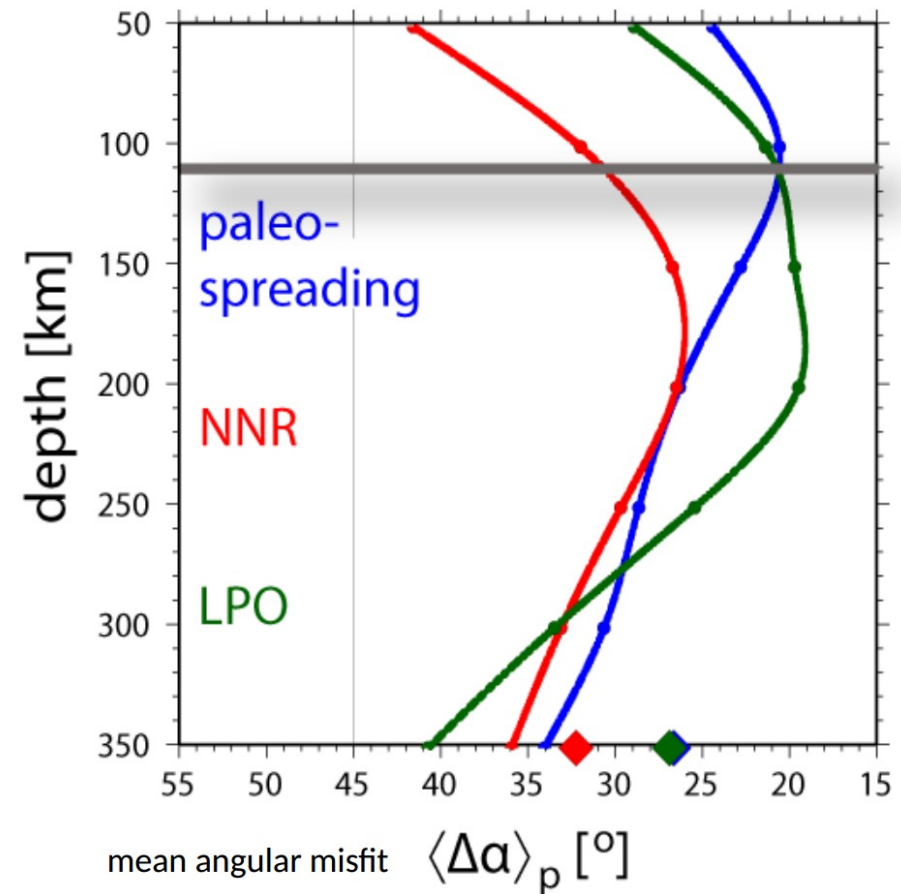
# Lower-mantle anisotropy (2)

Why is there a faster  $V_{SH}$  (horizontal flow) associated with regions of fast seismic velocity and faster  $V_{SV}$  (vertical flow) close to LLSVP's?

- PPV is the main candidate to explain CMB parallel LPO in the lowermost mantle, being less viscous than PV and intrinsically anisotropic
- Lithosphere sinks towards the CMB, undergoes the PPV phase transition, bends, and is subjected to horizontal deformation, causing LPO type anisotropy?
- Fast  $V_{SV}$  might be associated with vertical mass transport at zones of plume genesis near edges of LLSVPS

# Fossil (frozen-in) anisotropy vs. asthenospheric flow

- Paleo-spreading model match confined to shallow regions
- Define a mechanical Lithosphereasthenosphere “boundary” based on transition

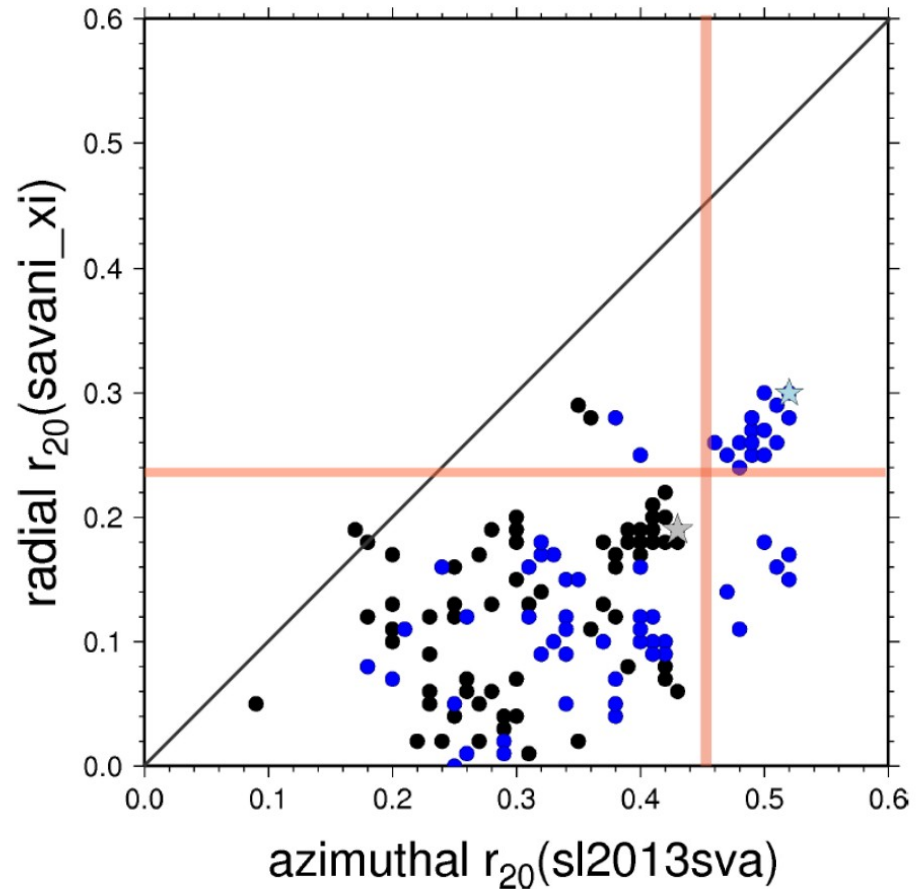


# Fit for different LPO models

## Azimuthal and radial

- Blue: oceanic correlation for 50-350 km
- Black: global correlation for 50-350 km

→ differently parameterized geodynamic models



# Azimuthal anisotropy (1)

Smith & Dahlen (1973) and Montagner & Nataf (1986) show that perturbations of surface wave phase velocities in media with weak hexagonal anisotropy is given by

$$\delta c = \frac{dc}{c} \approx D^0 + D_C^{2\phi} \cos(2\phi) + D_S^{2\phi} \sin(2\phi) \\ + D_C^{4\phi} \cos(4\phi) + D_S^{4\phi} \sin(4\phi),$$

Medium with a single fast or slow symmetry axis and isotropic velocities in the plane perpendicular to it.

Special cases: Azimuthal anisotropy (2 $\psi$  and 4 $\psi$  terms)

Radial anisotropy (0 $\psi$  terms)

No 1 $\psi$  and 3 $\psi$  terms

Coefficients in D depend on 13 independent elastic constants

# Azimuthal anisotropy (2)

From 21 to 13 independent elastic parameters, of which we resolve only a couple at once

- $2\Psi$ -azimuthal term:

$$\begin{aligned} \alpha_1 \cos 2\Psi & & \alpha_2 \sin 2\Psi \\ B_c = \frac{1}{2}(C_{11} - C_{22}) & & B_s = C_{16} + C_{26} \\ G_c = \frac{1}{2}(C_{55} - C_{44}) & & G_s = C_{54} \\ H_c = \frac{1}{2}(C_{13} - C_{23}) & & H_s = C_{36} \end{aligned}$$

- $4\Psi$ -azimuthal term:

$$\begin{aligned} \alpha_3 \cos 4\Psi & & \alpha_4 \sin 4\Psi \\ E_c = \frac{1}{8}(C_{11} + C_{22}) + \frac{1}{4}C_{12} - \frac{1}{2}C_{66} & & E_s = \frac{1}{2}(C_{16} - C_{26}) \end{aligned}$$

- Constant term ( $0\Psi$ -azimuthal term:  $\alpha_0$ )

$$\begin{aligned} A = \rho V_{PH}^2 &= \frac{3}{8}(C_{11} + C_{22}) + \frac{1}{4}C_{12} + \frac{1}{2}C_{66} \\ C = \rho V_{PV}^2 &= C_{33} \\ F = \frac{1}{2}(C_{13} + C_{23}) \\ L = \rho V_{SV}^2 &= \frac{1}{2}(C_{44} + C_{55}) \\ N = \rho V_{SH}^2 &= \frac{1}{8}(C_{11} + C_{22}) - \frac{1}{4}C_{12} + \frac{1}{2}C_{66} \end{aligned}$$

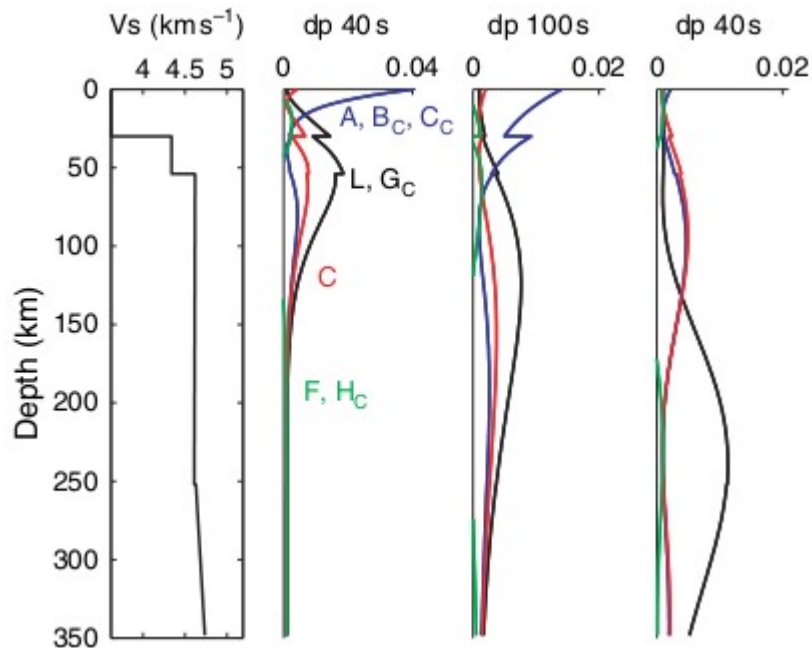
Azimuthal: 3 (Vs, Gs, Gc), horizontal symmetry, Radial: 2 (Vs, Xi) vertical symmetry

Vectorial tomography: 4 (Vs, Xi, two angles of orientation of symmetry axis).



# Azimuthal anisotropy (3)

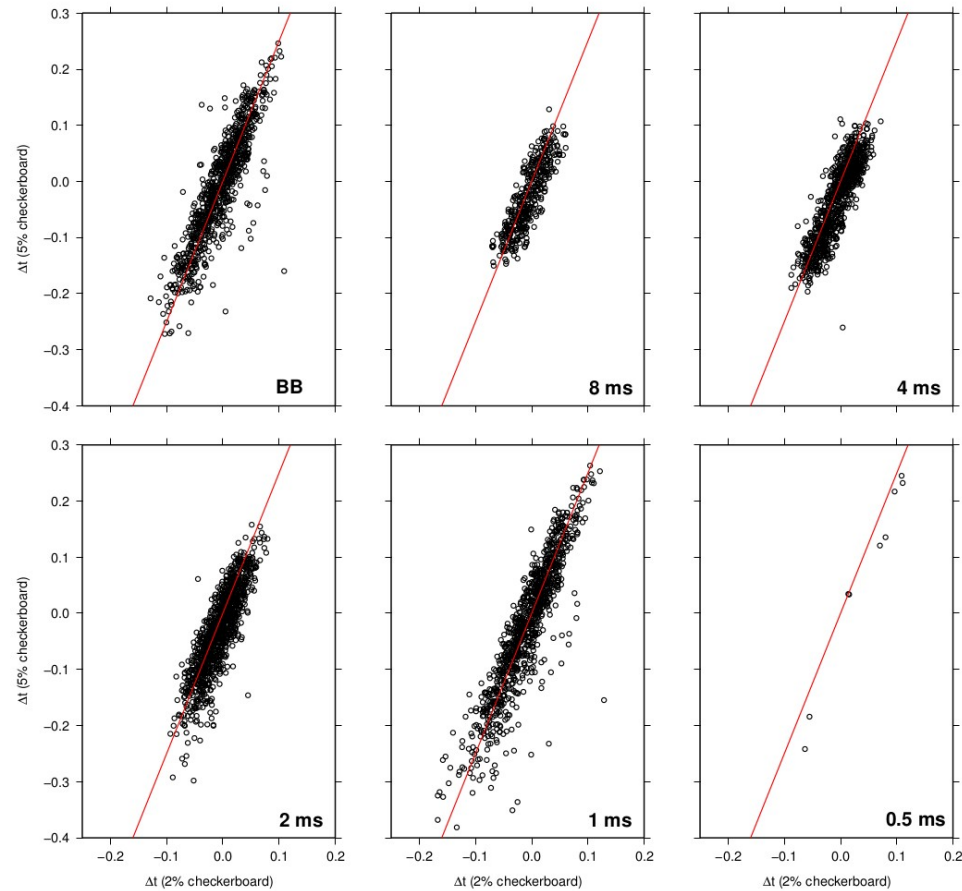
Montagner & Nataf (1986) show that A,C,L,N,F kernels are Equivalent to the azimuthal anisotropic kernels



We can derive The fast axis direction and the Amplitude of azimuthal anisotropy

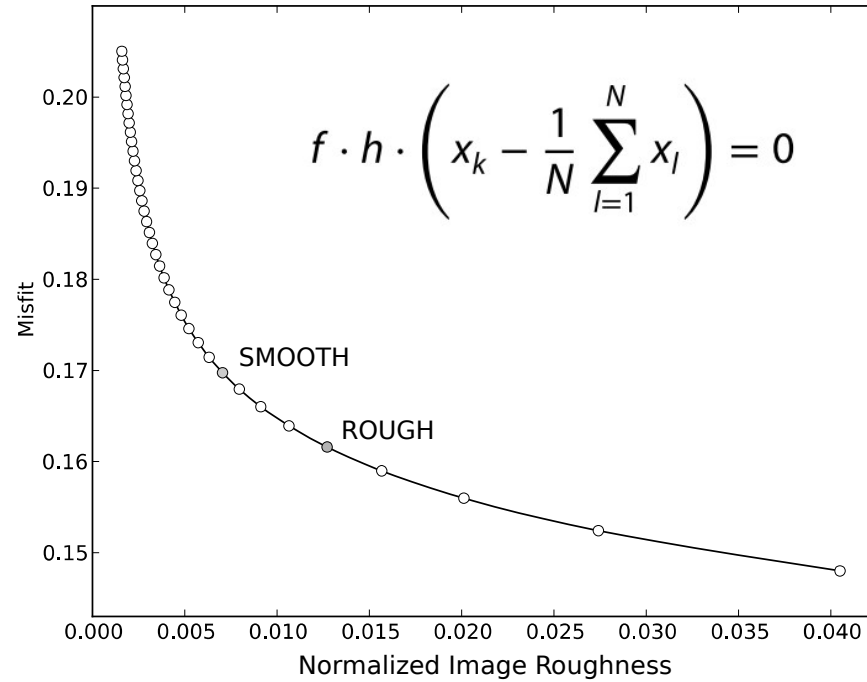
$$\left\{ \begin{array}{l} \Lambda_{2\Psi} = \sqrt{A_{2\Psi}^2 + B_{2\Psi}^2} \\ \Theta_{2\Psi} = \frac{1}{2} \arctan\left(\frac{B_{2\Psi}}{A_{2\Psi}}\right) \end{array} \right. \text{ and } \left\{ \begin{array}{l} \Lambda_{4\Psi} = \sqrt{A_{4\Psi}^2 + B_{4\Psi}^2} \\ \Theta_{4\Psi} = \frac{1}{4} \arctan\left(\frac{B_{4\Psi}}{A_{4\Psi}}\right) \end{array} \right.$$

# Linearity of CC delay times (1)



**Figure 4.** Cross-correlation delay times measured for the  $\pm 5\%$  versus those in the  $\pm 2\%$  model. The red line denotes a slope of 5/2. BB is for broadband data, and the other 5 bands with dominant periods of 8 ms, 4 ms, 2 ms, 1 ms and 0.5 ms. Very few correlations pass the condition that  $R > 0.8$  for the passband with dominant period of 0.5 ms.

# Regularization options



- Minimum norm model
- Smoothest model
- Fewest wavelets

# LPO model of Becker (2008)

*LPO*

(Becker et al., 2008)

reference frame shear	none (NNR)
density inferred from	S362WANI (Kustowski et al., 2008)
upper thermal boundary layer	excluded around cratons
upper mantle	
background viscosity, $\eta_{um}$	average $\approx 1.8 \times 10^{21}$ Pas, non-Newtonian
asthenospheric viscosity	temperature and stress dependent (Becker, 2006)  ~ three orders of magnitude variations in upper mantle
velocity gradients	form LPO when in dislocation creep
method of LPO estimate	full DREX (Kaminski et al., 2004) for A type LPO

# LPO model of Becker (2008)

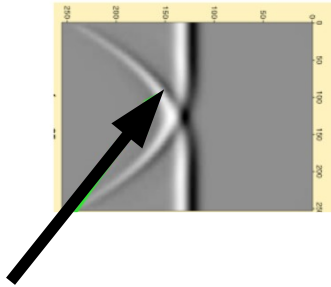
## Global circulation modeling

- × Instantaneous Stokes flow solution:  $u \propto \Delta \rho / \eta$
- × Solve with CitcomS (Zhong et al., 2000), with modifications as in Becker and Faccenna (2011), ~20 km resolution
- × Density anomalies:
  - Simple scaling of velocity anomalies from tomography, no chemical anomalies besides continental keels
  - Assigned to Wadati-Benioff zones for slabs
- × Rheology:
  - Layered viscosity structure, with lateral viscosity variations, Newtonian (can do power-law, realistic rheologies)
  - Plate boundaries are weak (low viscosity) zones

(Becker et al., 2008)

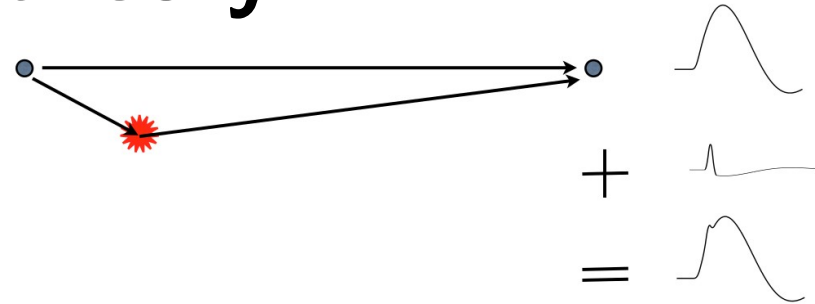
# Born theory

1)



A wavefront heals because energy  
Diffracts around the anomaly, making  
anomaly hard to detect

2)



In the first-order Born approximation  
this is modeled via single scattering

3)

Considering cross-correlation delay times:  
Scattered wave perturbs delay time

$$\delta T = -\frac{\delta \dot{\gamma}(0)}{\ddot{\gamma}(0)} = -\frac{\int_{-\infty}^{\infty} \dot{u}(t') \delta u(t') dt'}{\int_{-\infty}^{\infty} \ddot{u}(t') u(t') dt'}$$

$$\delta \bar{u}^{PP}(\omega) = \frac{\omega^2}{4\pi r V_P^2} \left[ \frac{\delta \rho}{\rho} - \frac{\delta \lambda}{\lambda + 2\mu} - 2 \frac{\delta \mu}{\lambda + 2\mu} \right] e^{ik_{pr}} dV$$

Born

4)

Line integrals

$$T \approx \int_{P_0} \frac{ds}{c(\mathbf{r})}$$

Turn into volume integrals

$$\delta T = \int K_P(\mathbf{r}_x) \frac{\delta V_P}{V_P} d^3 \mathbf{r}_x$$

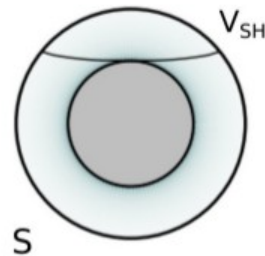
Can also use Adjoint method



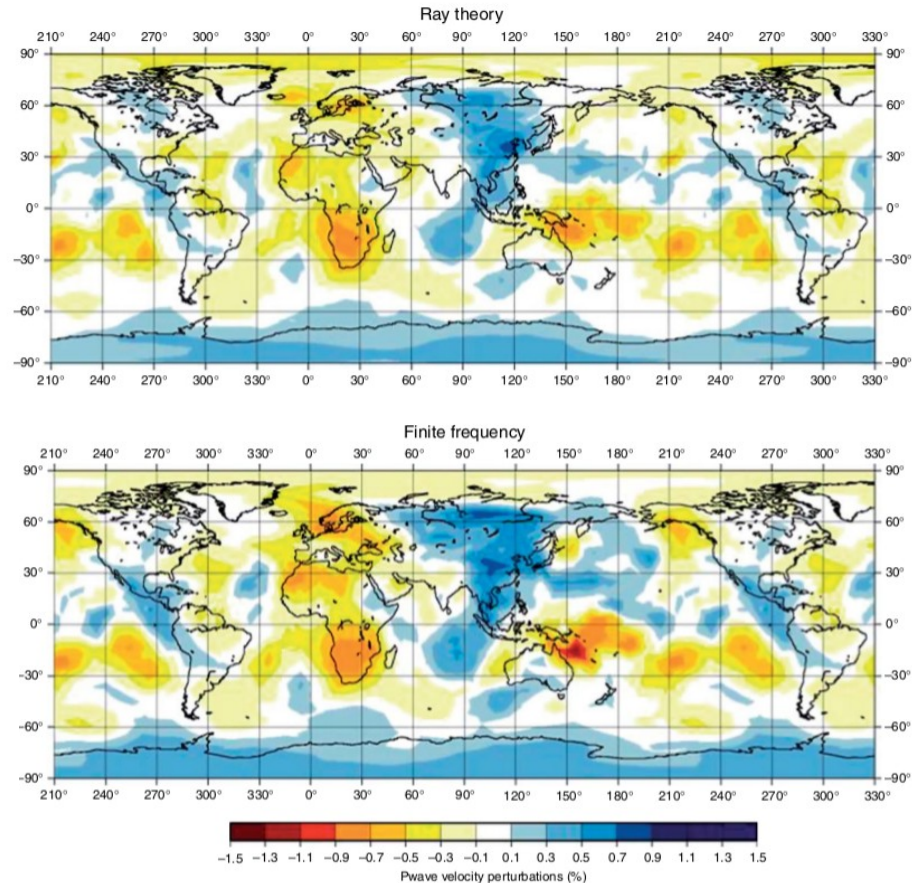
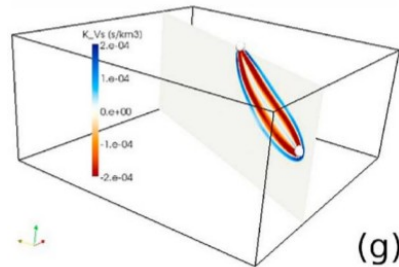
# Rays vs. finite-frequency kernels

Does replacing 0-th with 1<sup>st</sup> order (single scattering) theory play a role, globally?

Ray theory



FF-theory



Van der Hilst (2002)

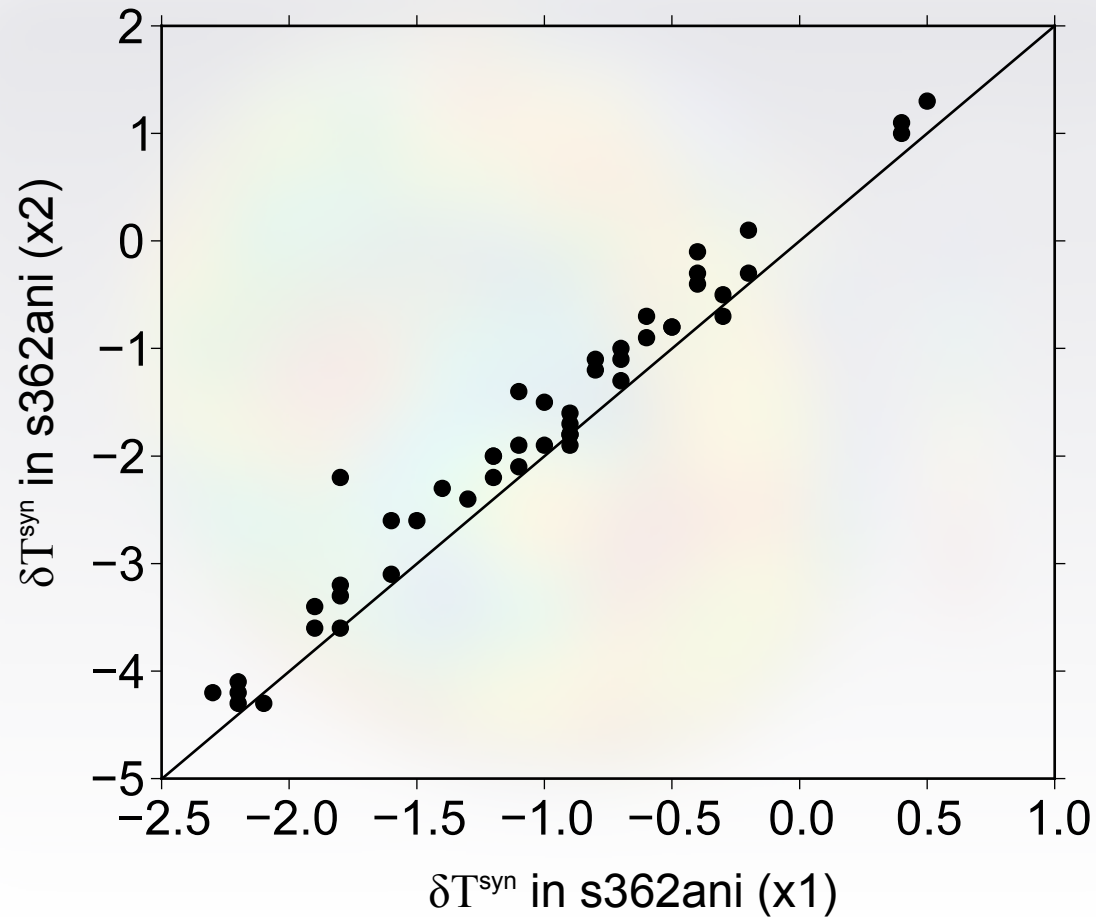
Montelli (2005)

Data coverage is much more of a problem. But regionally, it makes a difference!

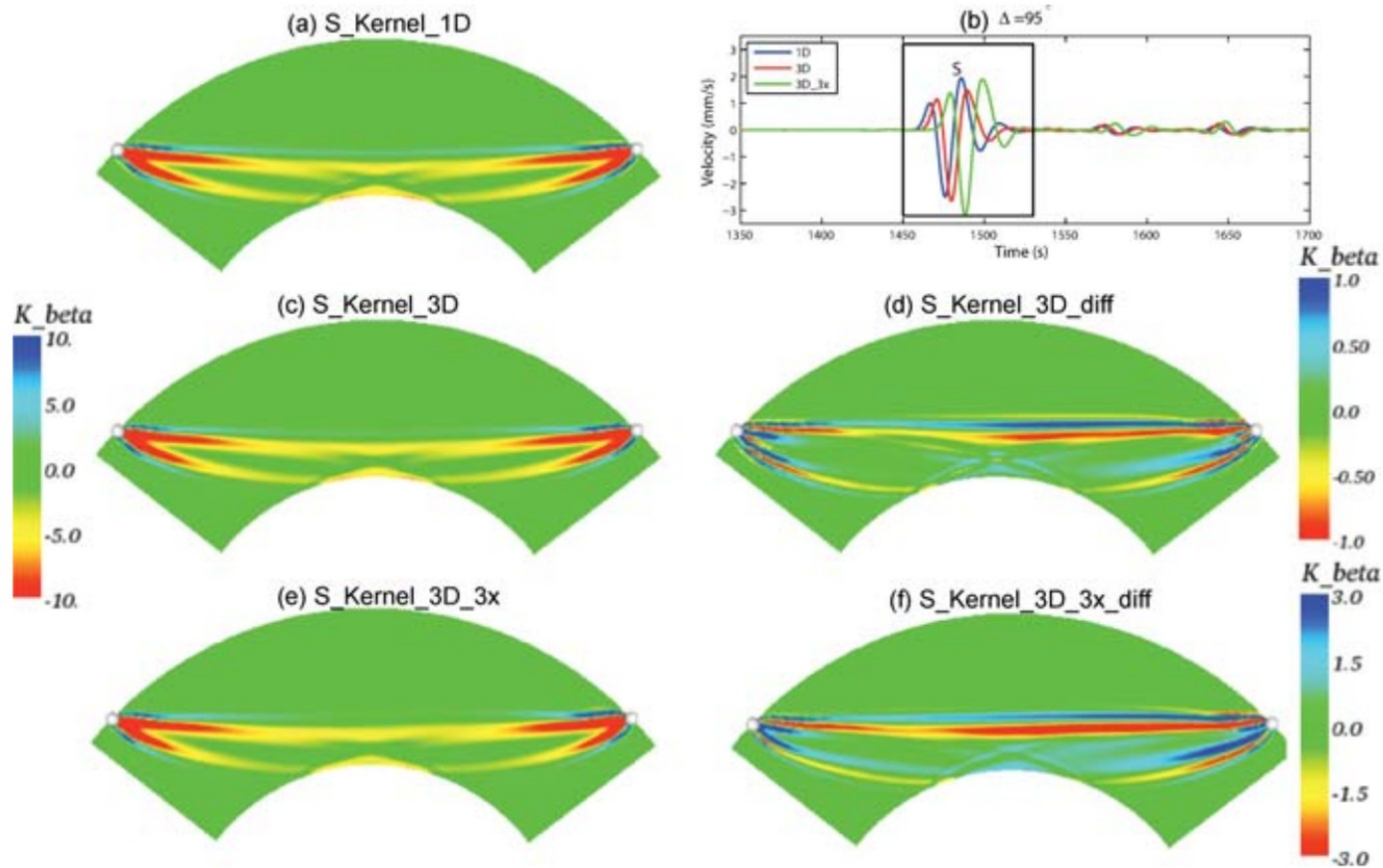
→ Ray theory where sufficient, FF-theory, where required!

# Linearity of CC delay times (1)

Quality of kernel linearity



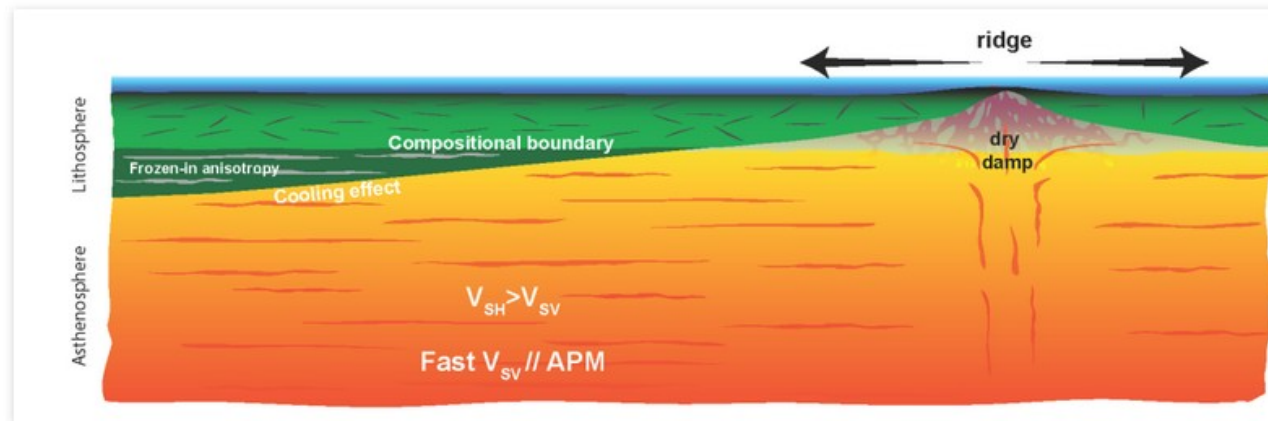
# Linearity of CC delay times (2)



**Figure 15.** (a) 9 s  $K_{\beta}$  kernel for the  $S$  phase at a distance  $\Delta = 95^\circ$  for the 1-D PREM model. (b) Transverse-component velocity seismograms for the source–receiver pair described in Fig. 14 for various reference models: 1-D (PREM), 3-D (S20RTS) and 3D\_3x (S20RTS enhanced by a factor of 3). (c) Same as (a), but for the 3-D model. (d) The difference between the kernels in (c) and (a). (e) Same as (a), but for the 3D\_3x model. (f) The difference between the kernels in (e) and (a). Note that the colour bars for figures (d) and (f) are different from the colour bar for figures (a), (c) and (e).

# Old results: Model of Beghein (2014)

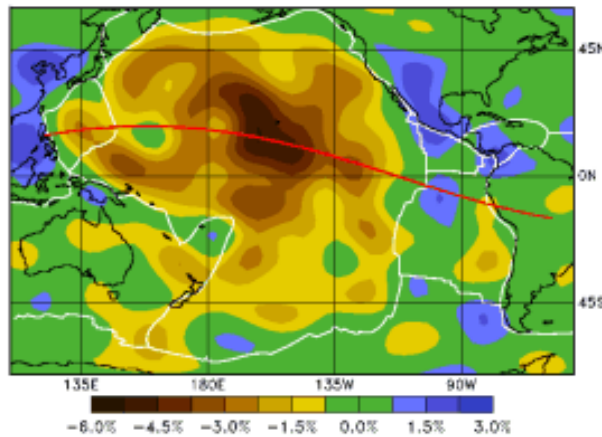
- Supplementary Material (42 pages) [here](#) - Article usage stats [here](#)



*This figure illustrates the Earth's upper mantle beneath the Pacific ocean. The orange layer represents the deformable, warm asthenosphere in which there is active mantle flow. The green layer on top represents the lithospheric plate, which forms at the mid ocean ridge, then cools down and thickens as it moves away from the ridge. The cooling of the plate overprints a compositional boundary that forms at the ridge by dehydration melting and is preserved as the plate ages. The more easily deformable, hydrated rocks align with mantle flow. The directions of past and present-day mantle flow can be detected by seismic waves, and changes in the alignment of the rocks inside and at the bottom of the plate can be used to identify layering. CREDIT: Nicholas Schmerr (University of Maryland)*

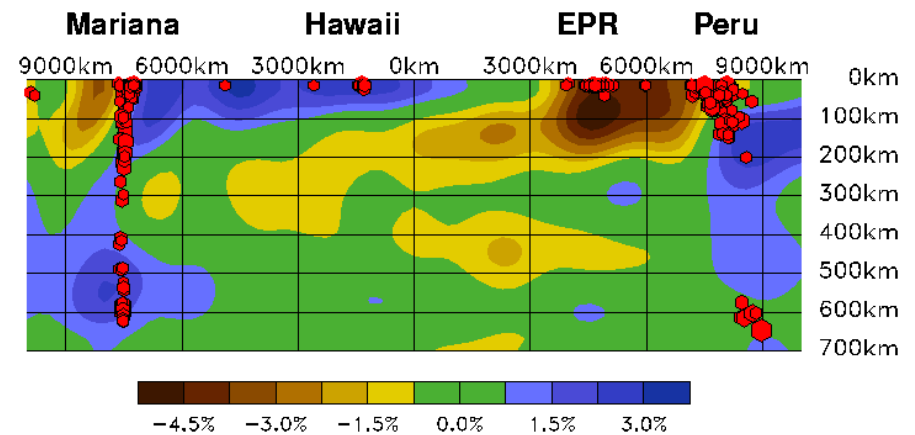
# Old results: Ekström (1998)

## Radial anisotropy in the Central Pacific

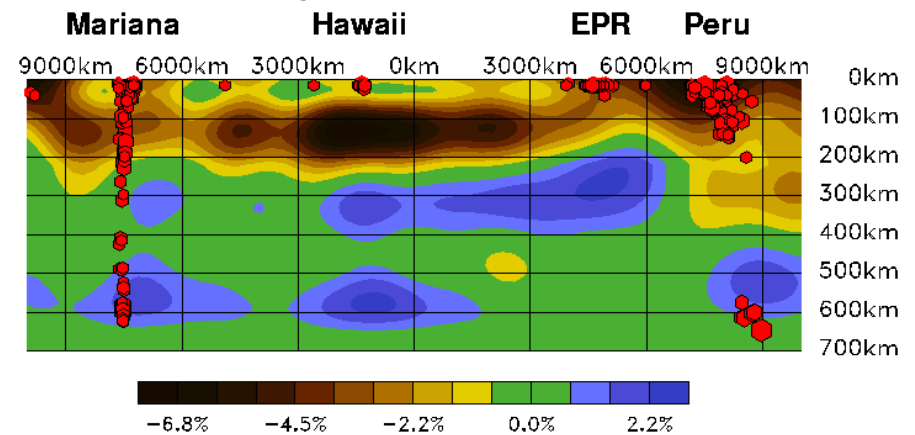


“Pancake” like structure of  
Above average  $\xi$  with a  
Peak at ~ 150 km depth

## Isotropic cross-section

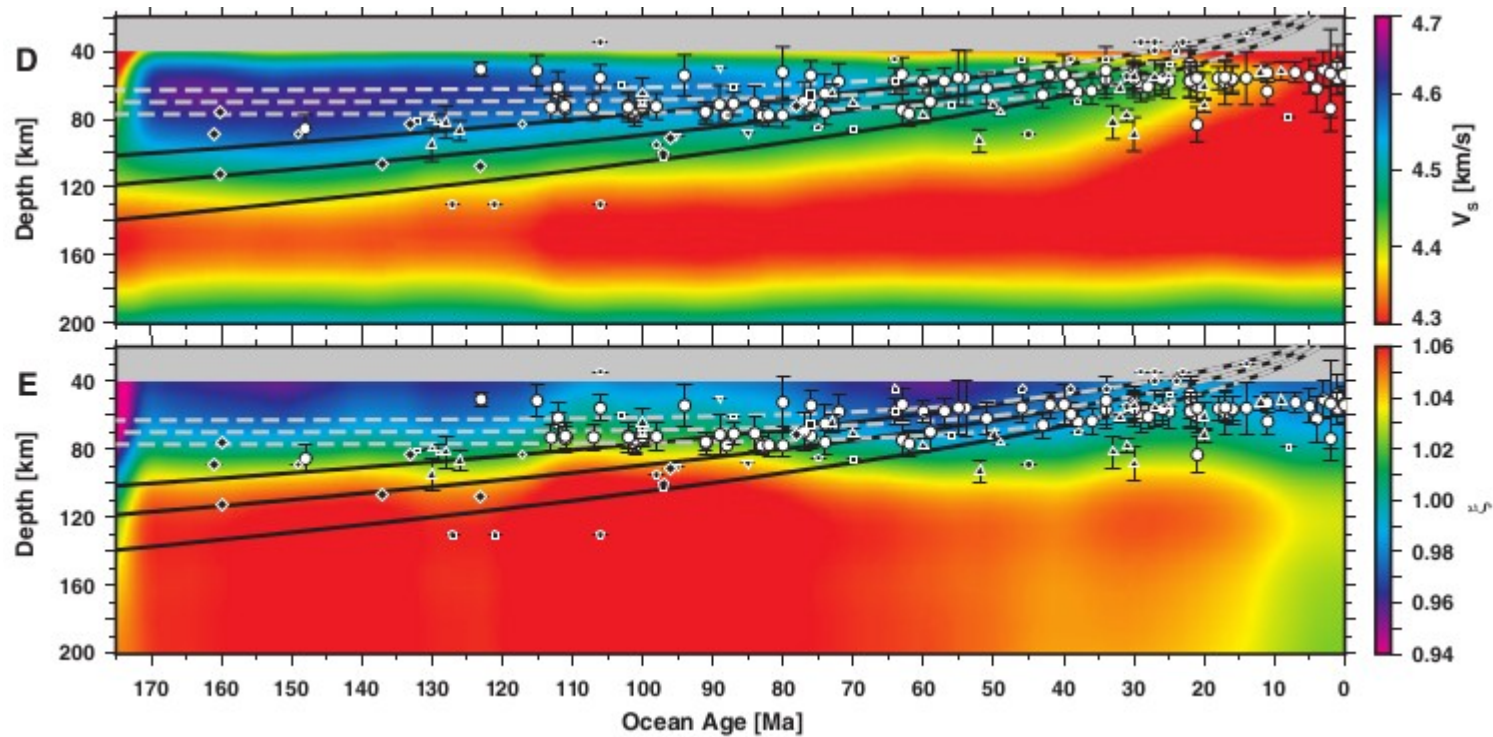


## Anisotropic cross-section





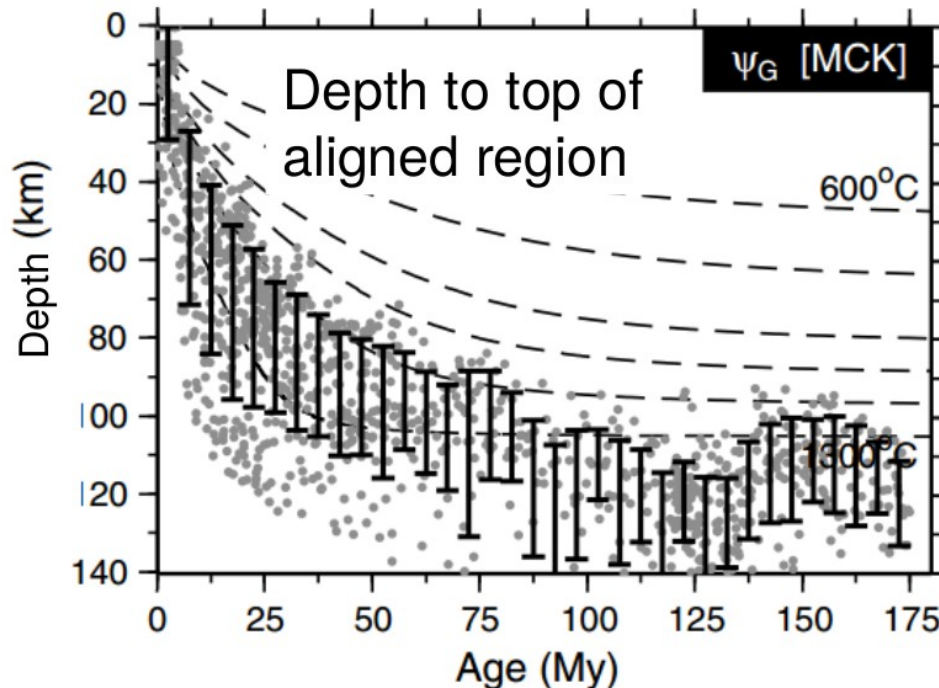
# Old results: Beghein (2015)





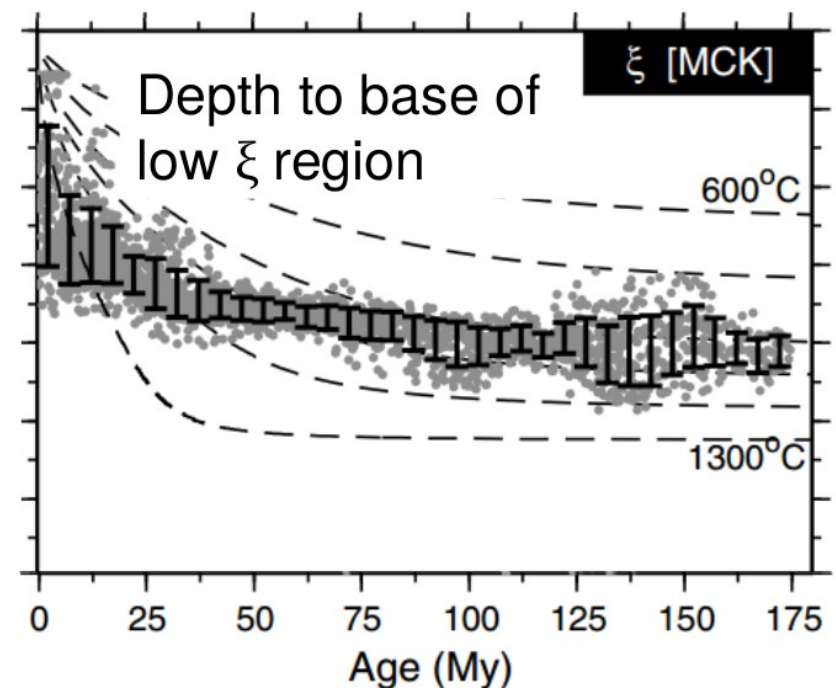
# Old results: Burgos (2014)

## azimuthal anisotropy



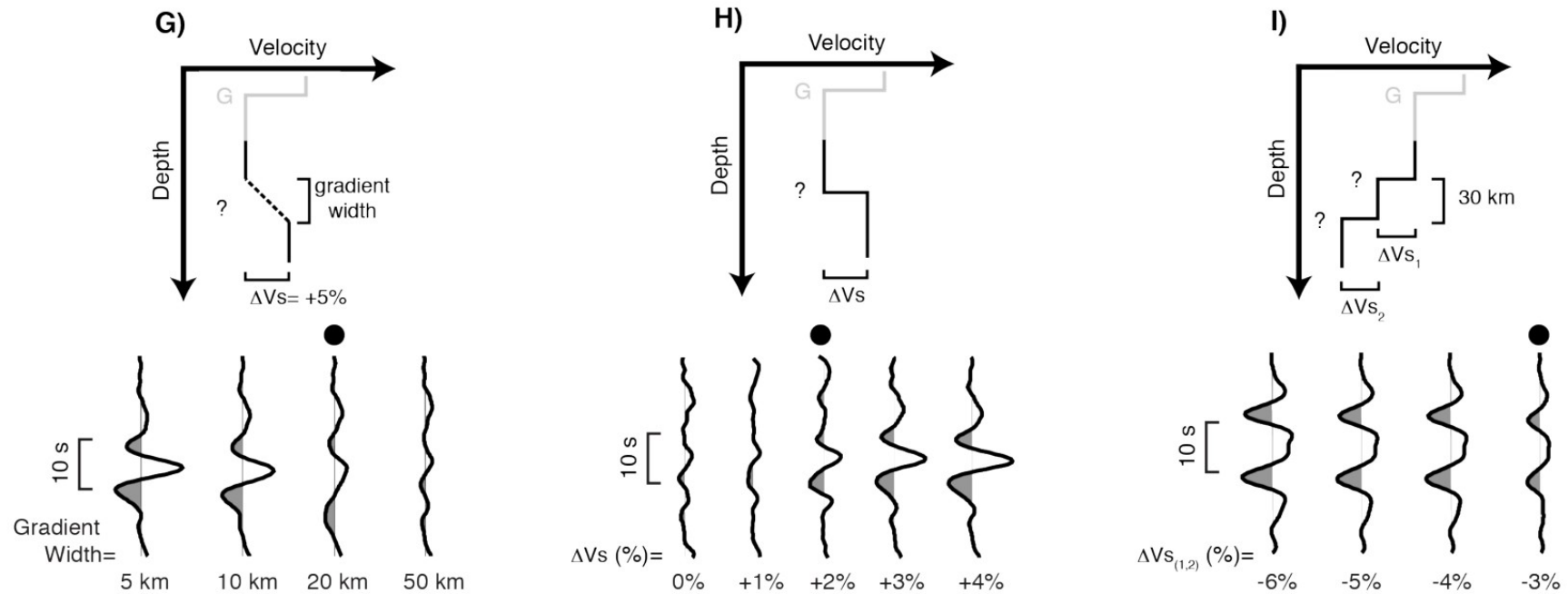
seems to follow  
thermal model - OK!

## radial anisotropy



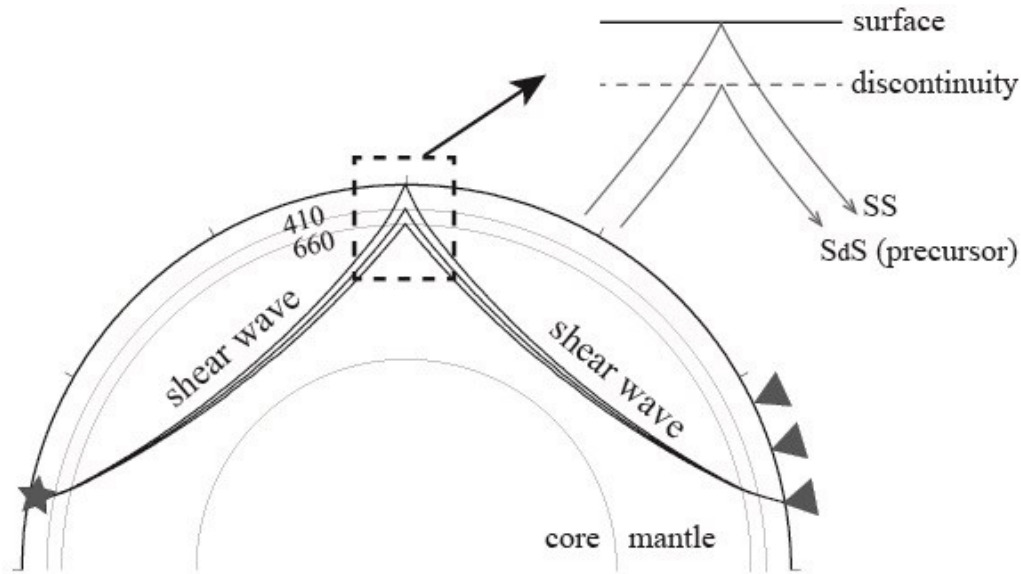
does not seem to follow  
thermal model - ?!

# SS precursor modeling, Schmerr (2012)

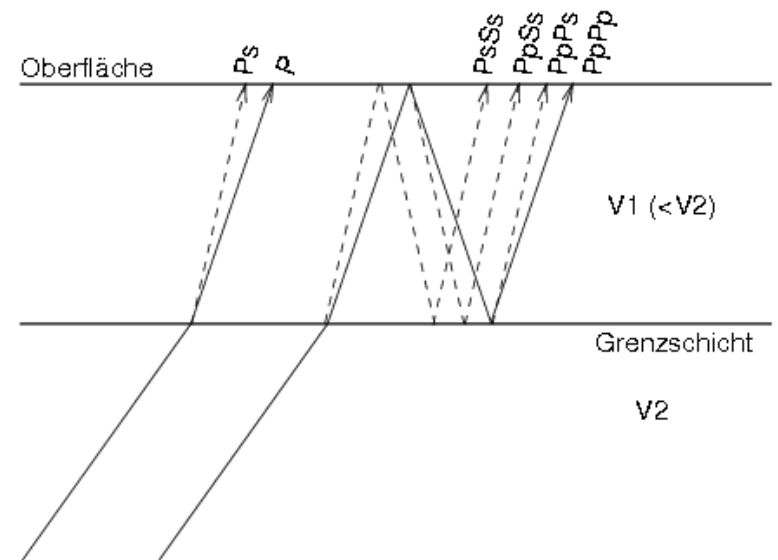


Which of them are compatible with our model?

# SS precursors, Receiver functions



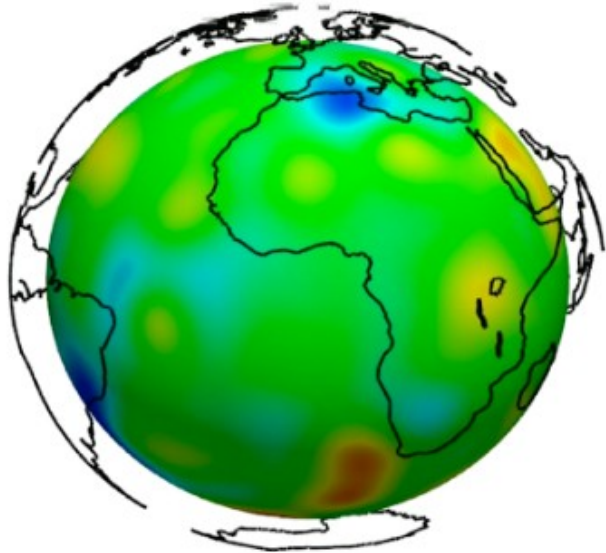
Look at SS precursors  
reflected at interfaces



Look at time difference  
With converted waves

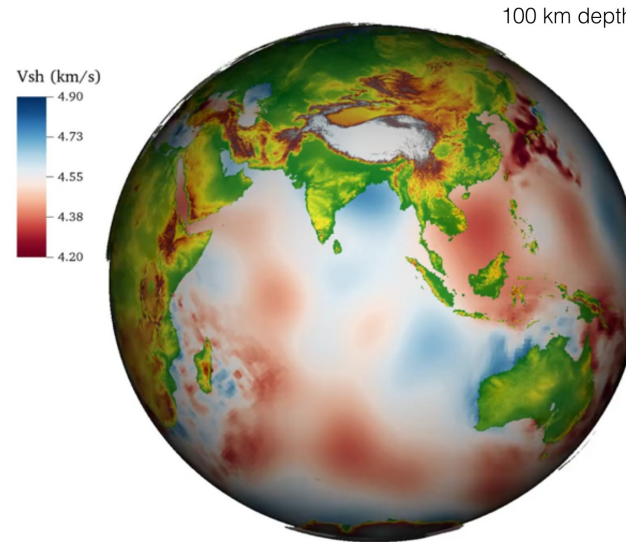
# Recent approaches to global modeling

Adjoint (Bozdag, 2012)



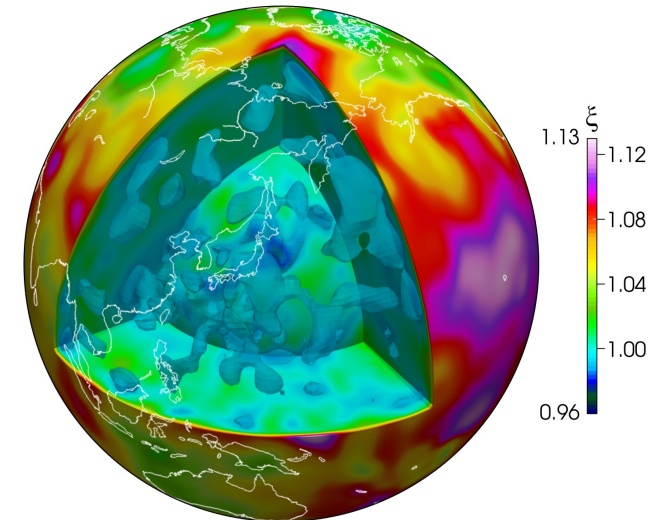
Global adjoint tomography  
 $T < 30$  s, 300 EQ's, 1 it.

CEM (Afanasiev et al. 2015)



Method agnostic composite  
model from regional waveform  
tomographys and S20RTS

SAVANI (Auer et al., 2014)



Compilation of multiple  
Global datasets, reinversion;  
Next: add regional data

# Adjoint vs. classical tomography

	classical tomography	adjoint tomography
reference model	1D	3D
physical domain	3D	3D
Born approximation	yes	yes
forward modelling technique	e.g., ray theory, modes, or banana-doughnut kernels	fully numerical (e.g., SEM)
gradient method	$\mathbf{g} = -\mathbf{G}^T \mathbf{d}$ $G_{ik} = \int_V K_i B_k d^3 \mathbf{x}$	$g_k = \int_V K B_k d^3 \mathbf{x}$
Newton method	$\mathbf{G}^T \mathbf{G} \delta \mathbf{m} \approx -\mathbf{g}$	(too costly)
number of iterations	1	multiple

# Adjoint vs. classical tomography

## Classical tomography

We can accumulate the full approximate Hessian

In the framework of non-linear optimization theory one would call it a “Gauss-Newton type” method

The approximate nature of our forward theory restricts ourselves to 1D reference models, thus we can only do 1 inversion step

## Adjoint tomography

One is restricted to compute so called “Event Kernels” i.e. a sum of misfit Kernels

Computing individual kernels for all different event station pairs would be to expensive

Would theoretically work with every available wiggle in the record but in practice people employ more robust misfit functionals (e.g. traveltimes)

Limited to rather low frequency and small domains

Focuses on the background model



# Kernel expressions

Table B.1: Sensitivity kernel expressions for three different parameterizations of elastic structure.  $\mathbf{D}$  is the deviatoric strain (e.g., *Liu and Tromp, 2006, Eq. 28*).

Model Parameter		Notation	Kernel Expression
Bulk modulus	$\kappa$	$K_{\kappa(\mu\rho)}(\mathbf{x})$	$-\kappa \int_0^T [\nabla \cdot \mathbf{s}^\dagger(\mathbf{x}, T-t)][\nabla \cdot \mathbf{s}(\mathbf{x}, t)] dt$
Shear modulus	$\mu$	$K_{\mu(\kappa\rho)}(\mathbf{x})$	$-2\mu \int_0^T \mathbf{D}^\dagger(\mathbf{x}, T-t) : \mathbf{D}(\mathbf{x}, t) dt$
Density	$\rho$	$K_{\rho(\kappa\mu)}(\mathbf{x})$	$-\rho \int_0^T \mathbf{s}^\dagger(\mathbf{x}, T-t) \cdot \partial_t^2 \mathbf{s}(\mathbf{x}, t) dt$
Bulk sound speed	$c$	$K_{c(\beta\rho)}(\mathbf{x})$	$2 K_{\kappa(\mu\rho)}$
S wavespeed	$\beta$	$K_{\beta(c\rho)}(\mathbf{x})$	$2 K_{\mu(\kappa\rho)}$
Density	$\rho$	$K_{\rho(c\beta)}(\mathbf{x})$	$K_{\rho(\kappa\mu)} + K_{\kappa(\mu\rho)} + K_{\mu(\kappa\rho)}$
P wavespeed	$\alpha$	$K_{\alpha(\beta\rho)}(\mathbf{x})$	$\left(2 + \frac{8\mu}{3\kappa}\right) K_{\kappa(\mu\rho)}$
S wavespeed	$\beta$	$K_{\beta(\alpha\rho)}(\mathbf{x})$	$2 K_{\mu(\kappa\rho)} - \frac{8\mu}{3\kappa} K_{\kappa(\mu\rho)}$
Density	$\rho$	$K_{\rho(\alpha\beta)}(\mathbf{x})$	$K_{\rho(\kappa\mu)} + K_{\kappa(\mu\rho)} + K_{\mu(\kappa\rho)}$

# “Bayesian” generalized non-linear least-squares inversion (vs. discrete regularization)

Assume: a posteriori model parameters are distributed according to a Gaussian PDF

$$\rho(\mathbf{m}) \propto \exp\left(-\frac{1}{2}\mathbf{m} \cdot \mathbf{C}_m^{-1} \cdot \mathbf{m}\right) \quad C_m(\mathbf{r}_1, \mathbf{r}_2) = \sigma^2 \exp\left(-\frac{|\mathbf{r}_1, \mathbf{r}_2|^2}{2L^2}\right)$$

For a linear problem the solution is given by (Tarantola & Valette, 1982)

$$\begin{aligned} \mathbf{m} &= \mathbf{C}_m \cdot \mathbf{G}^T \cdot (\mathbf{C}_d + \mathbf{G} \cdot \mathbf{C}_m \cdot \mathbf{G}^T)^{-1} \cdot \mathbf{d} \\ &= (\mathbf{C}_m^{-1} + \mathbf{G}^T \cdot \mathbf{C}_d^{-1} \cdot \mathbf{G})^{-1} \cdot \mathbf{G}^T \cdot \mathbf{C}_d^{-1} \cdot \mathbf{d} \end{aligned}$$

- To solve slightly non-linear problems, this algorithm can be iterated
- Provides a formalized way to introduce prior knowledge

Compare to “Classic” formulation:

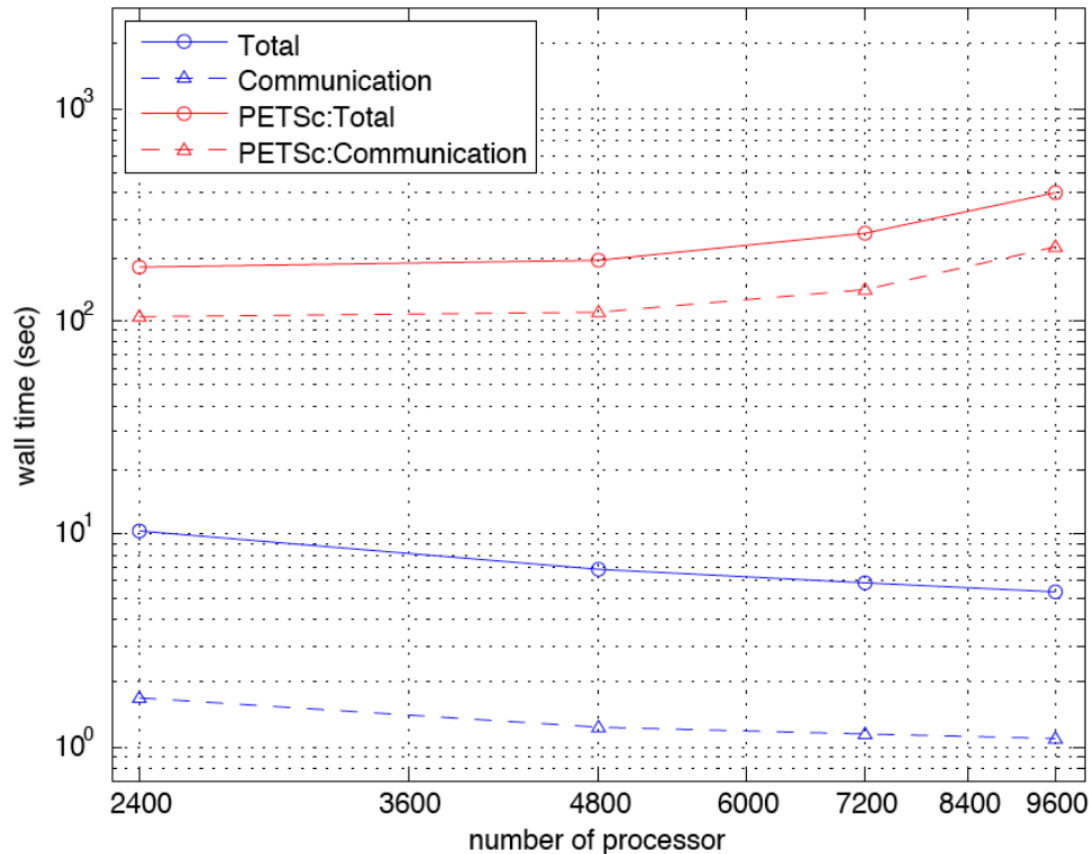
Solving the linear system

$$\begin{bmatrix} \mathbf{A}^{-1/2} \cdot \mathbf{G} \\ \mathbf{B}^{-1/2} \end{bmatrix} \cdot \mathbf{m} = \begin{bmatrix} \mathbf{A}^{-1/2} \cdot \mathbf{d} \\ 0 \end{bmatrix} \xrightarrow{\text{LSQR}} \mathbf{m} = (\alpha \mathbf{I} + \beta \mathbf{D}_1^2 + \mathbf{G}^T \cdot \mathbf{C}_d^{-1} \cdot \mathbf{G})^{-1} \cdot \mathbf{G}^T \cdot \mathbf{C}_d^{-1} \cdot \mathbf{d}$$

$$\mathbf{C}_m^{-1} = \alpha \mathbf{I} + \beta \mathbf{D}_1^2$$

# Computational aspects (petscinv)

SPQLSR vs. PETSc (KSP-LSQR)



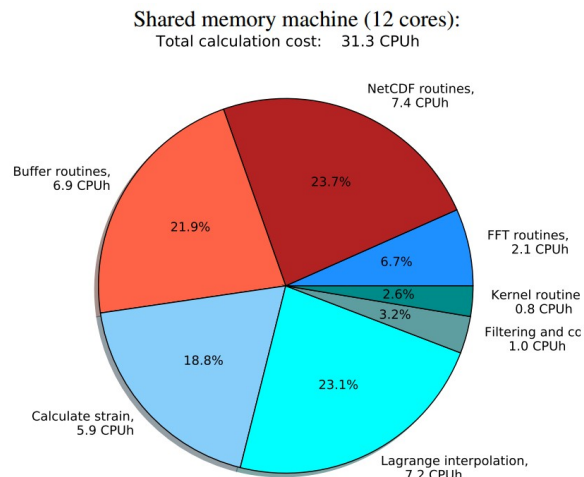
Tested our code on system matrices up to around 200 GB, takes around 10 hrs for one point in the L-curve on 12 Cores 3.45 GHz shared memory node

Tomography solver from Univ. of Wyoming

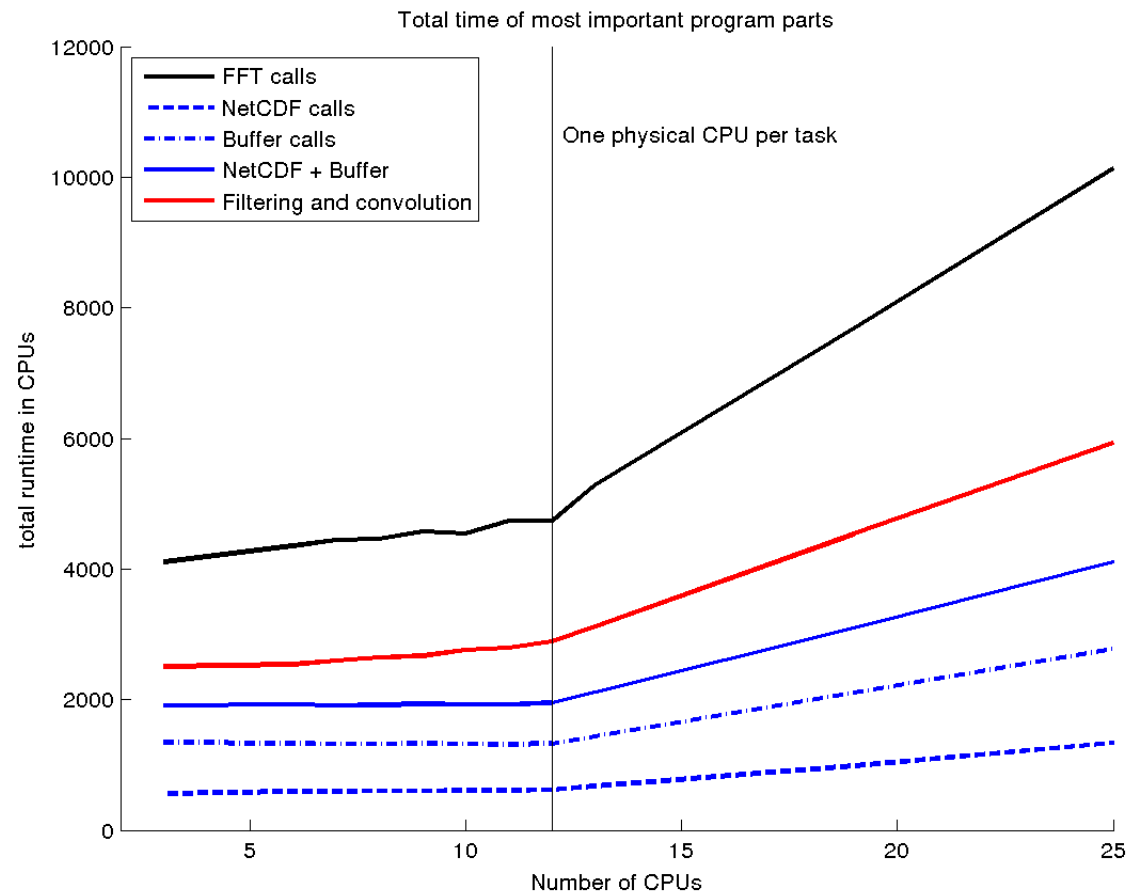
# Computational aspects (MCkernel)

Strong scaling test, fixed problem size:

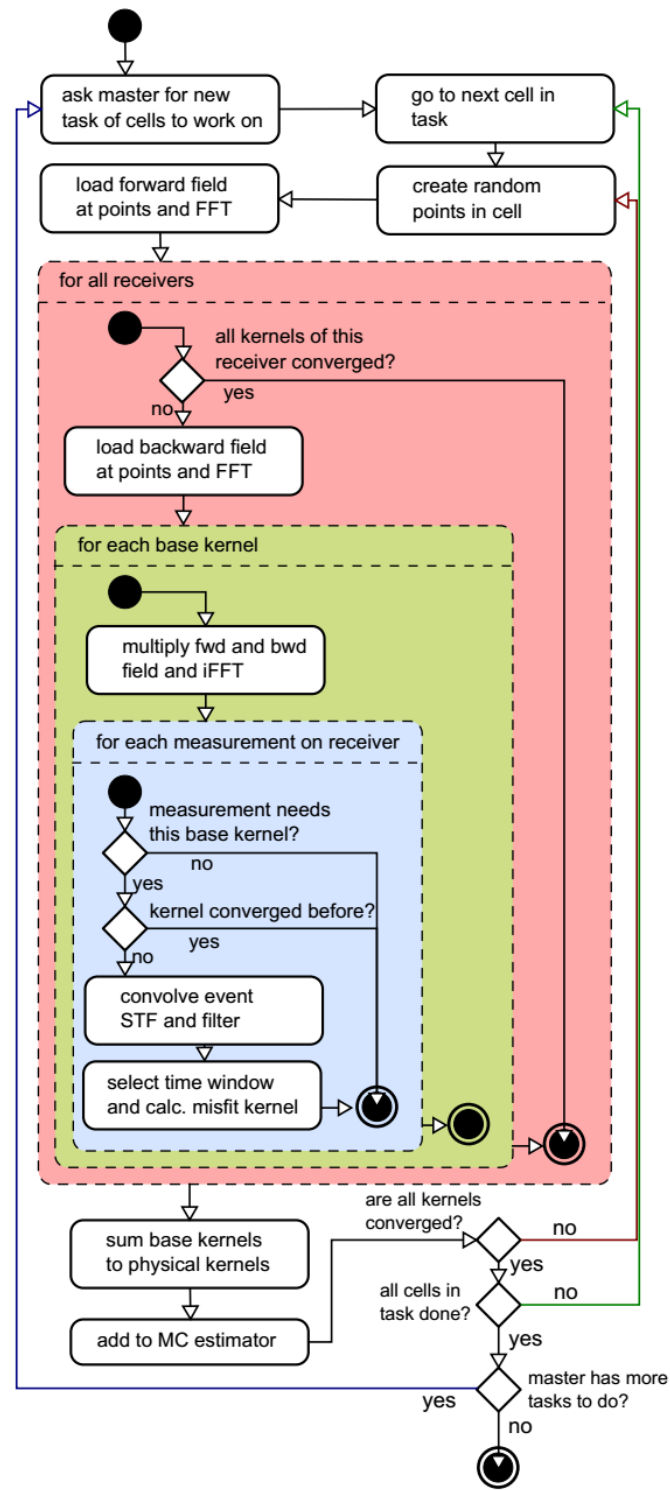
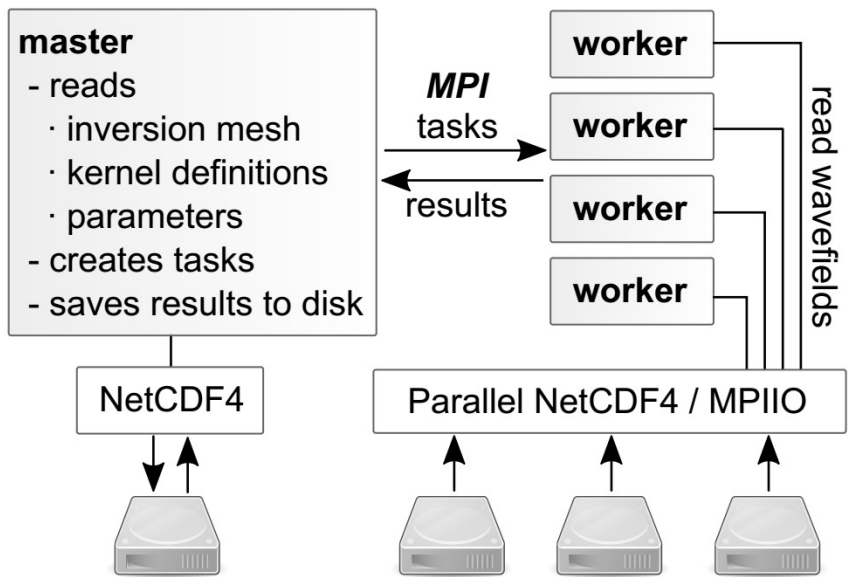
- 587 SRC-REC pairs
- 4 period bands from 10 – 30 s
- Triangular mesh with  $n=7260$



**Figure 6.** Time shares of different program parts on a small-scale run with 5 receivers and 160 kernels. Reddish colors show program parts related to IO, while blueish show CPU-intensive parts. It can be seen that the code benefits highly from a low-latency file system, as the one on the shared-memory machine. Note that the CPU frequency of the Shared memory machine was 3.45 GHz, compared to 2.2 GHz for the HPC one.

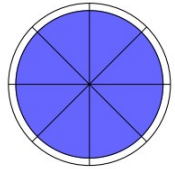


# Mckernel: How it works

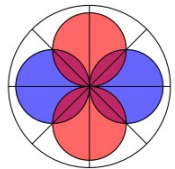


# AxiSEM, basic principle

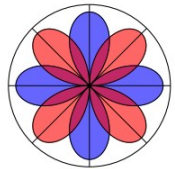
Source Decomposition:



$$\mathbf{u} = \mathbf{u}(s, z)$$

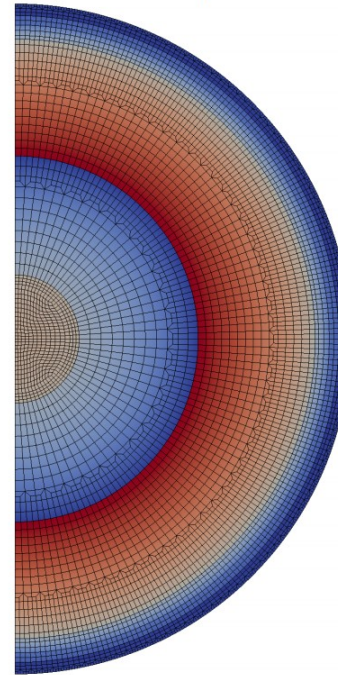


$$\mathbf{u} = \mathbf{u}(s, z) \cdot f(\sin \phi, \cos \phi)$$



$$\mathbf{u} = \mathbf{u}(s, z) \cdot f(\sin(2\phi), \cos(2\phi))$$

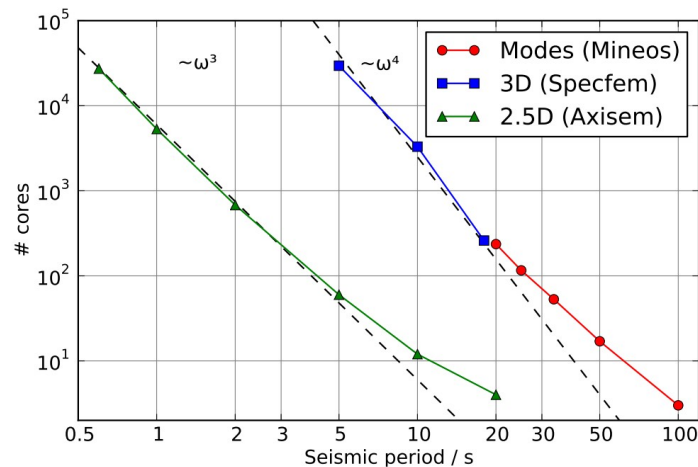
2D numerical problems:



1D models → one can use properties arising from axial symmetry (where source is located on a axis through the center of the earth) to decompose moment tensor sources in a series of multipole sources that can be solved on a 2D disk.

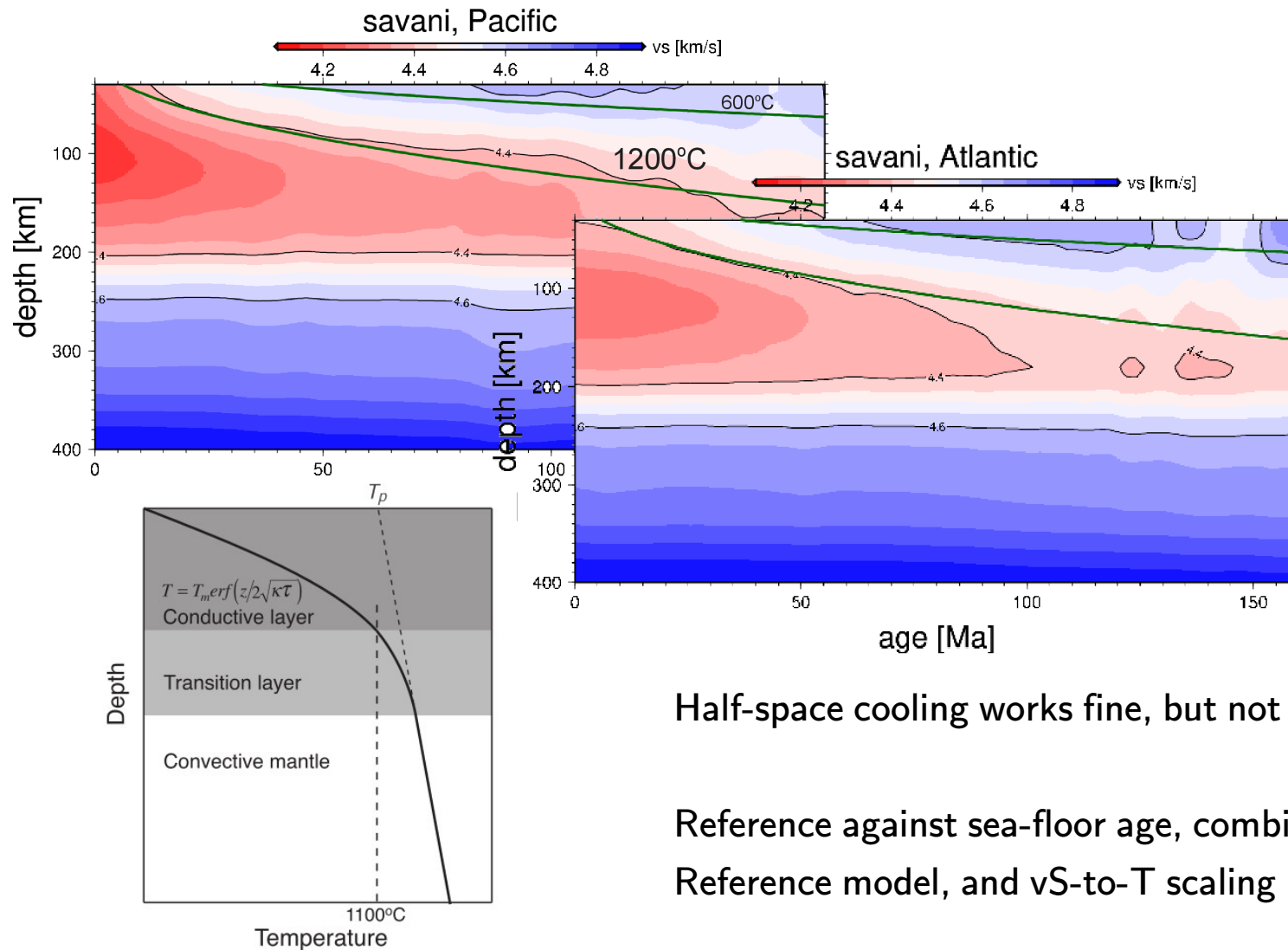
3D wavefields reconstructed from 2D solutions via Analytical relations

→ Global 1D wavefields up to 2 Hz



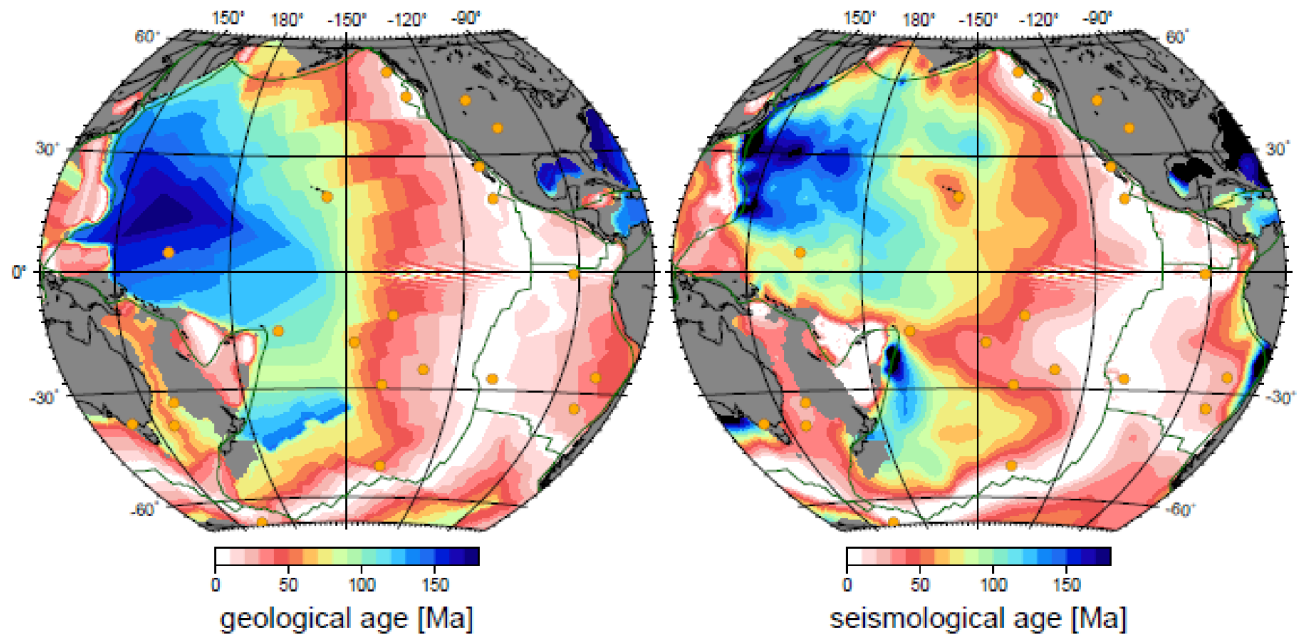


# HSC comparsion (1)



Half-space cooling works fine, but not at  $> \sim 80$  Ma

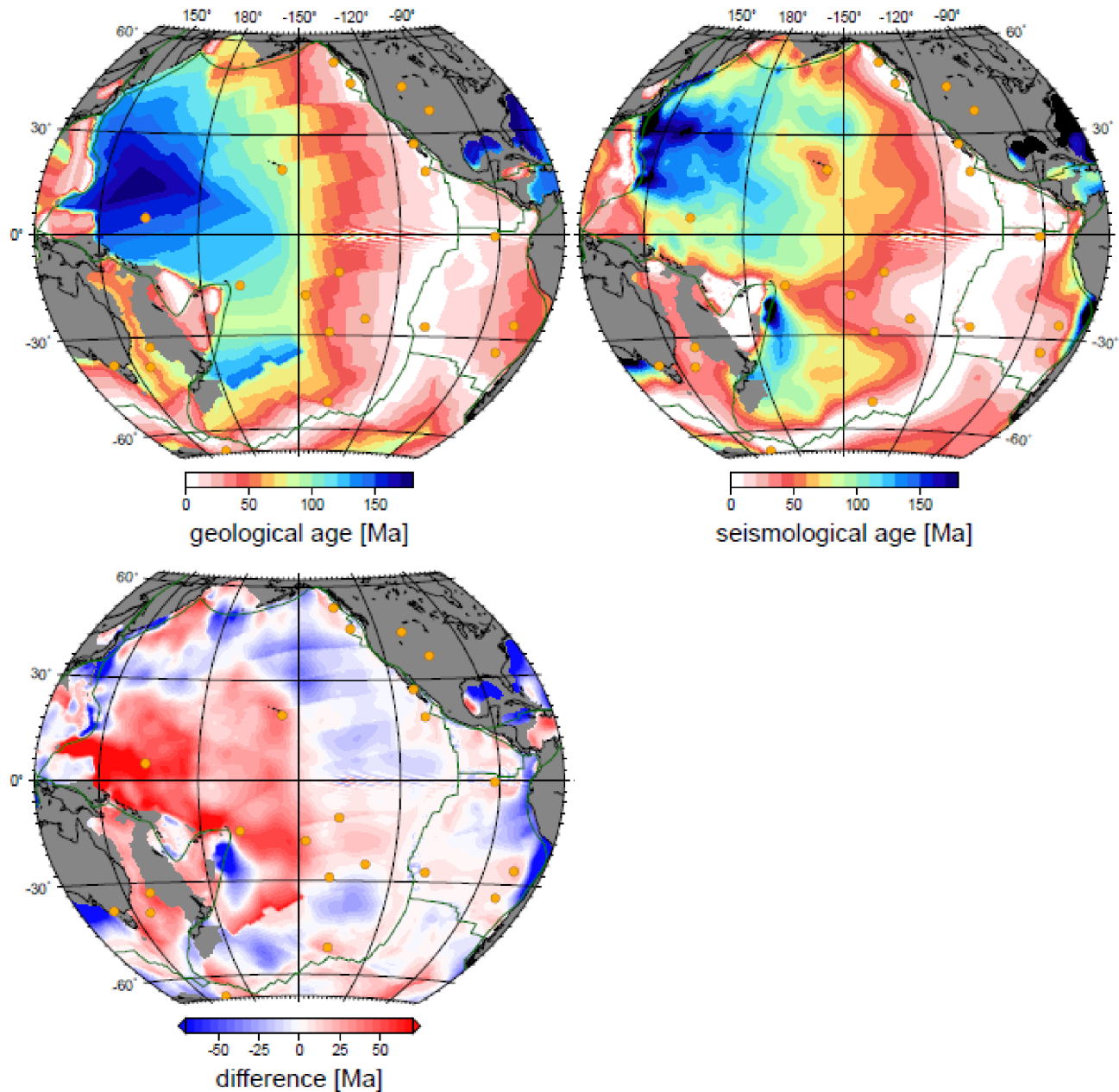
Reference against sea-floor age, combine with a thermal Reference model, and  $vS$ -to- $T$  scaling relationship



SAVANI,  
equivalent  
half-space  
cooling age,  
referenced  
to Pacific

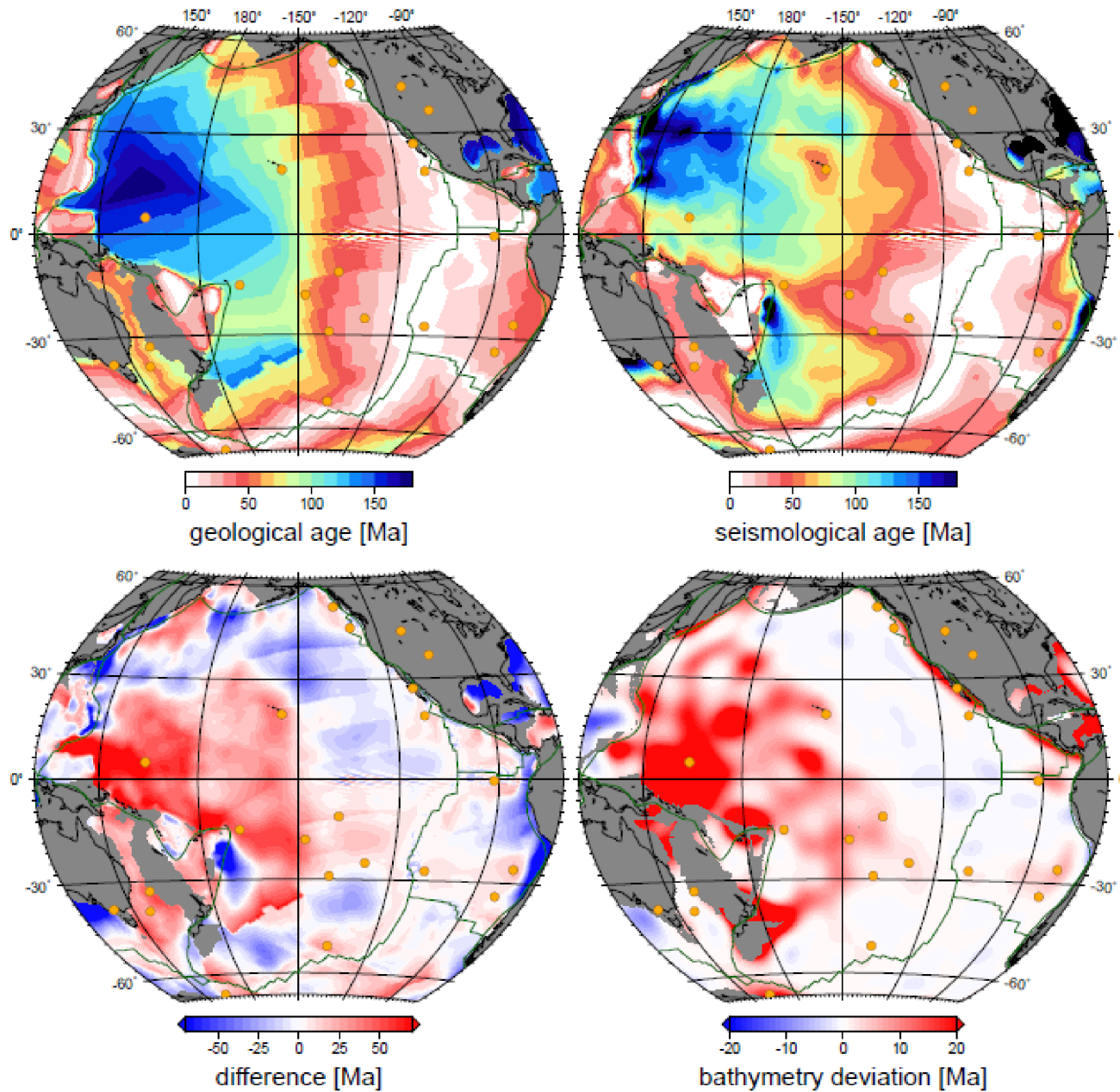
cf. Ritzwoller et al. (2004)

SAVANI,  
equivalent  
half-space  
cooling age,  
referenced  
to Pacific

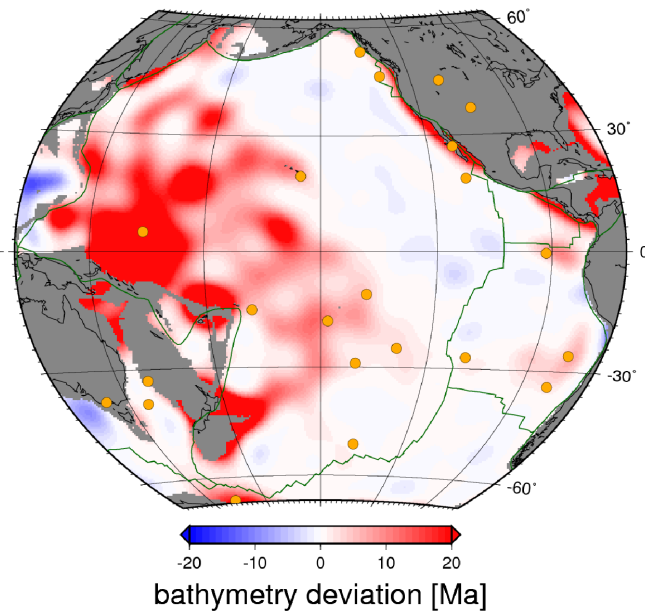
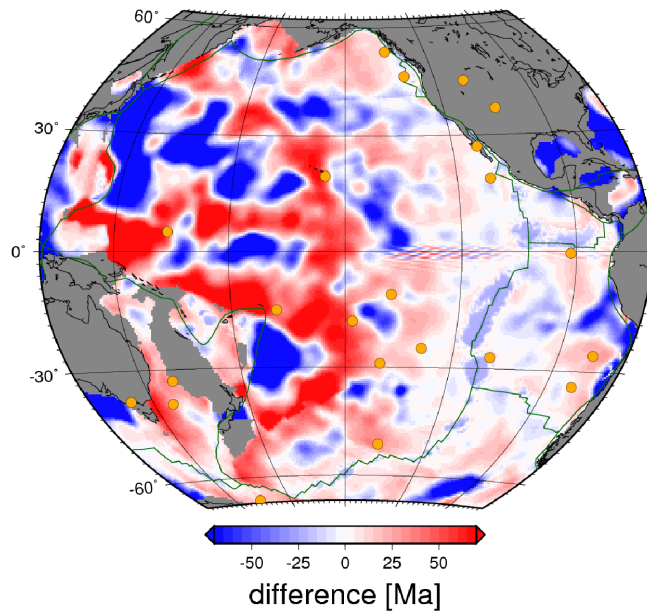
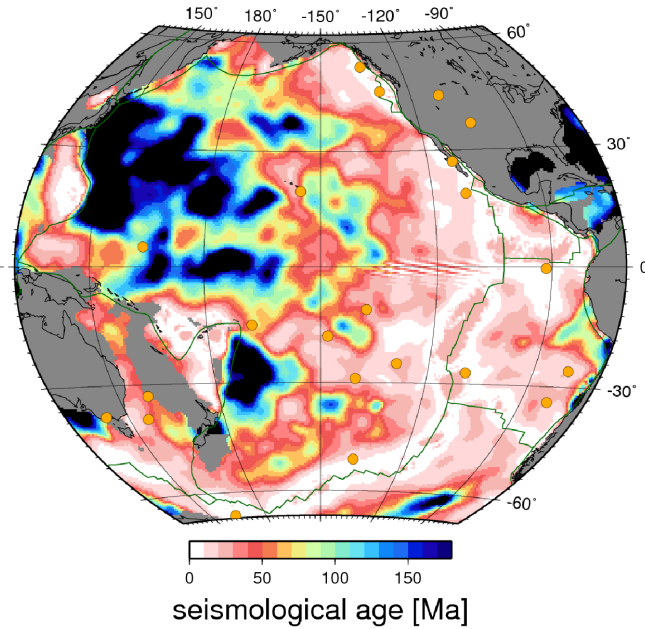
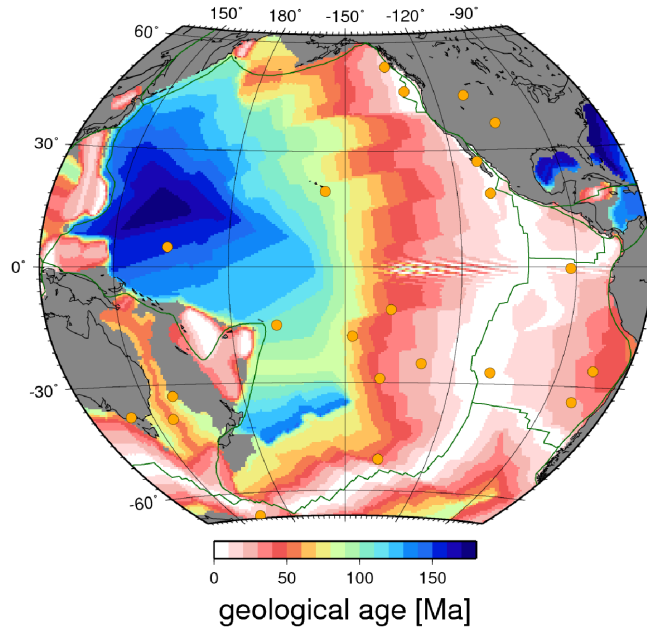


cf. Ritzwoller et al. (2004)

# SAVANI, equivalent half-space cooling age, referenced to Pacific



cf. Ritzwoller et al. (2004)



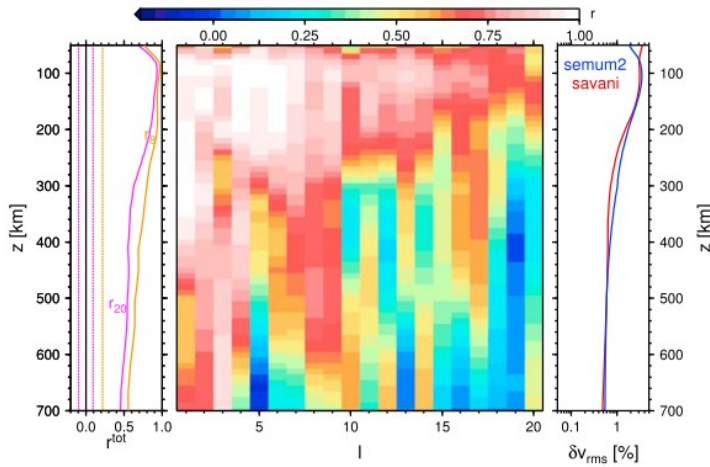
SEMum2,  
equivalent  
half-space  
cooling age,  
referenced  
to Pacific

Ritzwoller et al. (2004)

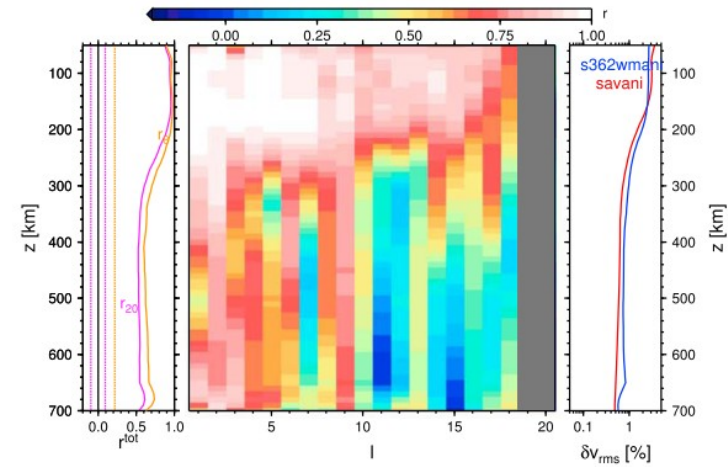


# SH cross-model comparisons

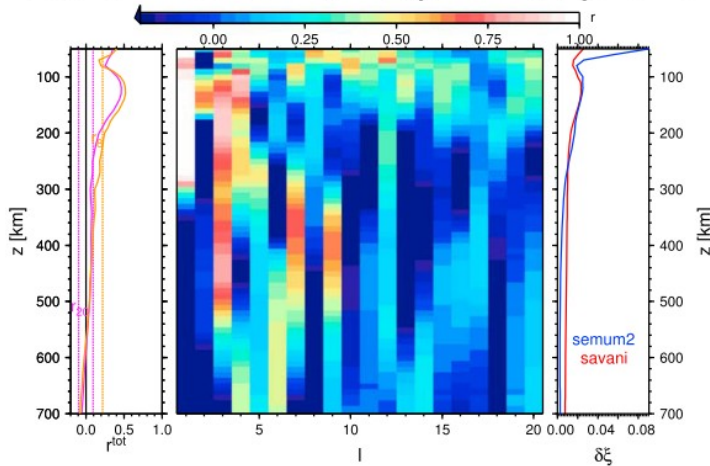
a)  $v_s$  savani - semum2,  $\langle r_{20} \rangle = 0.66$ ,  $\langle r_8 \rangle = 0.76$



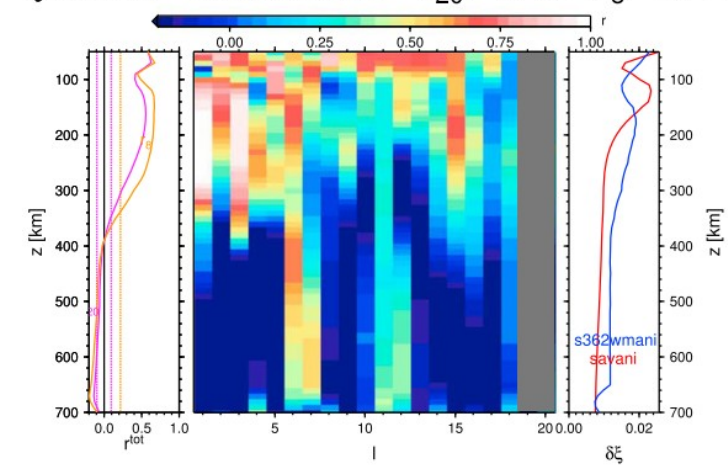
b)  $v_s$  savani - s362wmani,  $\langle r_{20} \rangle = 0.68$ ,  $\langle r_8 \rangle = 0.75$



c)  $\xi$  savani - semum2,  $\langle r_{20} \rangle = 0.12$ ,  $\langle r_8 \rangle = 0.15$

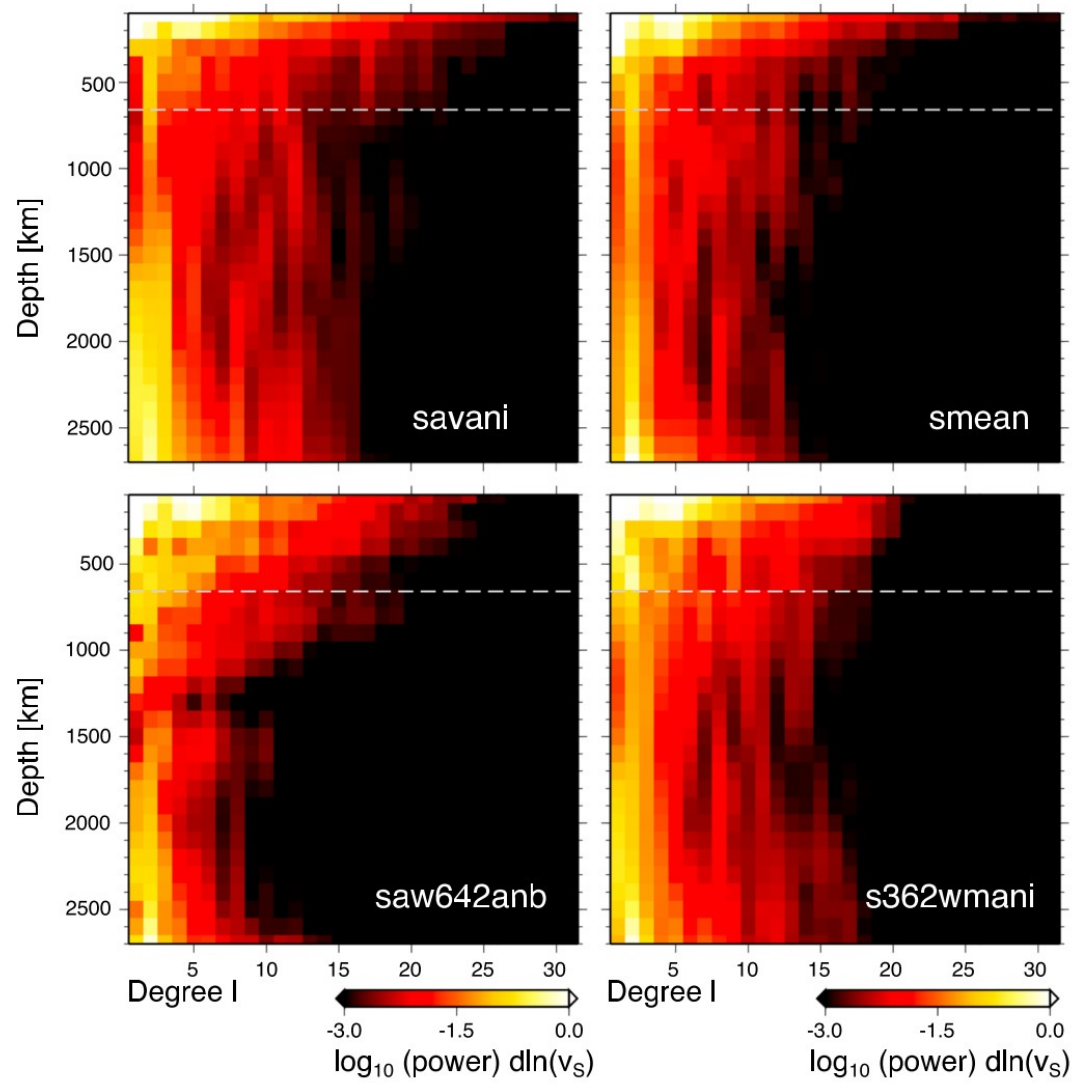


d)  $\xi$  savani - s362wmani,  $\langle r_{20} \rangle = 0.18$ ,  $\langle r_8 \rangle = 0.22$





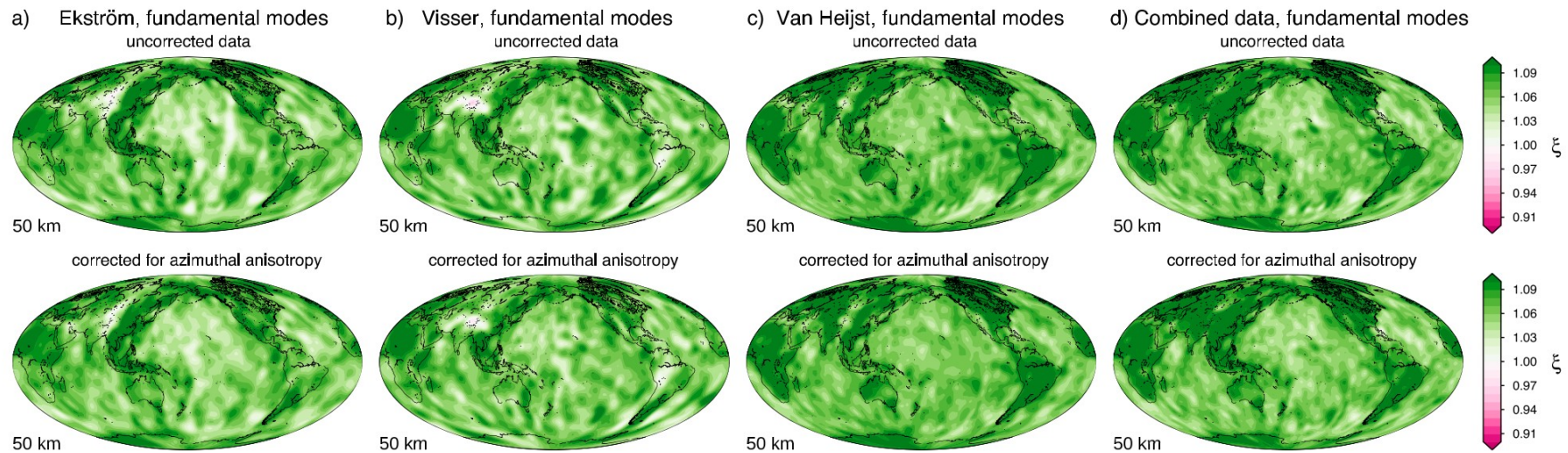
# Model power, comparison







# Azimuthal anisotropy correction



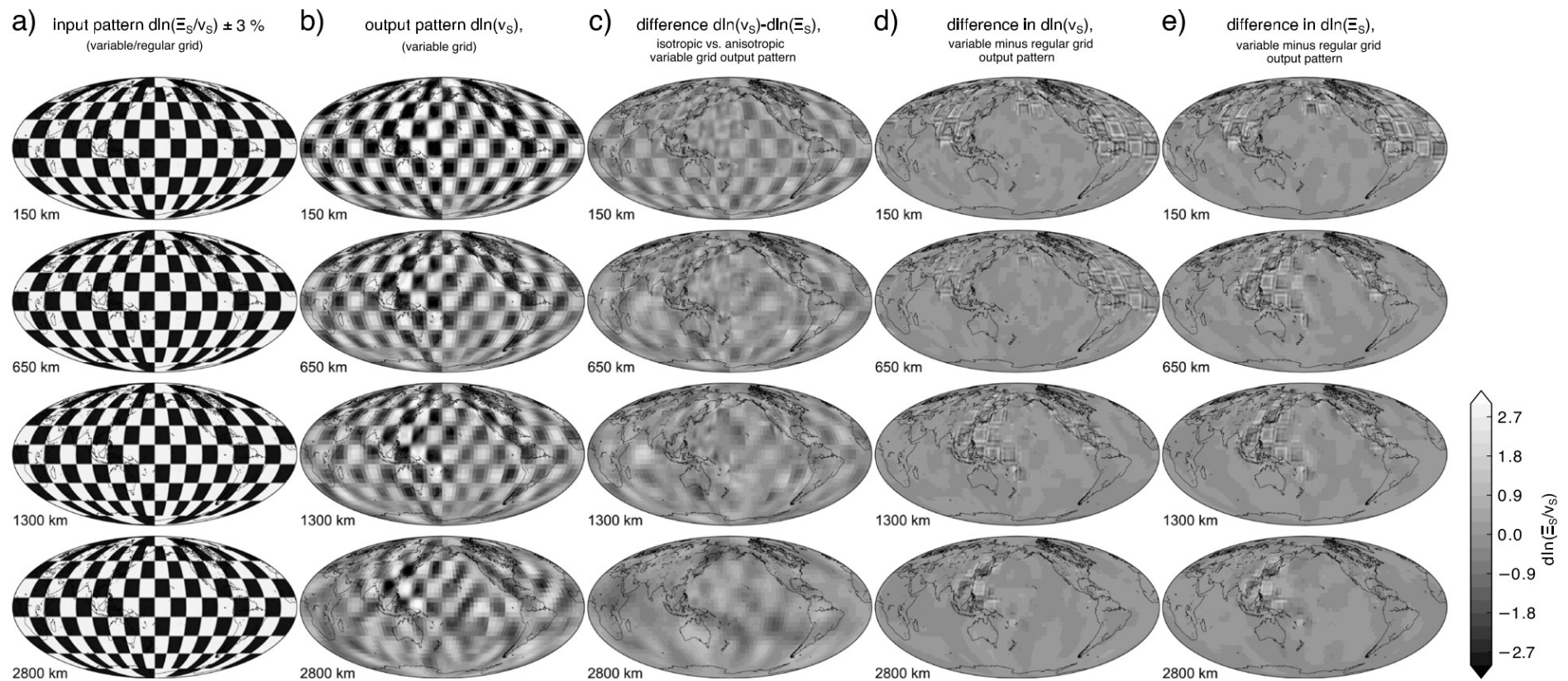
**Figure 8.** Anisotropic variations at a depth of 50 km, (a–d, top) before and (a–d, bottom) after correcting fundamental modes for azimuthal anisotropy. The corrections have profound effects [Ekström, 2011] and seem to remove a number of spurious anomalies such as the streak-like feature extending from Tasmania toward the East Pacific Rise (Ekström data set). Combining all data sets sharpens the continental signature at shallow depths.

# Checkerboard tests (Savani)

$$\mathbf{d} = \mathbf{A}\mathbf{m}; \quad \mathbf{m}^\dagger = \mathbf{A}^\dagger\mathbf{d} = \mathbf{A}^\dagger\mathbf{A}\mathbf{m} = \mathfrak{R}\mathbf{m}$$

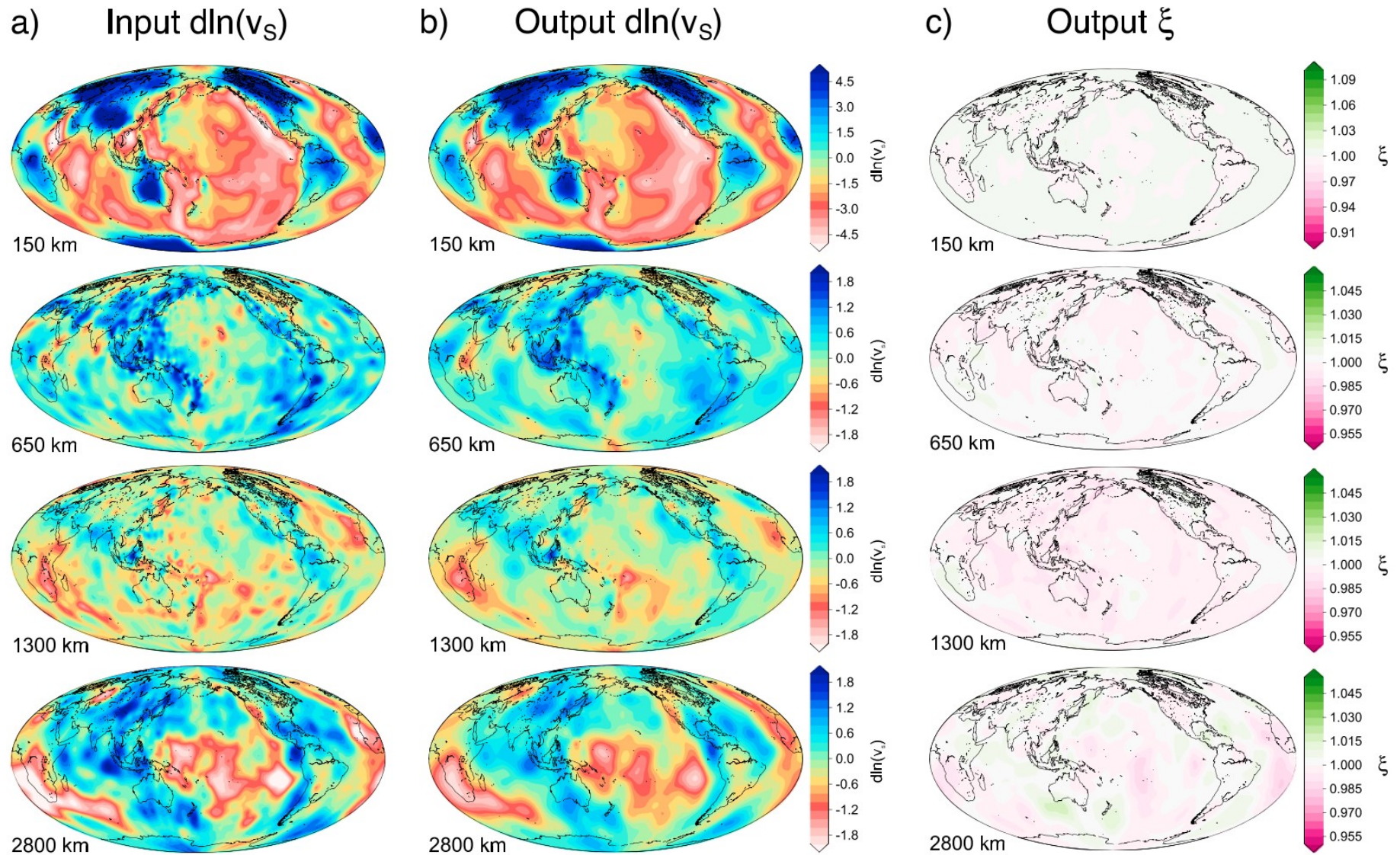
$\mathfrak{R}$  describes how the “real” Earth  $\mathbf{m}$  is project

into the tomographic model  $\mathbf{m}^\dagger$  (Menke, 1989)



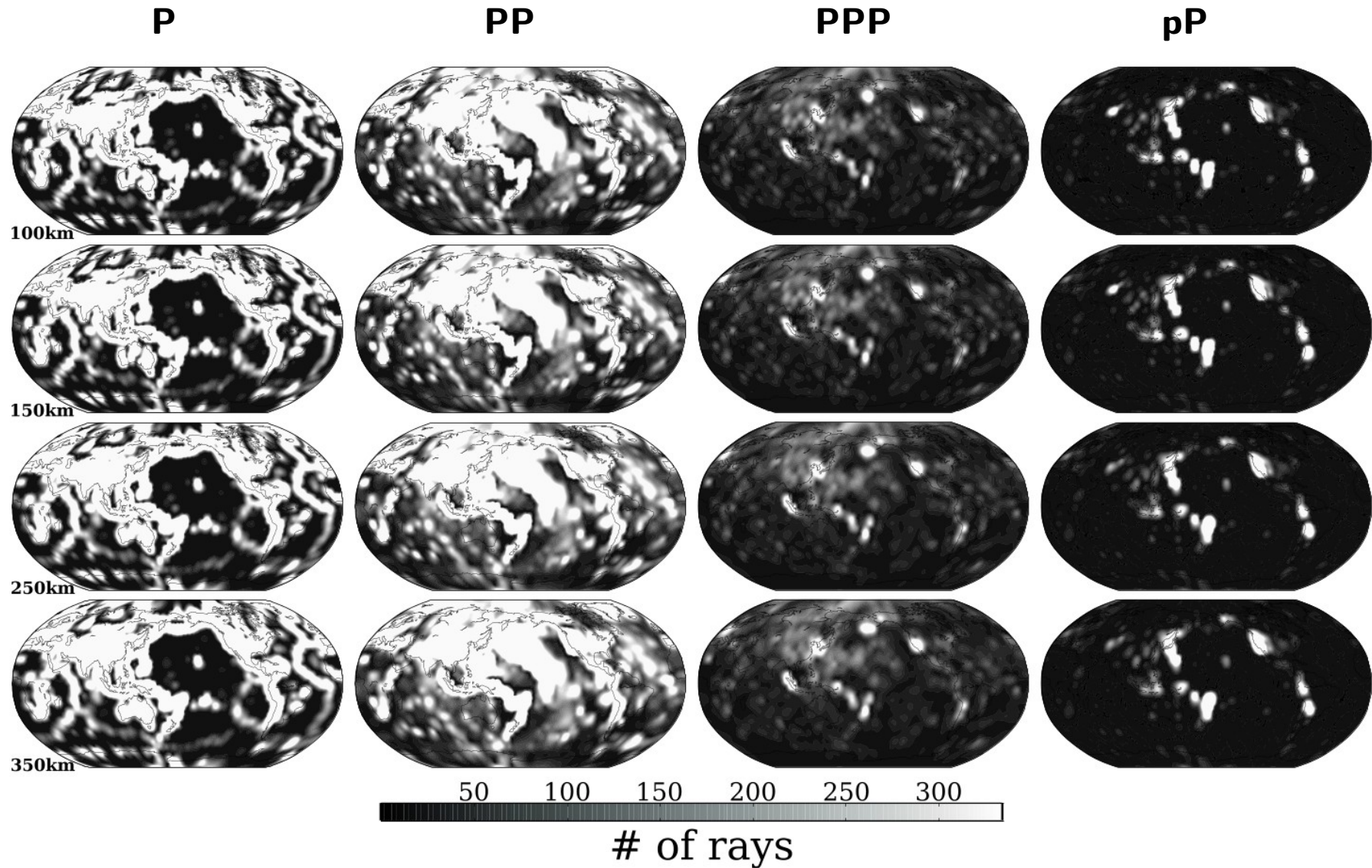


# Leakage tests (Savani)



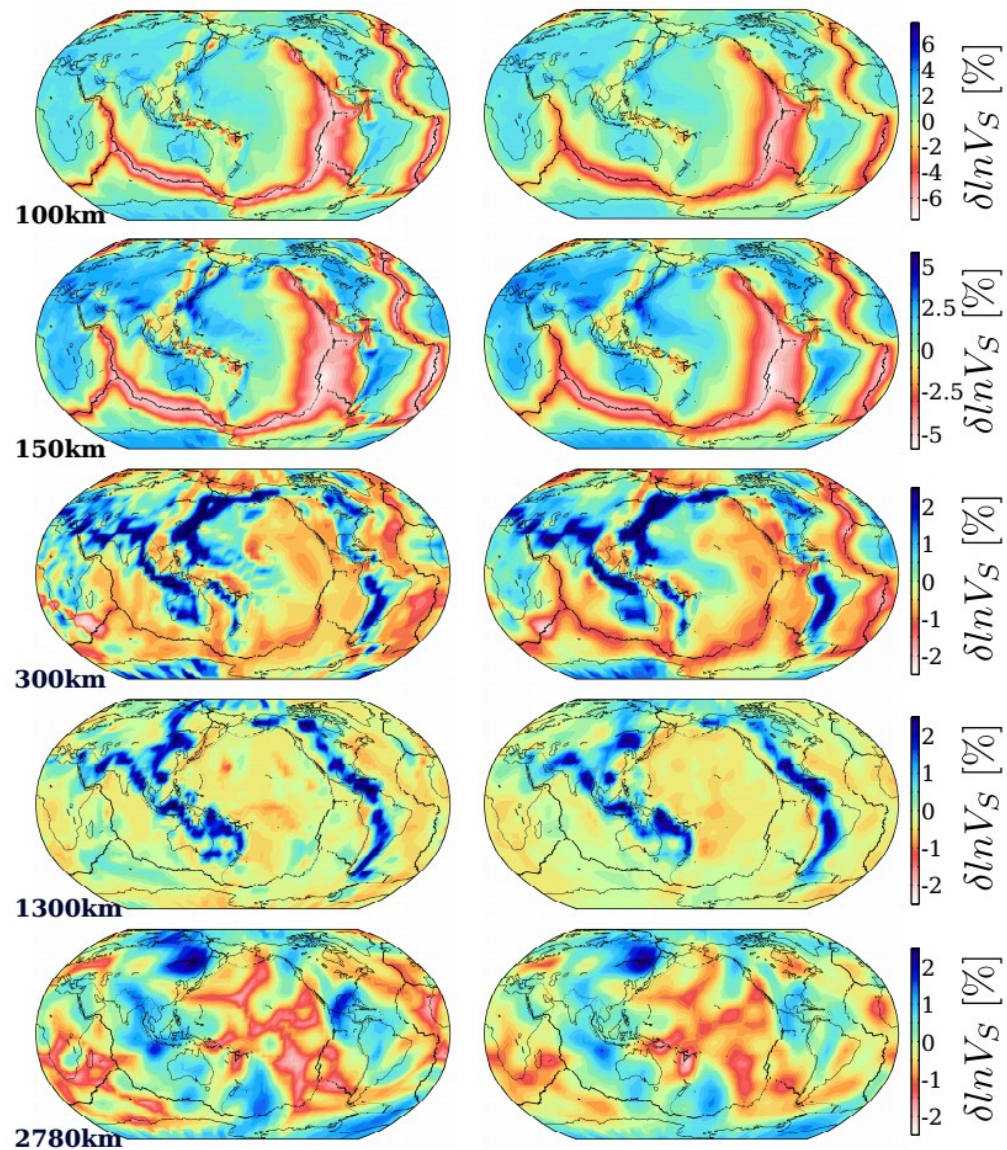


# Ray hitcounts (SPani)

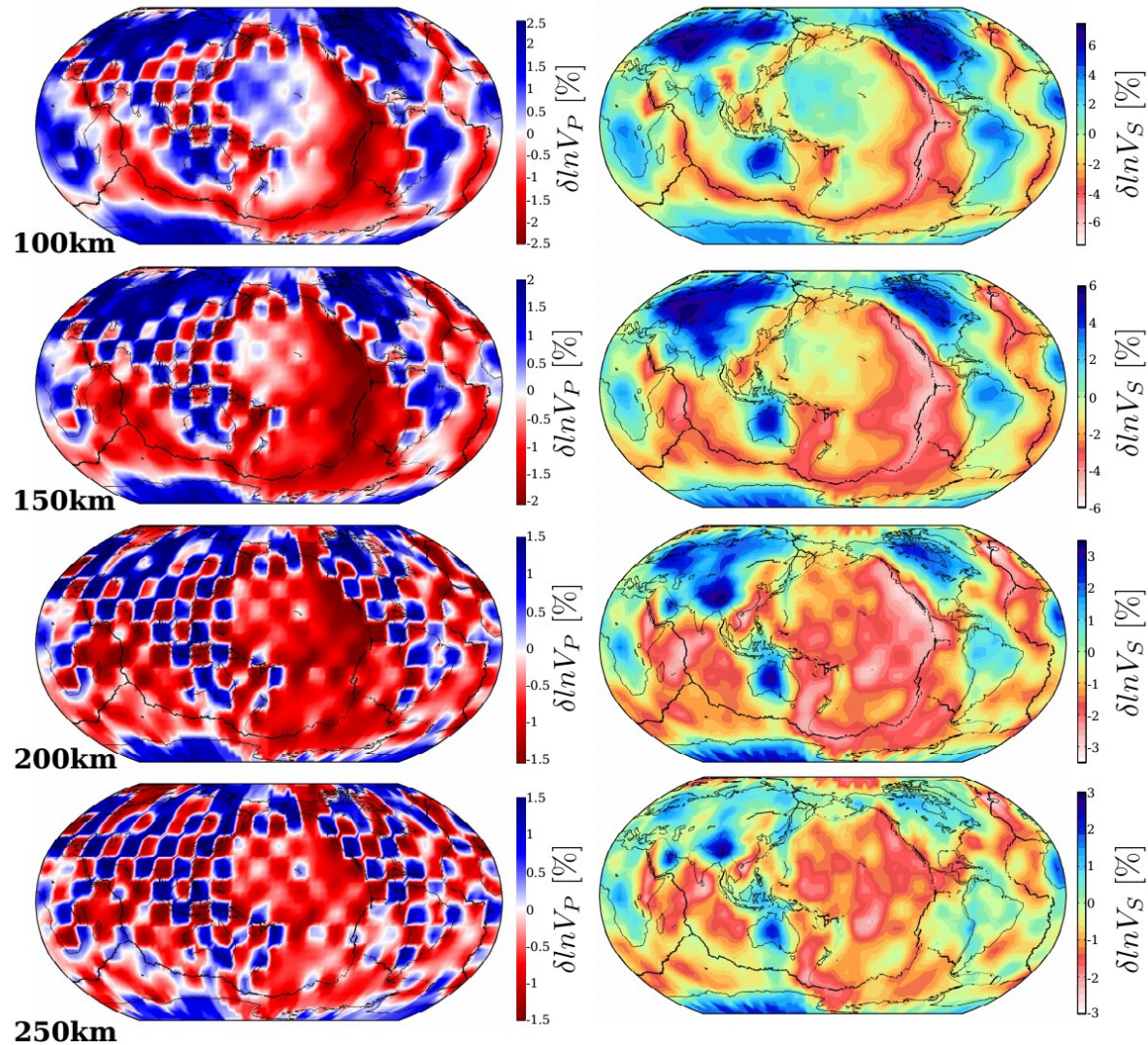




# Tomographic filtering (SPani)

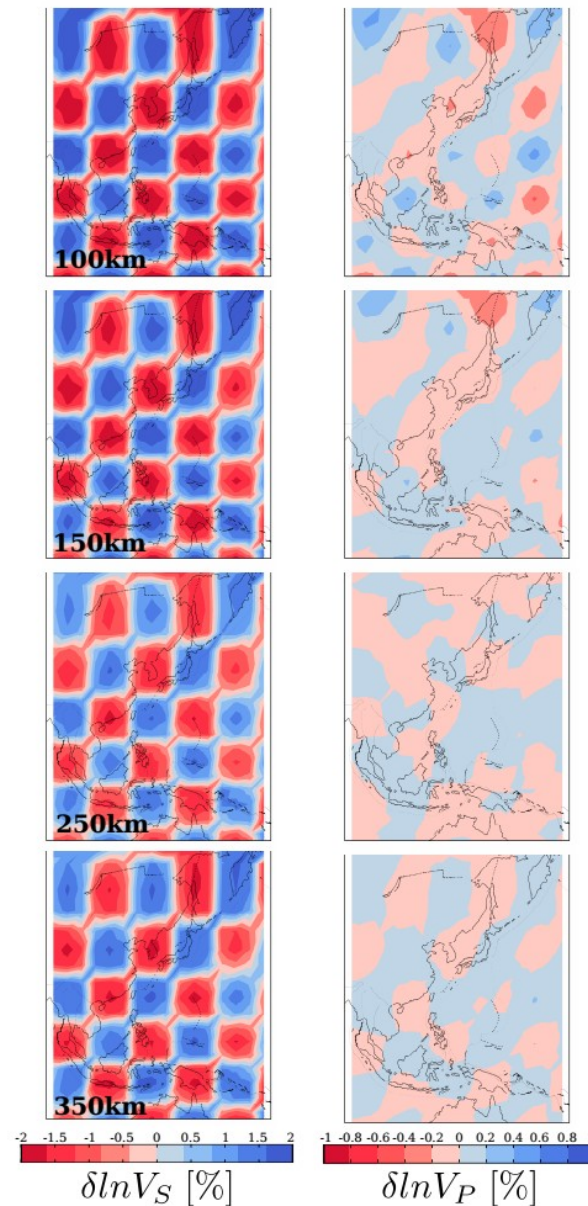


# Leakage test (SPani)





# Leakage test (SPani)

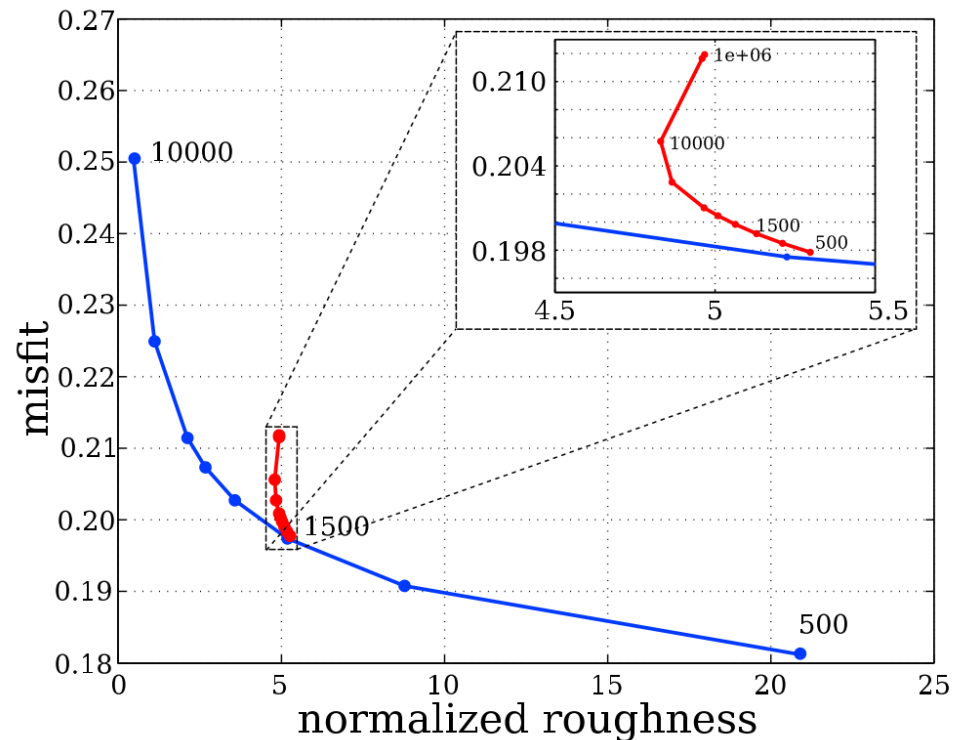


# L-curve analysis (SPani)

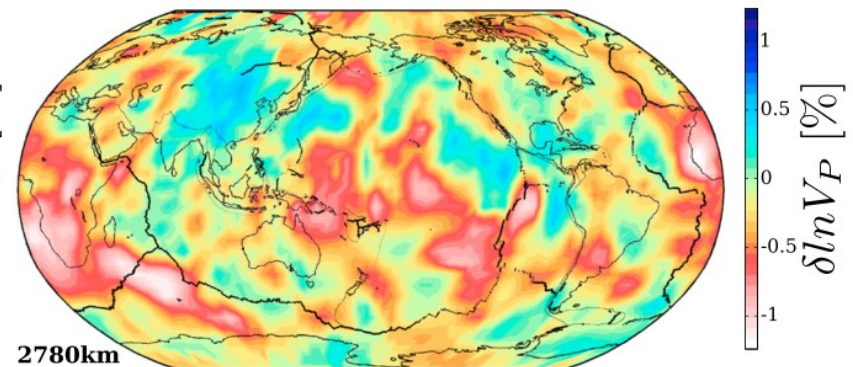
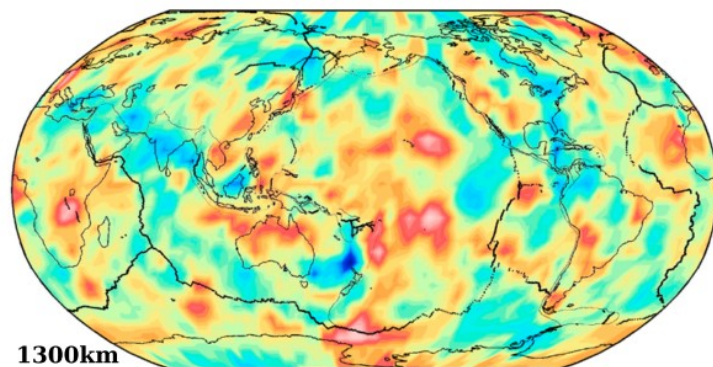
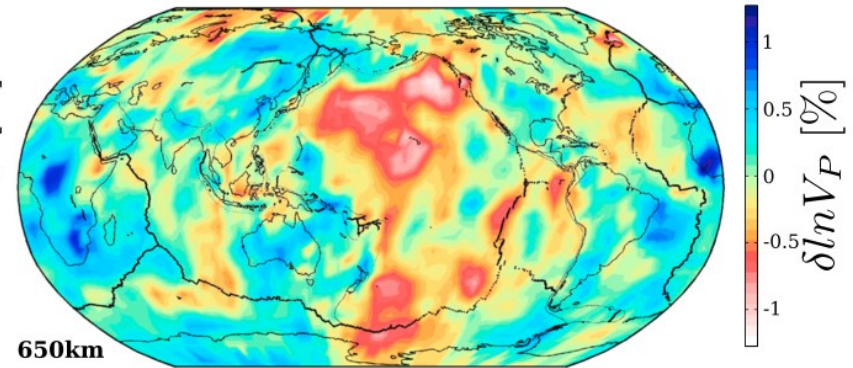
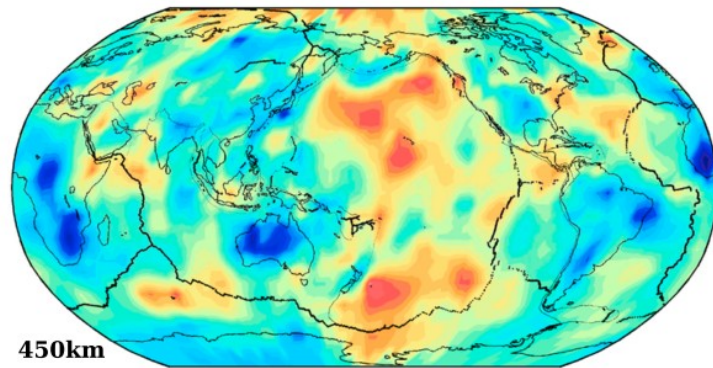
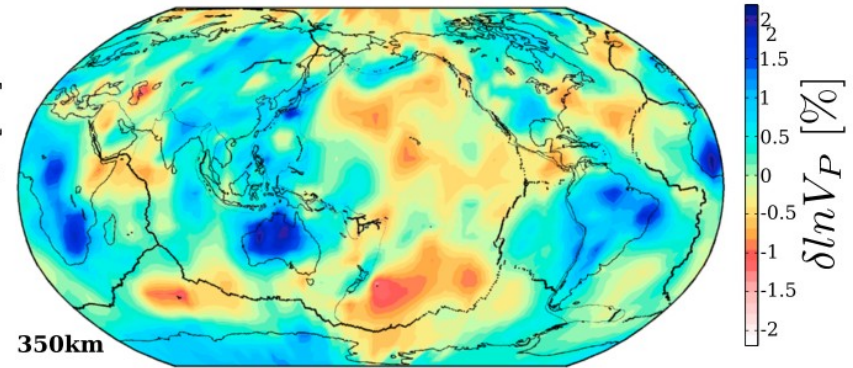
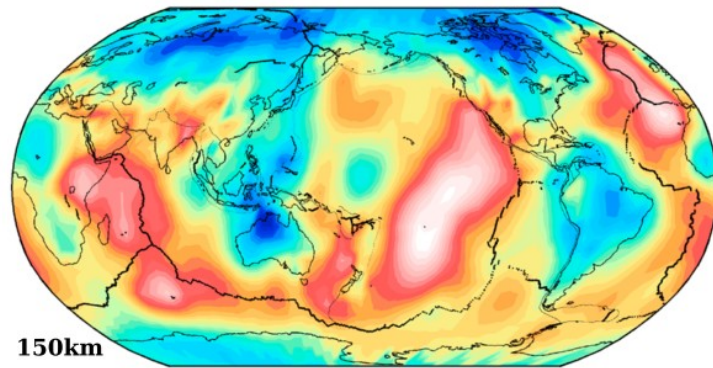
Rayleigh + P have sensitivity  
For P-wave structure, but not  
Enough to resolve it.

We “help” the inversion to  
converge to a geological  
plausible model

Need to find a balance  
between perfect scaling  
(leading poor data fit) but not  
to hide contribution from P-  
waves in the regions where  
we have large sensitivity

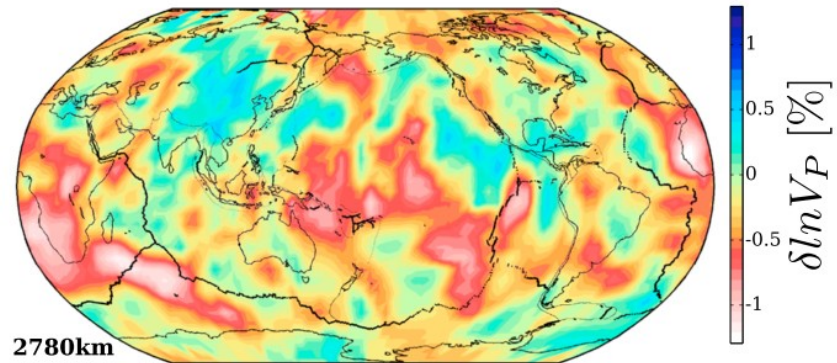
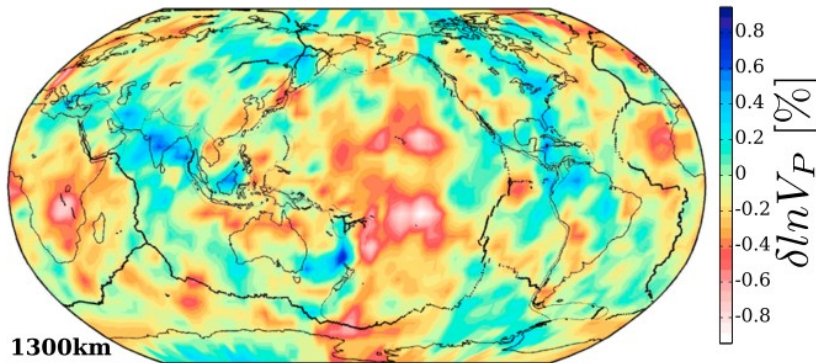
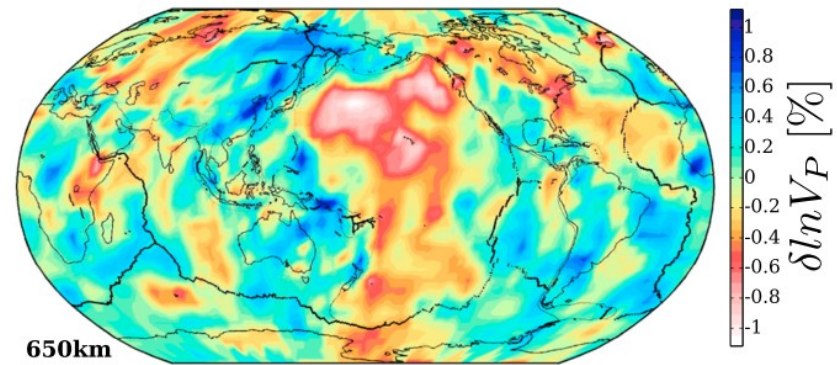
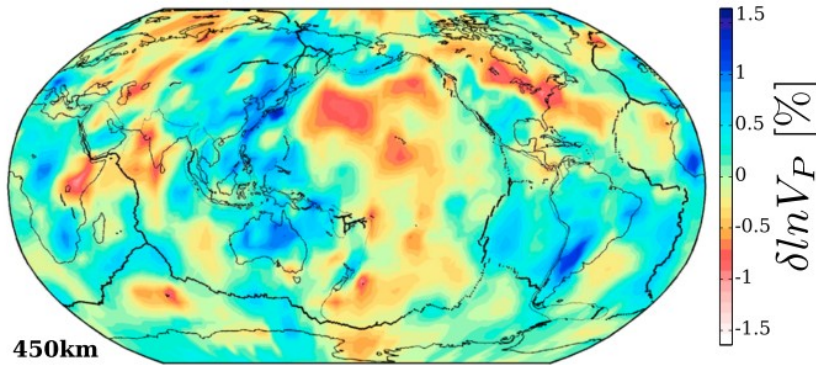
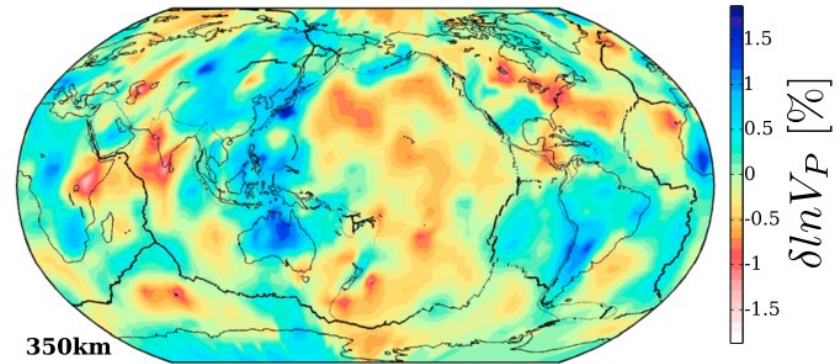
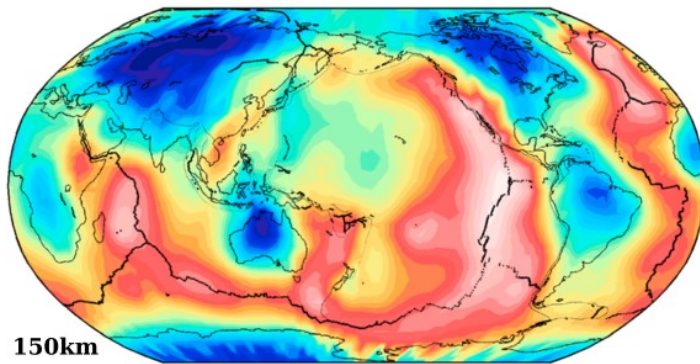


# SPani test (without scaling constraints)





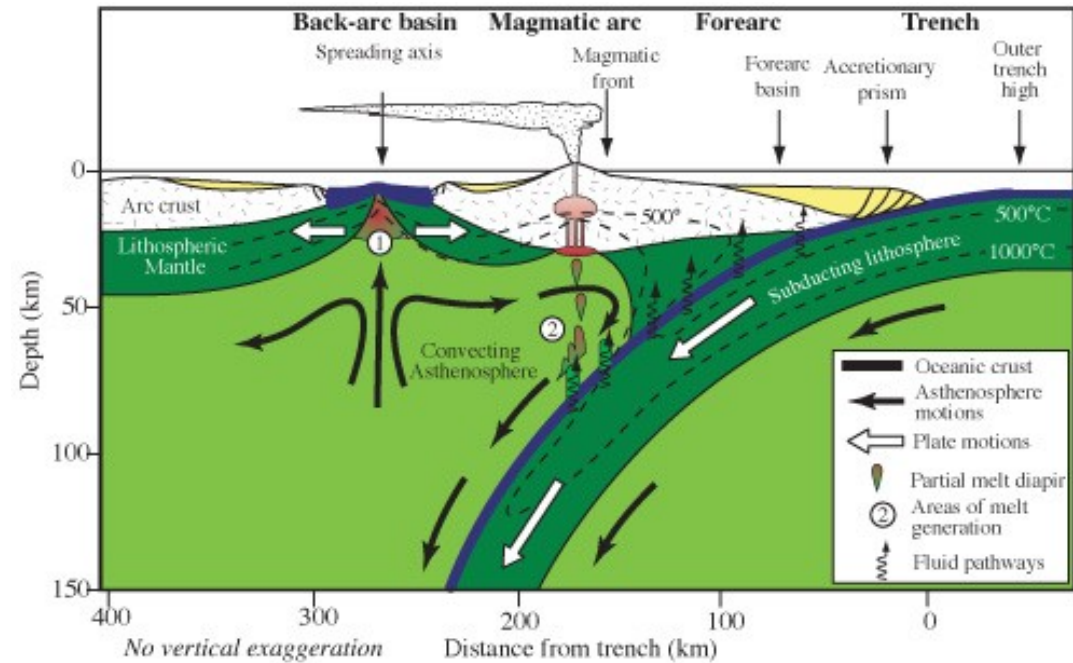
# SPani (with scaling constraints)



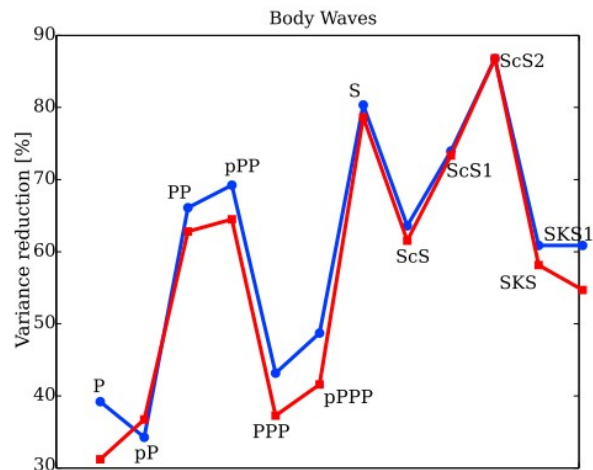
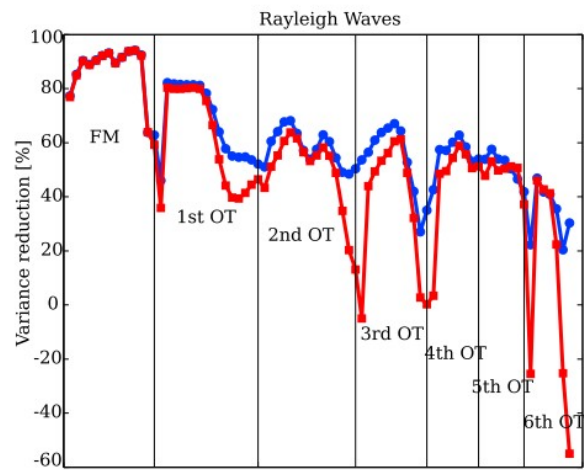
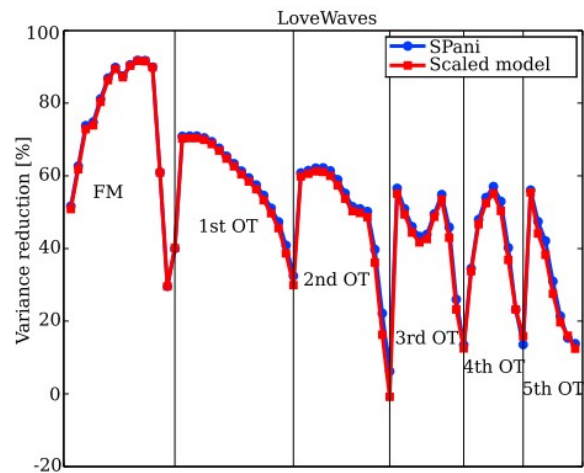


# SPani (with scaling constraints)

Partial melting occurring in the subducting slab creates a volcanic arc. The stretching of the crust caused by the upwelling of molten magma creates a basin behind the volcanic chain (e.g. Japan sea, Philippine sea)



Water transported down via the subducted slab, enables partial melting and the creation of a magmatic. Probably not degassed entirely, since we find a lot of Serpentinized sea mountains (only forms if water present). Water trapped inside the mantle in the form of hydrous minerals? Fast eastward retreatment → lowered Viscosity? → small scale convection?



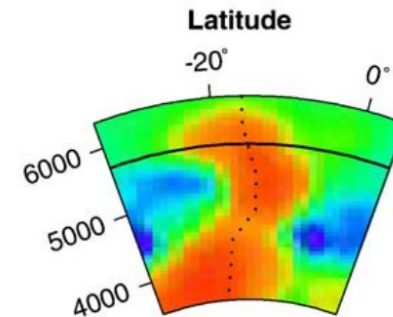
# SPani + scaled model: datafit

# A statistical condundrum

We assume normally distributed errors, define

$$\chi^2 = \sum_i \frac{|\sum_j A_{ij} m_j - d_i|^2}{\sigma_i^2}$$

And accept models with  $\chi^2 \approx N$  we now take an acceptable model, and remove one anomaly (e.g. the Tahiti plume): Small enough that Chi squared remains acceptable!



Does this mean Tahiti is not resolved?

# A statistical condundrum

We

Could be done for any anomaly! Nothing resolved?

$\chi$

However there are subsets of the data, that violate the criterion (e.g. the station under Tahiti)

And

take

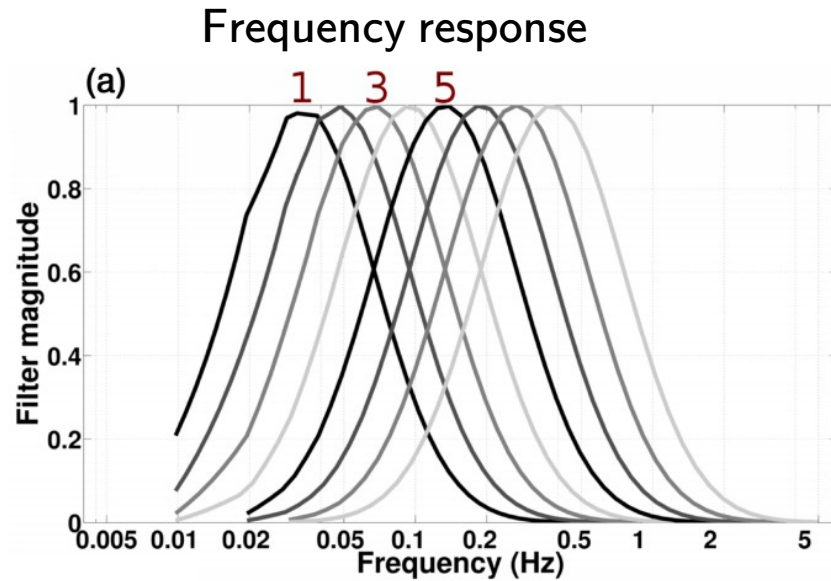
→ Create subset criterion, regionalized datasets!

anon

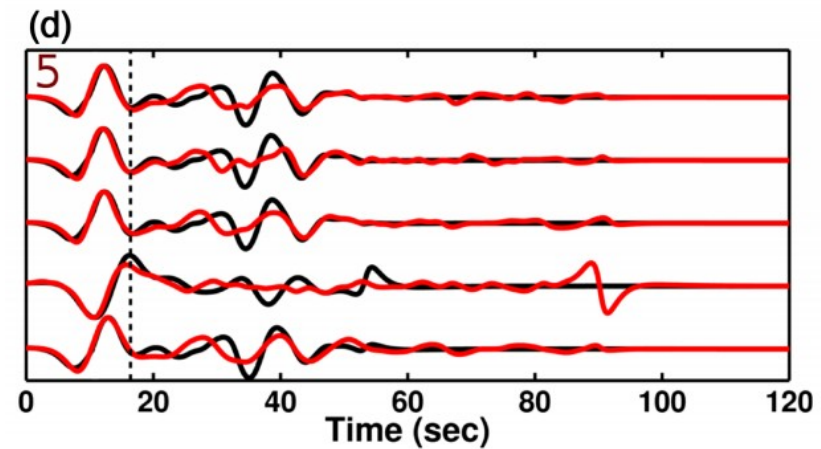
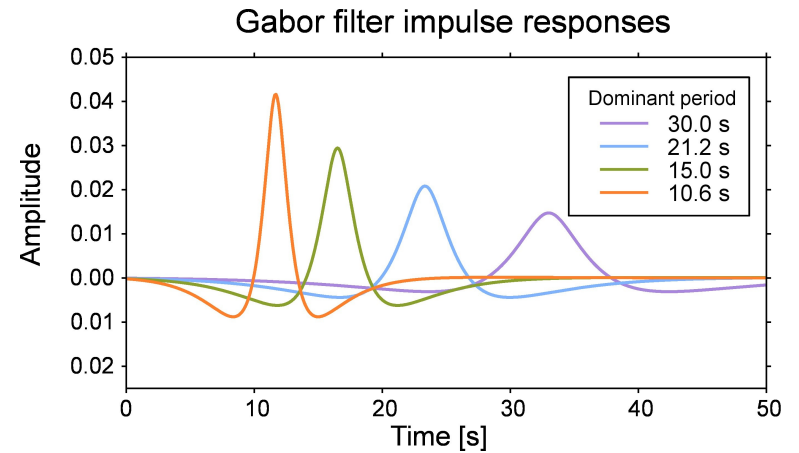
ough that Chi squared remains acceptable!

Does this mean Tahiti is not resolved?

# Measuring CC traveltimes



Constant fractional bandwidth





# Measuring phase anomalies

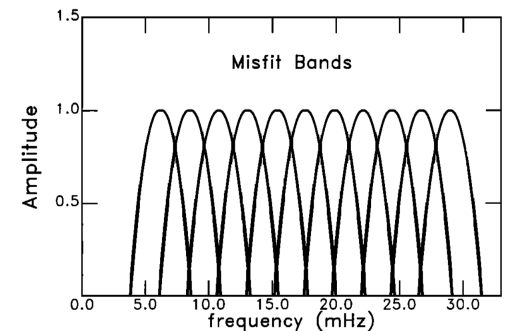
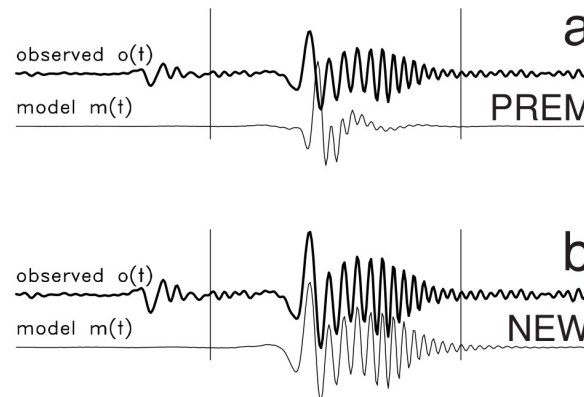
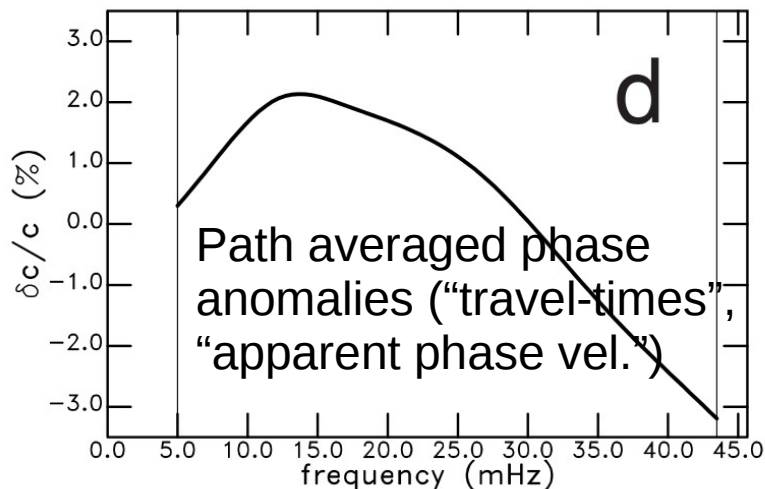
Using ray theory on a sphere (e.g. Tromp & Dahlen, 1993) we can write surface wave displacement, for a reference earth, as a function of frequency as

$$u(\omega) = A(\omega) \exp \left[ \frac{i\omega\Delta}{c(\omega)} \right] \leftarrow \text{Source Phase/Receiver Phase/Propagation Phase}$$

The observed surface wave displacement, is given as a perturbation to the reference

$$u(\omega) + \delta u(\omega) = [A(\omega) + \delta A(\omega)] \exp \left[ \frac{i\omega\Delta}{c(\omega)} + \delta\Phi(\omega) \right]$$

→ Compare seismogram with its reference synthetic in a waveform fitting procedure, that involves phase- and amplitude matching procedure in different frequencies



Dispersion properties of 1D medium iteratively modified until synthetic fits measured waveform

# Phase anomaly → 2D phase velocity maps

Phase anomaly is related to local slowness perturbation via

$$\delta\Phi(\omega) = \omega \int_0^{\Delta} \delta p(\theta(s), \phi(s); \omega) ds$$

Slowness perturbation is related to structure

$$\delta p(\theta, \phi; \omega) = \int_0^a \sum_{i=1}^I K_i(r; \omega) \delta\pi_i(r, \theta, \phi) dr$$

Phase slowness  $p$  is expanded in some basis functions, e.g.: splines  
Pixels, Spherical Harmonics, or ...

$K_i$  can be frequency dependent  
frechet derivatives or 1D-sensitivity functions

**2 options:** Inversion can be done using a 3D parameterization in one step or via the detour of first making 2D phase velocity “maps” or regionalized dispersion curves

$$\delta\Phi(\omega) = \omega \int_0^{\Delta} \int_0^a \sum_{i=1}^I K_i(r; \omega) \delta\pi_i[r, \theta(s), \phi(s)] dr ds$$

# Savani: modeling details (1)

- Horizontally: Local blocks, either constant  $3^\circ$  or with size depending on coverage ( $1.25^\circ$ - $5.0^\circ$ ); Vertically: 12/27 layers
- Physical parameterization: Radial anisotropy in  $V_{SH}$  and  $V_{SV}$  or  $V_{S_{Voigt}}$  and  $\chi$
- Weighting: downweighting outliers and upweighting for differential number of measurements
- Regularization: Roughness minimization  $x_i - \sum_{l=1}^N x_l = 0$
- Reference model: PREM

## Savani: modeling details (2)

Surface wave phase velocity perturbation  $\delta c(\omega)$  related to selected inversion parameter  $\delta\pi_i(r, \theta, \phi)$ , using

$$\frac{\delta c_j(\omega)\Delta}{\omega} = \sum_{i=1}^I \sum_{k=1}^K x_{ik} \int_0^{\Delta_j} \int_0^a K_i(r; \omega) f_k(r, \theta, \phi) dr ds \quad (1)$$

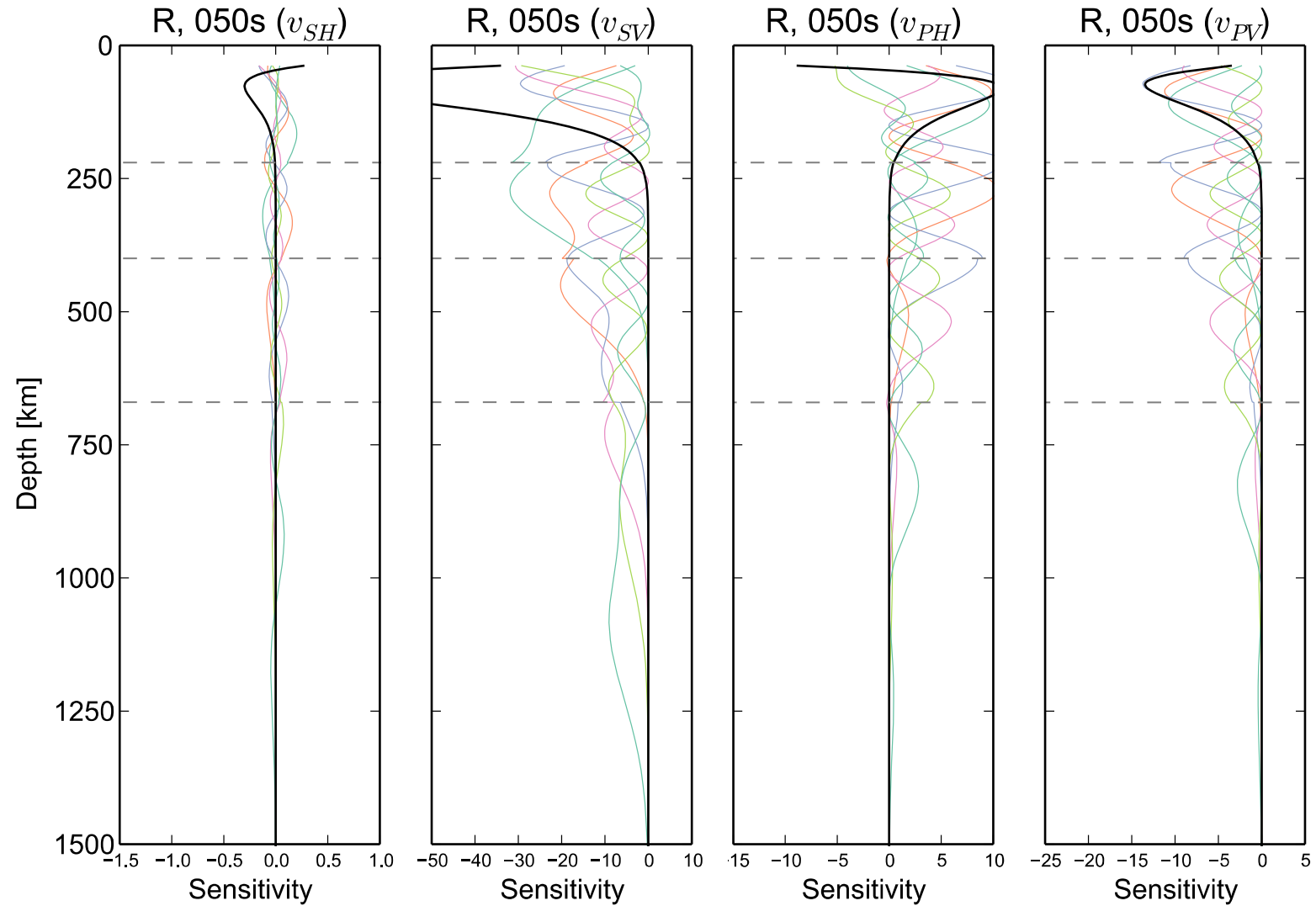
For body waves we get the similar expression

$$\delta T = - \sum_{k=1}^K x_{ik} \int_0^{\Delta_j} K_i(r; \omega) f_k(r, \theta, \phi) ds \quad (2)$$

Leads to a linear system to be solved in a least-squares sense

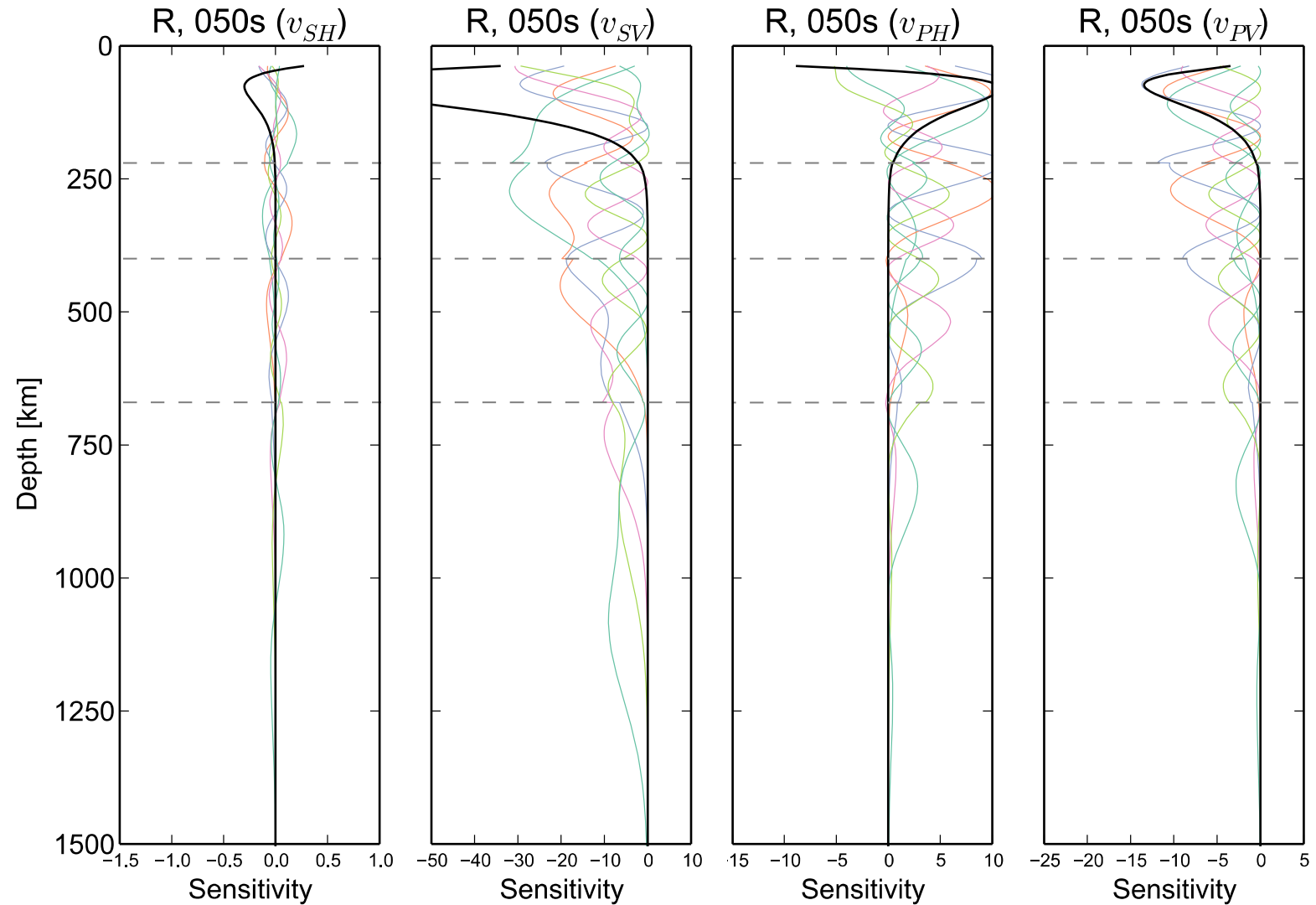
$$\mathbf{x} = (\mathbf{A}^T \cdot \mathbf{A})^{-1} \cdot \mathbf{A} \cdot \mathbf{d} \quad (3)$$

# Sensitivity (2)



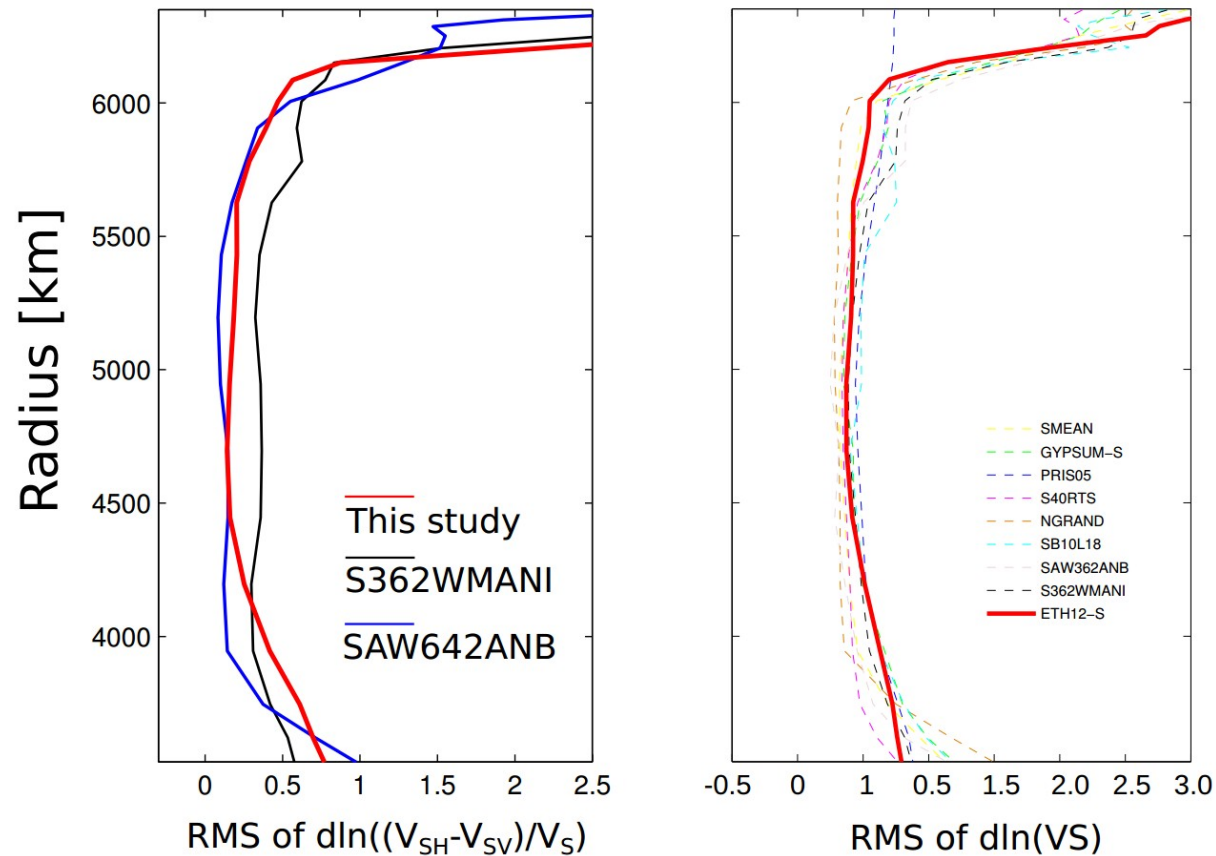


# Sensitivity (1)



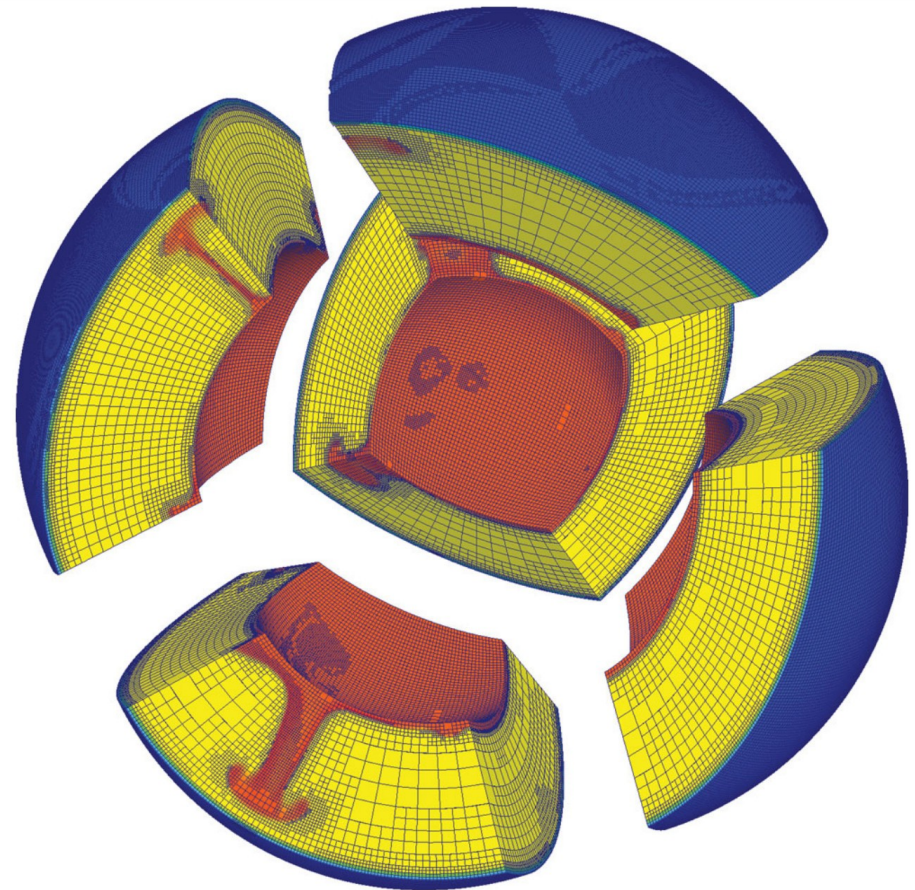
# RMS comparison

RMS profiles of our global model compared to previously published global models of radial anisotropy



# (1) Introduction: Tomography approach

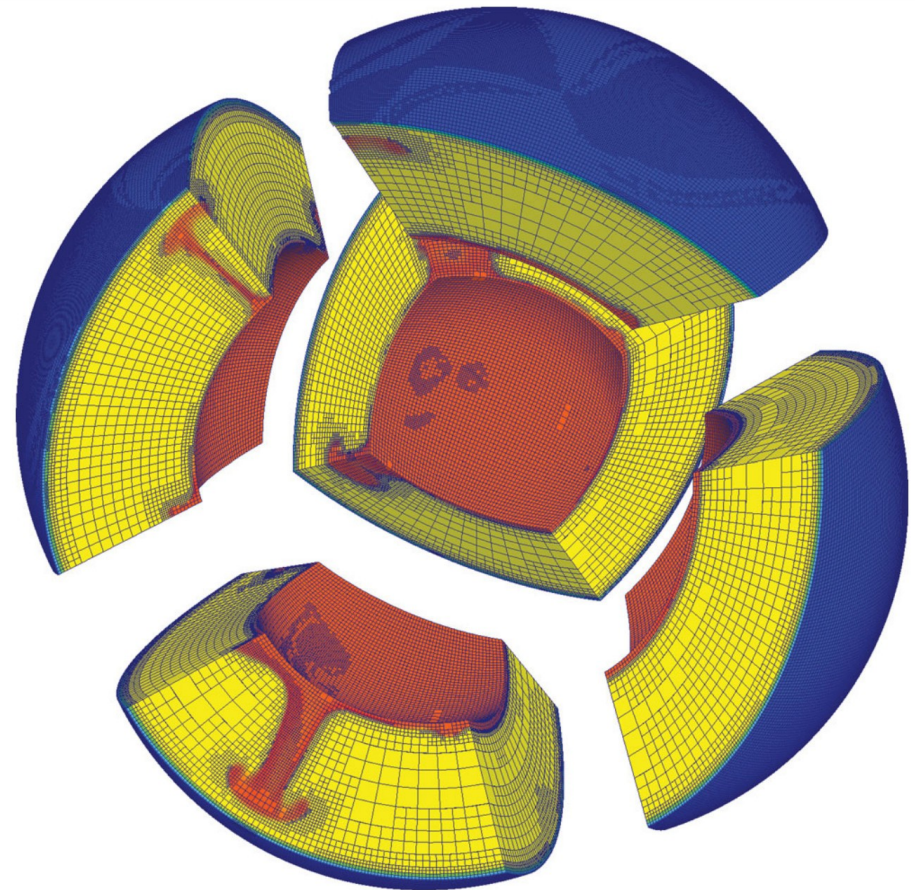
Imaging multi-scale structure and dynamics



# (1) Introduction: Tomography approach

Imaging multi-scale structure and dynamics

- Global mantle convection embraces multiple, coupled scales

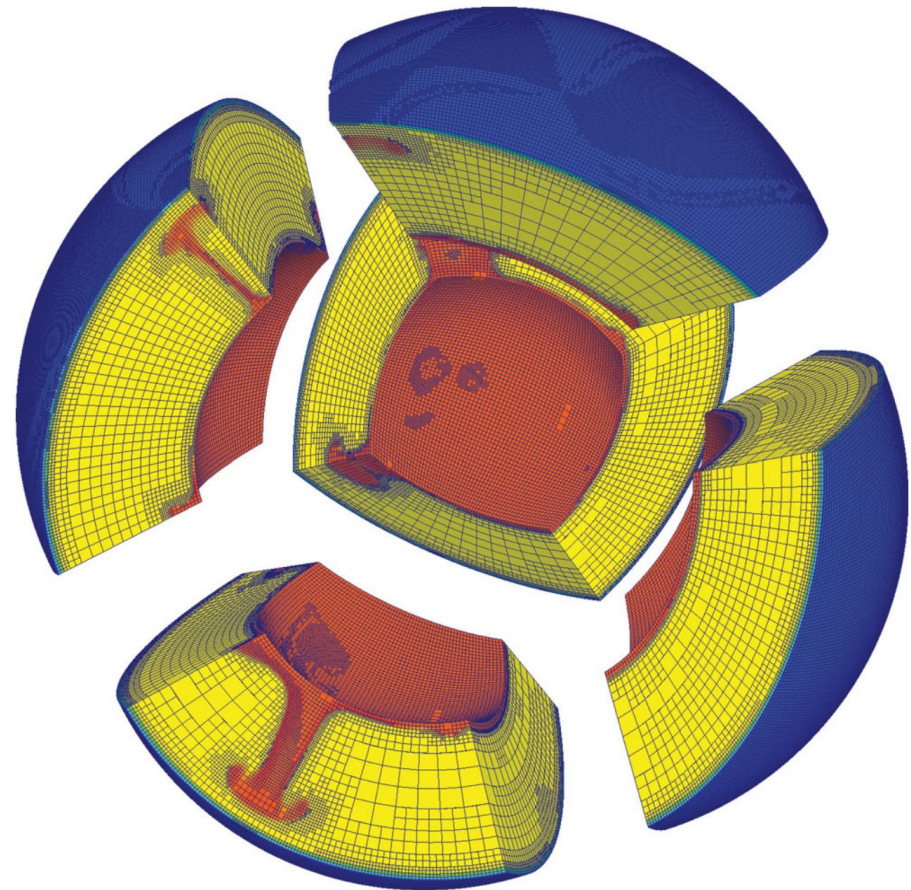




# (1) Introduction: Tomography approach

Imaging multi-scale structure and dynamics

- Global mantle convection embraces multiple, coupled scales
- Convection models optimally account for them (see, e.g. *Burstedde et al. 2013*)

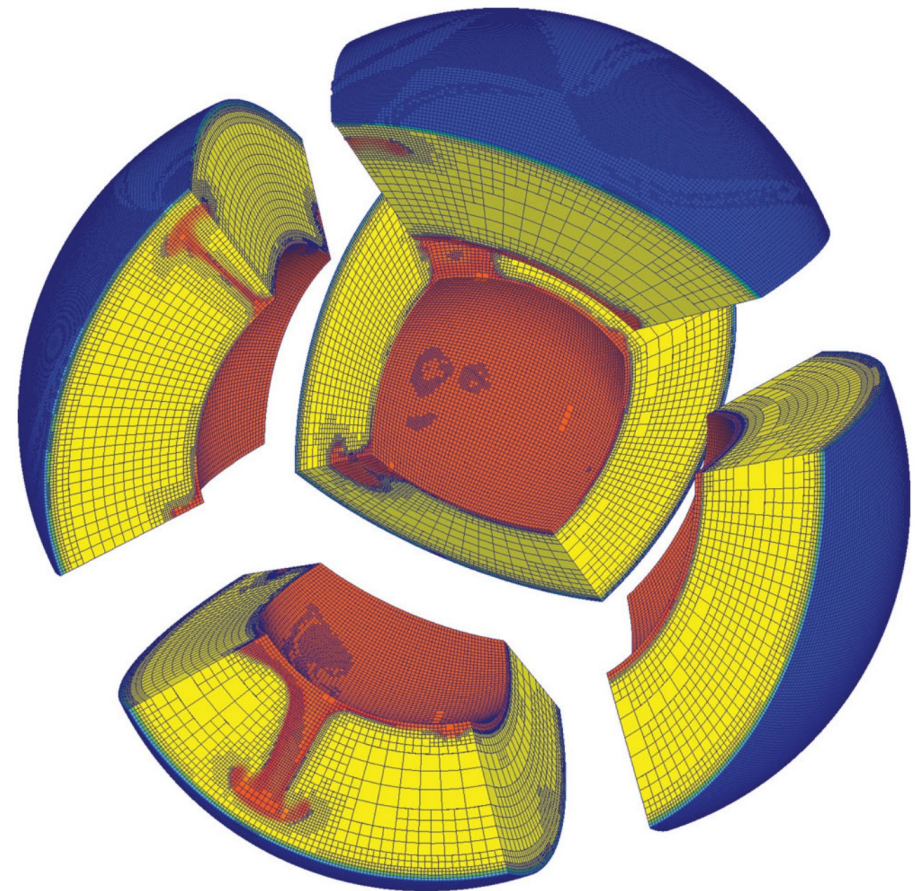




# (1) Introduction: Tomography approach

## Imaging multi-scale structure and dynamics

- Global mantle convection embraces multiple, coupled scales
- Convection models optimally account for them (*see, e.g. Burstedde et al. 2013*)
- Tomography models also should reflect multiple scales



# (1) Introduction: Tomography approach

## Geodynamically relevant parameters

Anisotropy, e.g.  $\xi = \left( \frac{v_{SH}}{v_{SV}} \right)^2$

- Provides a more direct link to mantle kinematics
- Common to assume azimuthal anisotropy or radial anisotropy

Composition, e.g.  $R = \frac{d \ln v_P}{d \ln v_S}$

- Provides a more direct link to mantle composition
- Can be inferred through ratios of shear and compressional wave-speeds

# (1) Introduction: Tomography approach

## Summary of the inverse problem

### 1. Symbolic representation

$$\begin{pmatrix} \mathbf{C}_d^{-1/2} \mathbf{G} \\ \beta \mathbf{D} \\ \alpha \mathbf{I} \end{pmatrix} \cdot (\delta \mathbf{m}) = \begin{pmatrix} \mathbf{C}_d^{-1/2} \delta \mathbf{d} \\ 0 \\ 0 \end{pmatrix}$$

### 2. Regularization

$$\mathbf{D} = \begin{pmatrix} \dots & 0 & 0 & 0 \\ -1 & 1 & 0 & 0 \\ 0 & -1 & 1 & 0 \\ 0 & 0 & -1 & 1 \\ 0 & 0 & 0 & \dots \end{pmatrix}$$

### 3. Normal equations

$$\delta \mathbf{m} = [\mathbf{G}^T \mathbf{C}_d^{-1} \mathbf{G} + \mathbf{C}_m^{-1}] \mathbf{G}^T \mathbf{C}_d^{-1} \delta \mathbf{d}$$

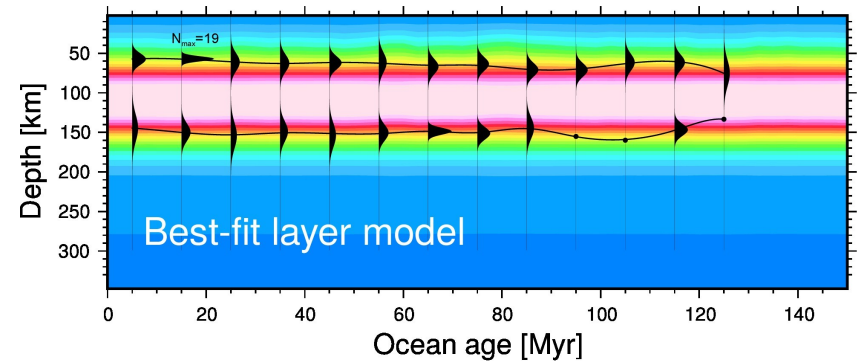
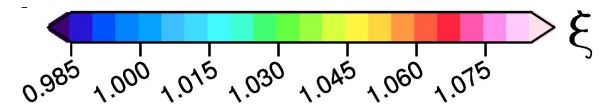
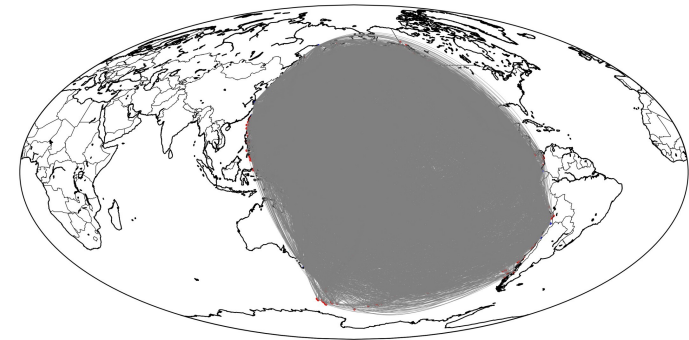
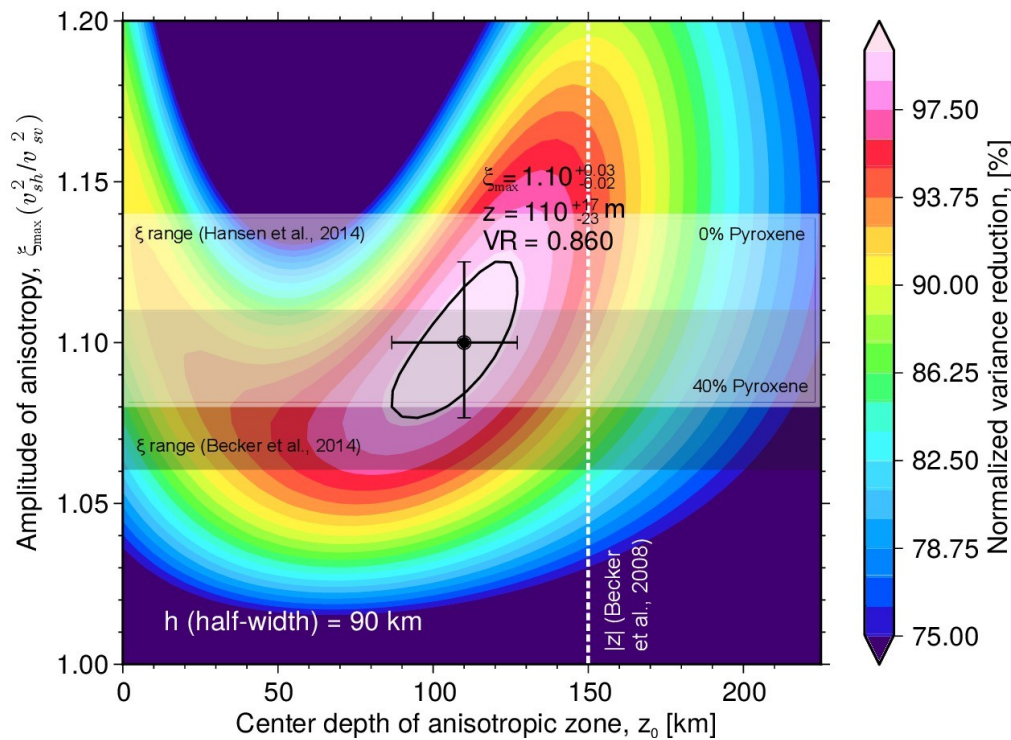
### 4. Solver

Parallel solver PETSCINV  
available for download @  
<https://github.com/auerl/petscinv>

# (3) Hypothesis test: flat layer

Grid search over layer depth  $z_0$   
and the anisotropy strength  $\xi_{\max}$

Misfit (VR) for Pacific dataset:  
$$VR = 1 - \frac{|\mathbf{G} \cdot \mathbf{m} - \mathbf{d}_{\text{obs}}|}{|\mathbf{d}_{\text{obs}}|}$$

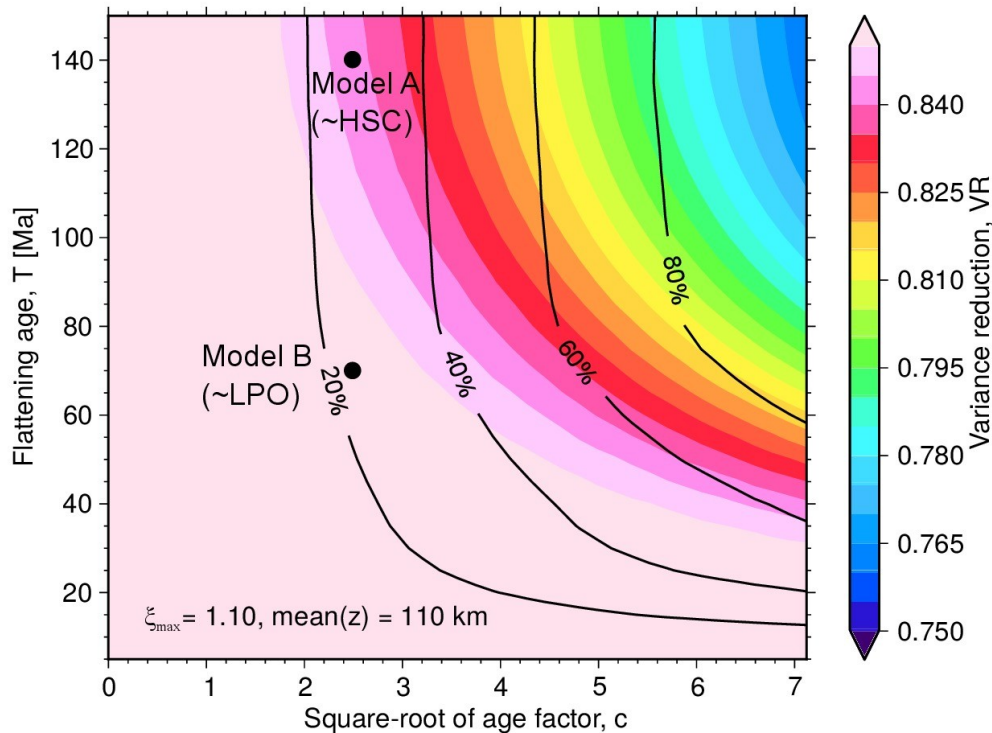


**Again:** LPO anisotropy too deep

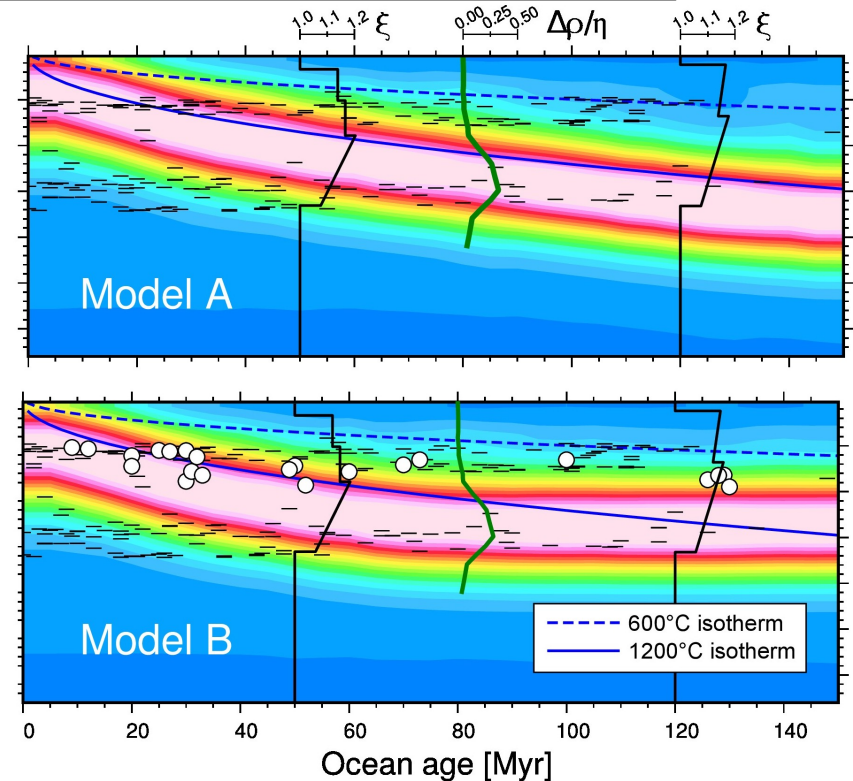
# (3) Hypothesis test: $\sqrt{\tau}$ case

**Model:**

$$z_1 = a + c\sqrt{\kappa\hat{\tau}(\mathbf{x}, f)} \quad \hat{\tau}(\mathbf{x}, f) = \begin{cases} \tau(\mathbf{x}) & \text{for } \tau(\mathbf{x}) < f \\ f & \text{for } \tau(\mathbf{x}) \geq f \end{cases}$$



Grid search over age-slope,  
 $c$ , and flattening age,  $T$



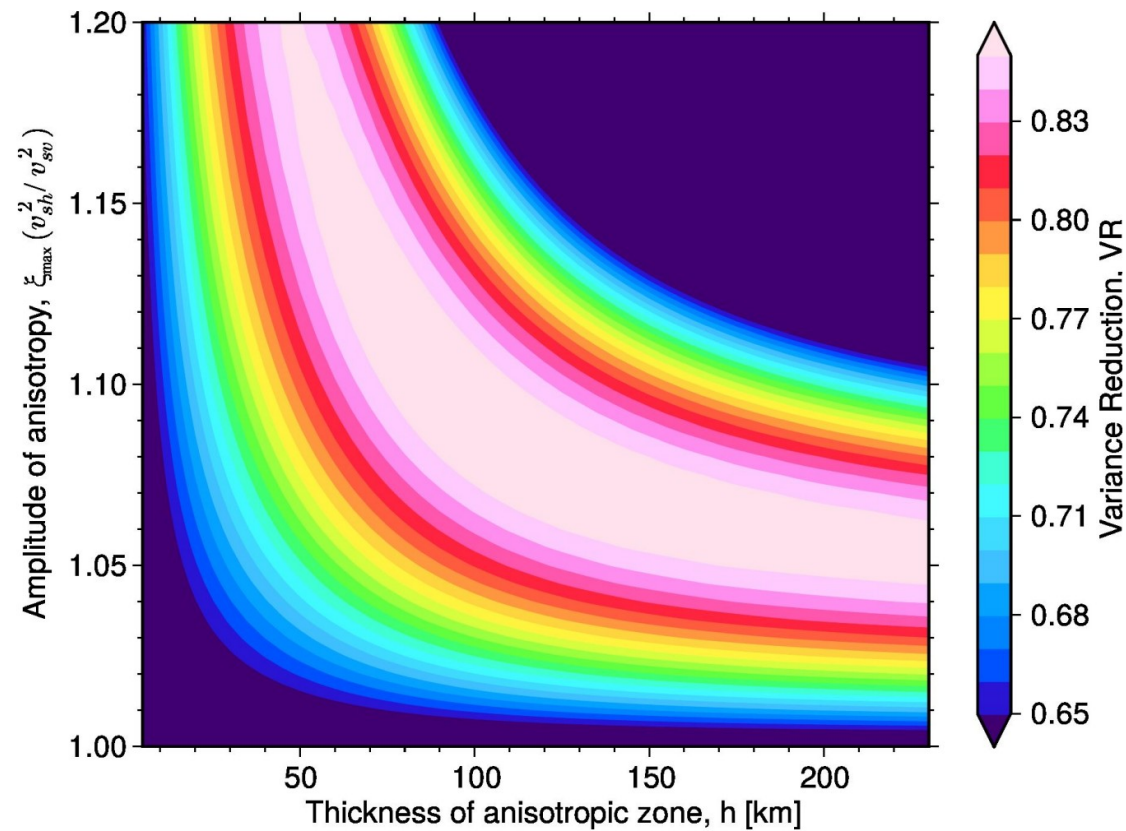
Geodynamic LPO model  
borderline compatible with data!



### (3) Hypothesis test: $\sqrt{\tau}$ case

**Model:**

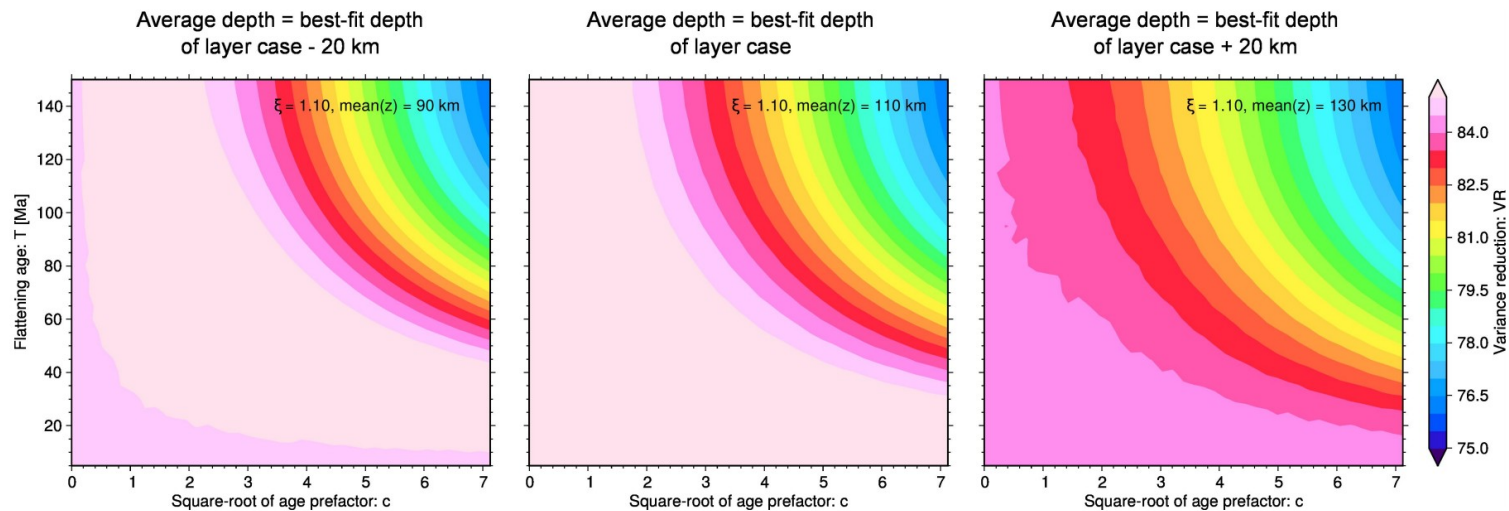
$$z_1 = a + c\sqrt{\kappa\hat{\tau}(\mathbf{x}, f)} \quad \hat{\tau}(\mathbf{x}, f) = \begin{cases} \tau(\mathbf{x}) & \text{for } \tau(\mathbf{x}) < f \\ f & \text{for } \tau(\mathbf{x}) \geq f \end{cases}$$



# (3) Hypothesis test: $\sqrt{\tau}$ case

**Model:**

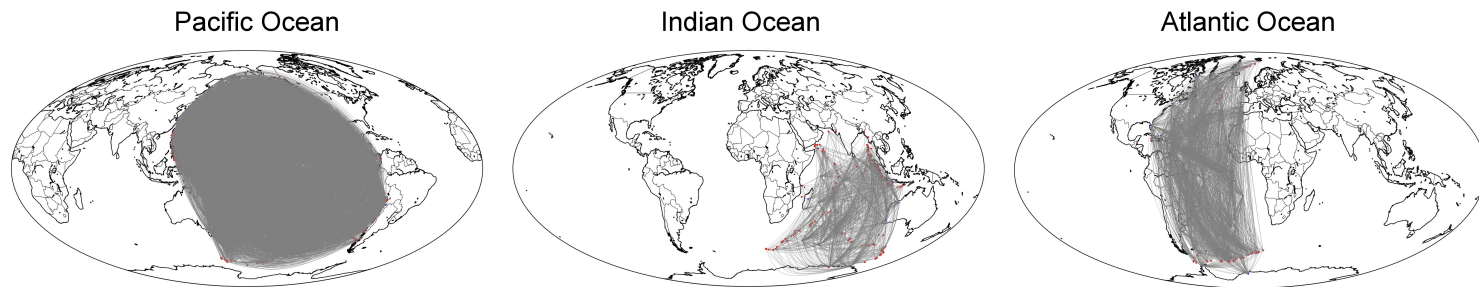
$$z_1 = a + c\sqrt{\kappa\hat{\tau}(\mathbf{x}, f)} \quad \hat{\tau}(\mathbf{x}, f) = \begin{cases} \tau(\mathbf{x}) & \text{for } \tau(\mathbf{x}) < f \\ f & \text{for } \tau(\mathbf{x}) \geq f \end{cases}$$



**Figure 3.13.:** Additional age-dependent hypothesis test (Pacific region), where the parameter  $z_0$ , controlling the average depth of the zone, was chosen to be 20 km shallower (leftmost panel) or 20 km deeper (rightmost panel) than the best-fit depth of the layer case (middle panel). The latter case shows the highest variance reductions and the general shape of the misfit contours remains the same.

# (3) Hypothesis test: flat layer

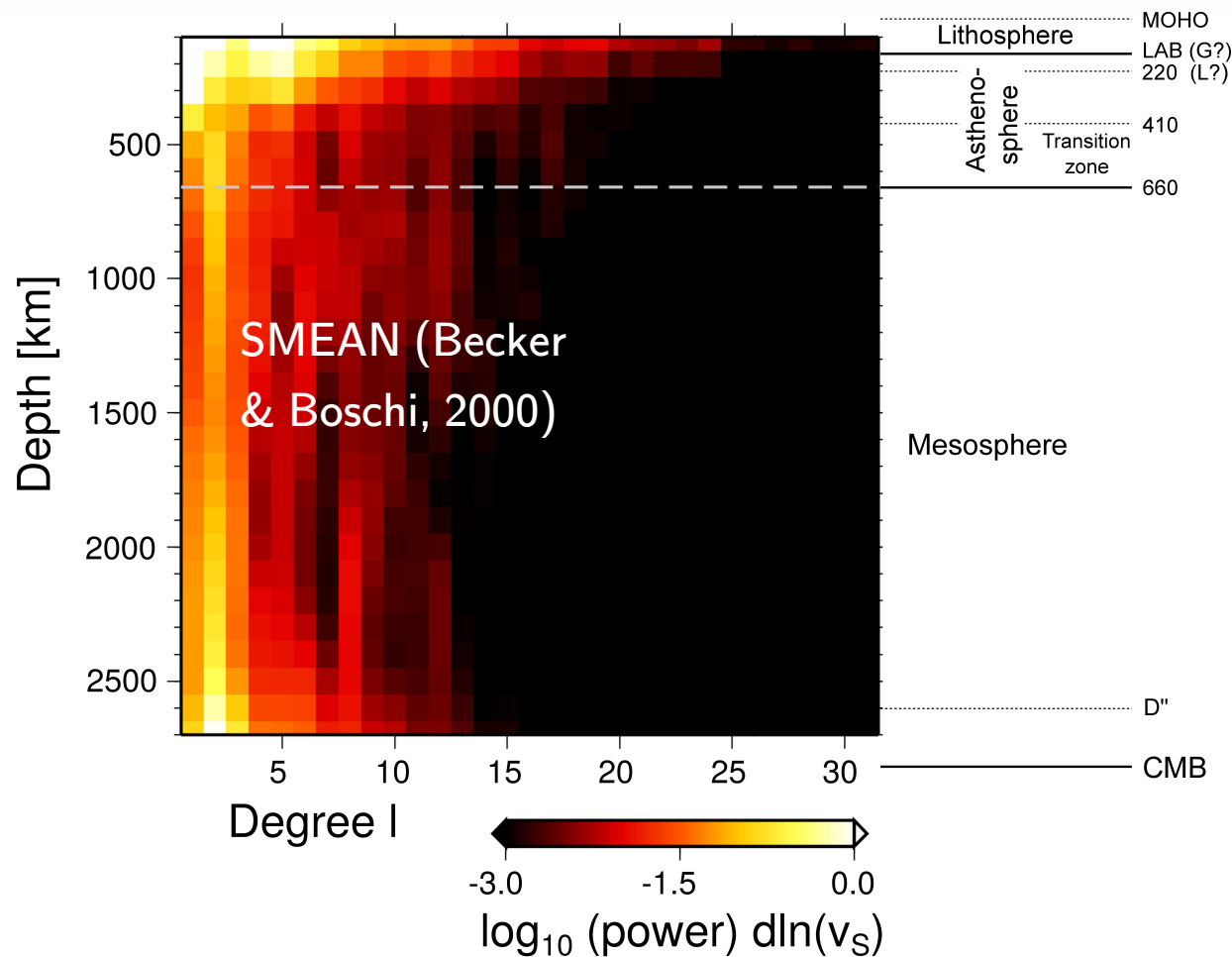
## Application to different oceanic regions



Model case	Pacific	Indian	Atlantic	Global
depth $z_0$ [km]	110	85	80	105
radial anisotropy, $\xi_{\max}$	1.1	1.09	1.07	1.09
variance reduction, VR	0.860	0.733	0.781	0.854

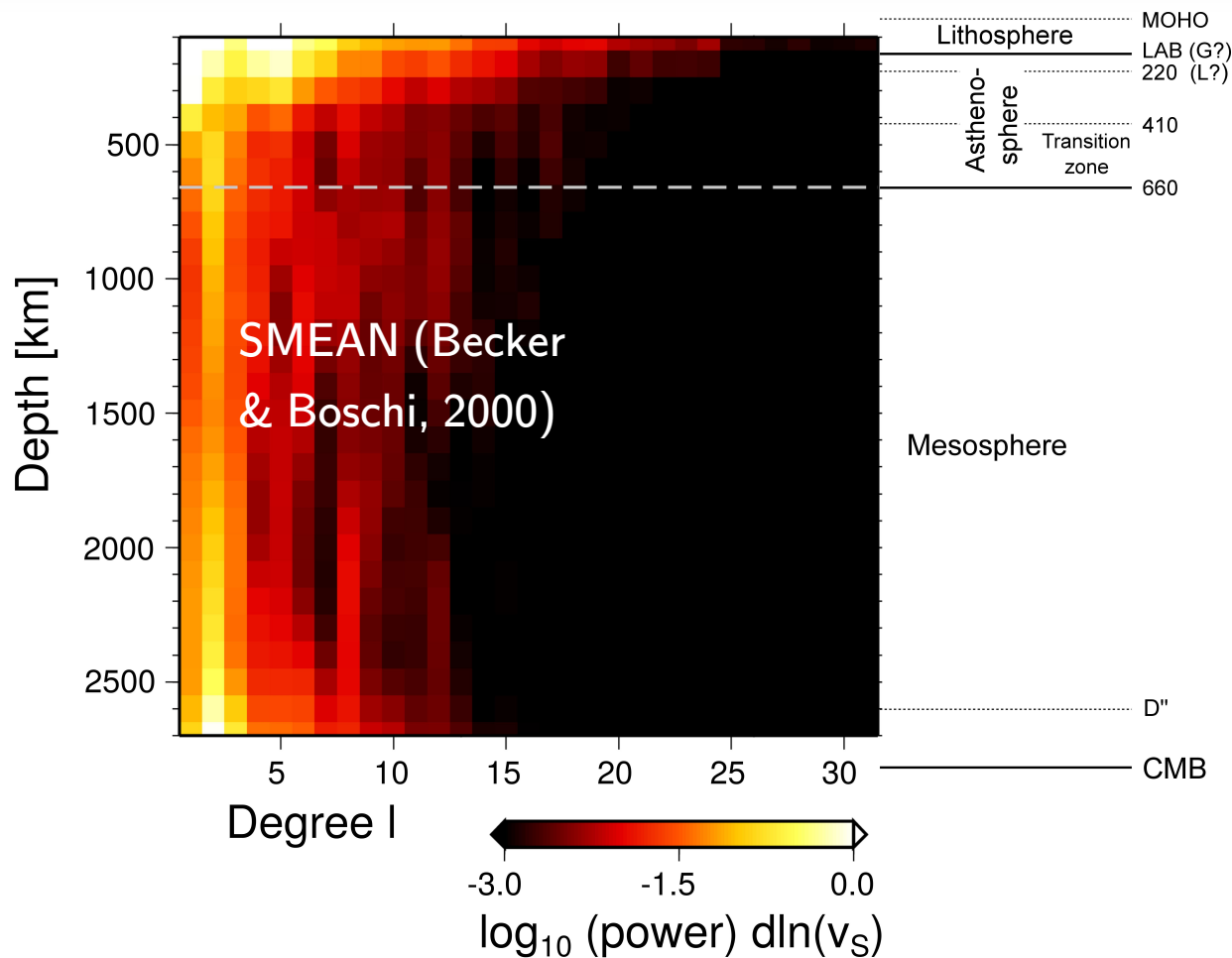
# (1) Introduction: Earth's 3D structure

Visualization in the spherical-harmonic domain



# (1) Introduction: Earth's 3D structure

## Visualization in the spherical-harmonic domain

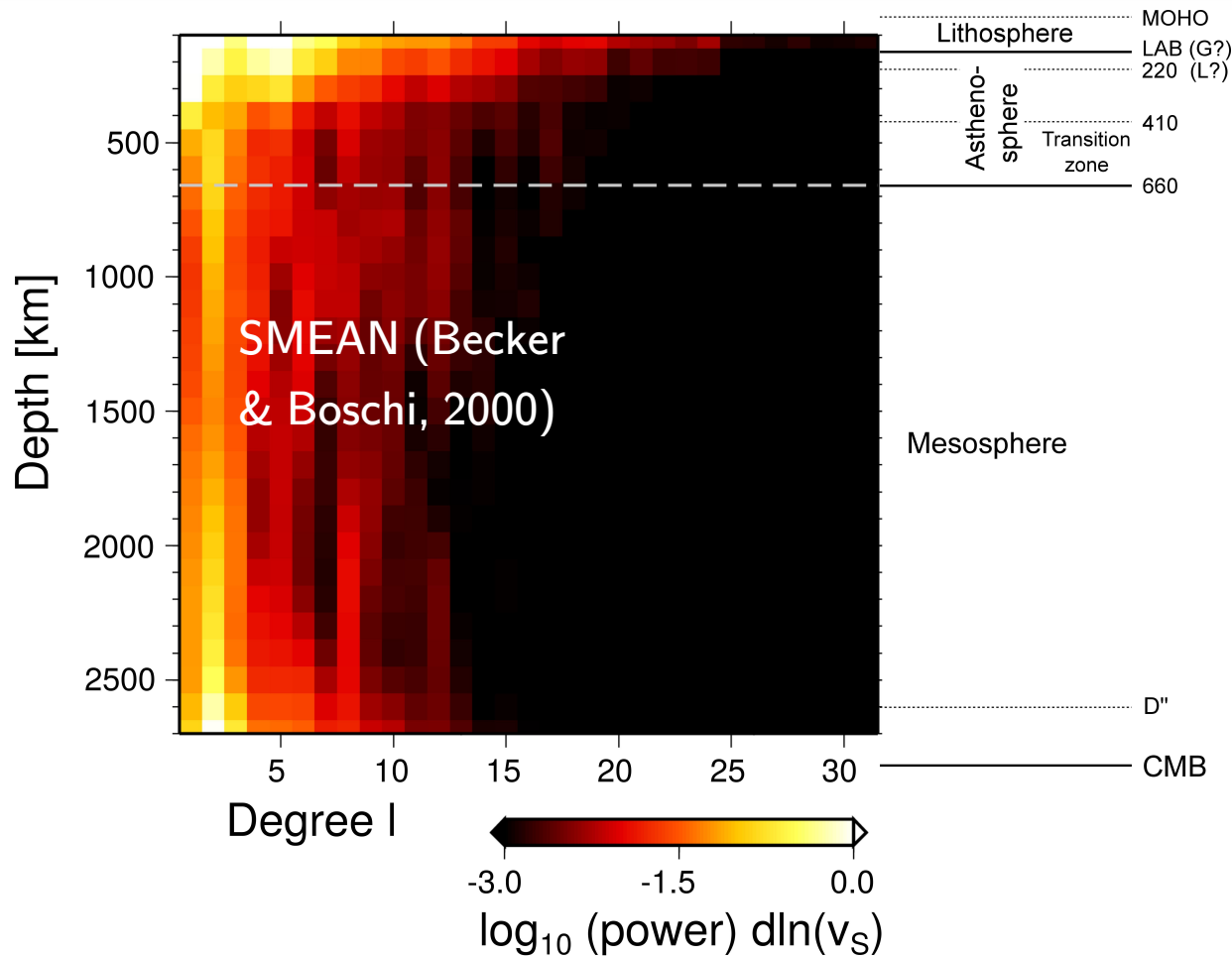


- Strong degree 5 in the upper mantle



# (1) Introduction: Earth's 3D structure

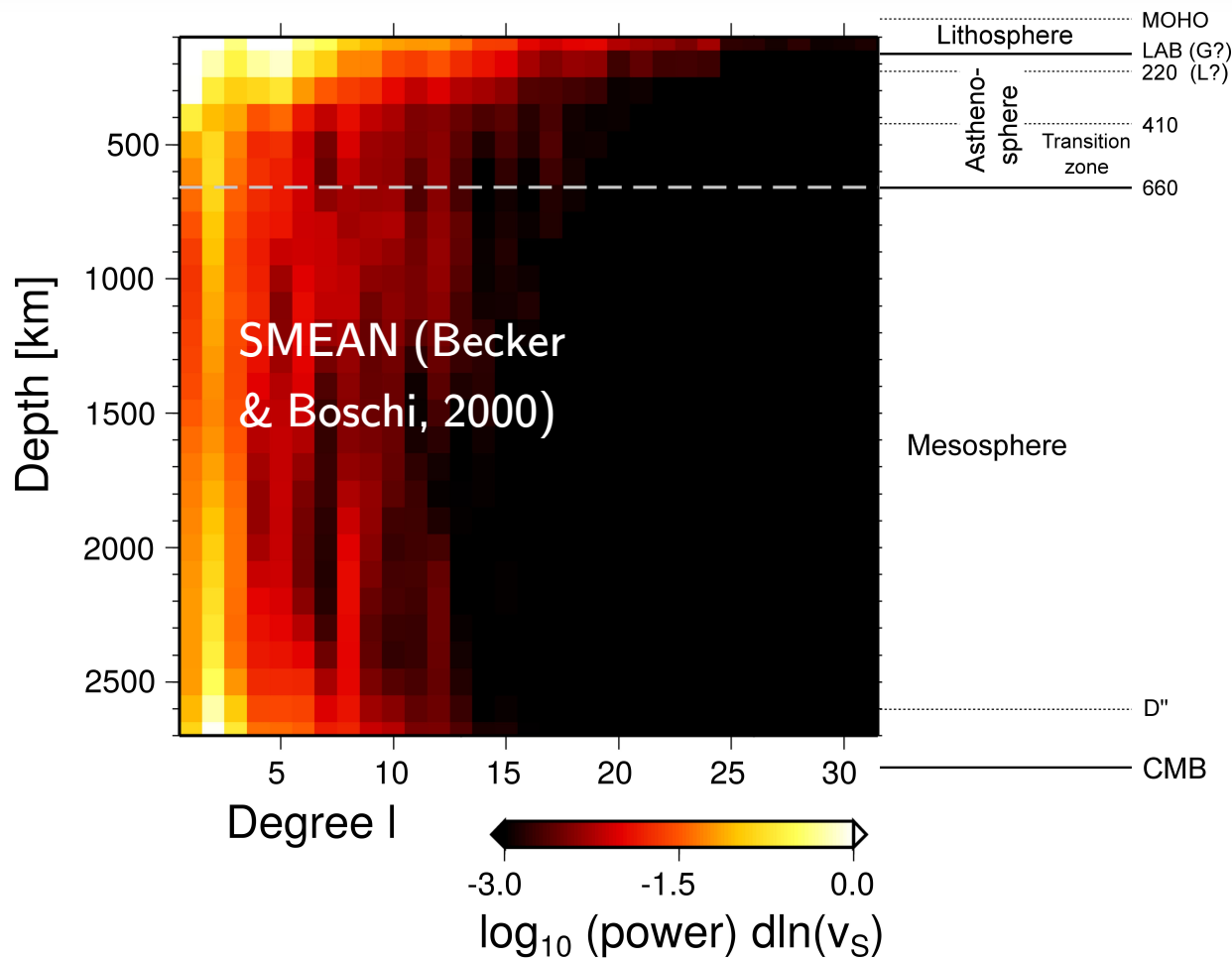
## Visualization in the spherical-harmonic domain



- Strong degree 5 in the upper mantle
- Strong degree 2 in the lower mantle

# (1) Introduction: Earth's 3D structure

## Visualization in the spherical-harmonic domain



- Strong degree 5 in the upper mantle
- Strong degree 2 in the lower mantle
- Large model power near thermal boundary layers

